



SSD VALIDATION FACILITIES AND PROCEDURES IN LECCE

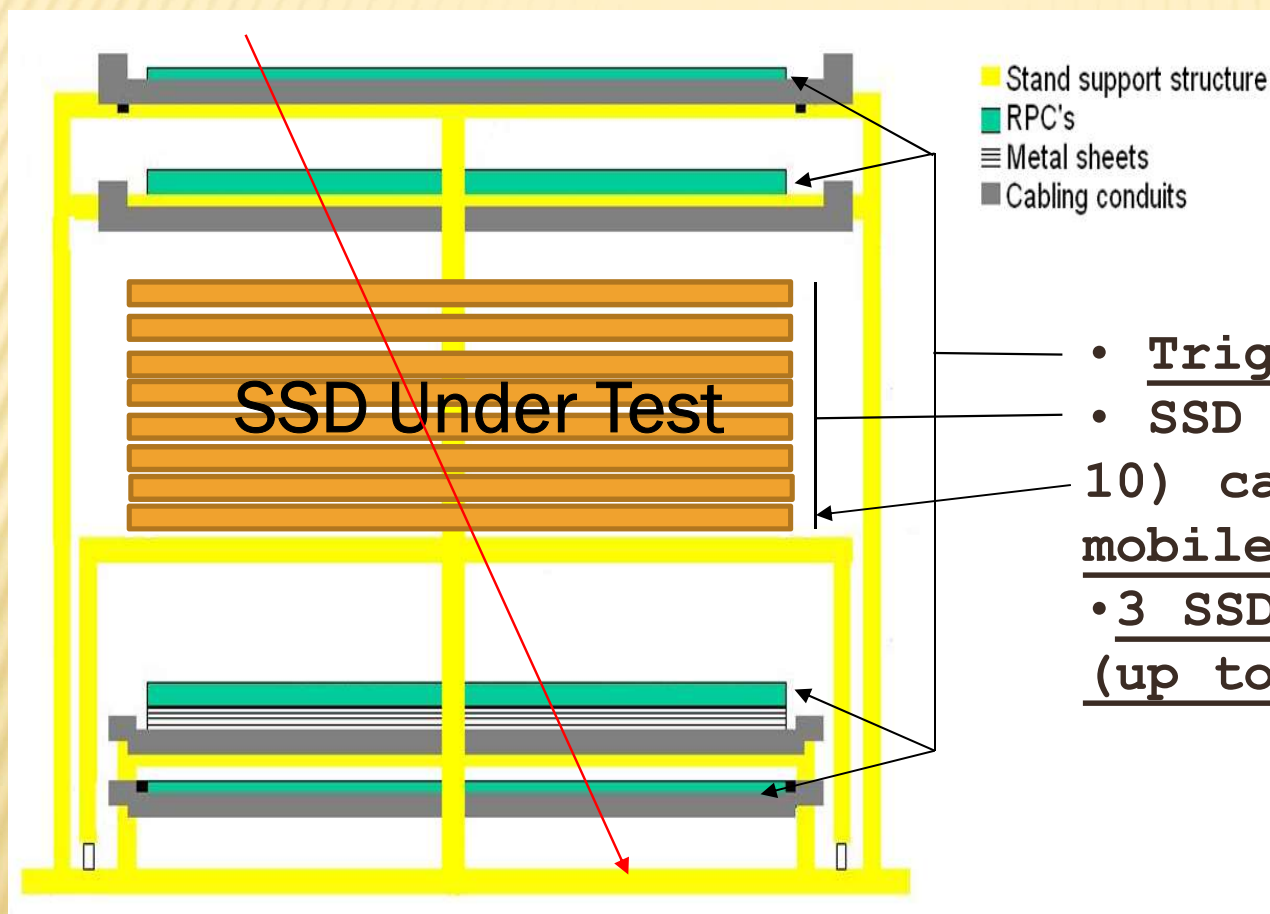
Gabriella Cataldi
on behalf of
Lecce Auger Group



STARTING POINT: ATLAS RPC TEST FACILITY

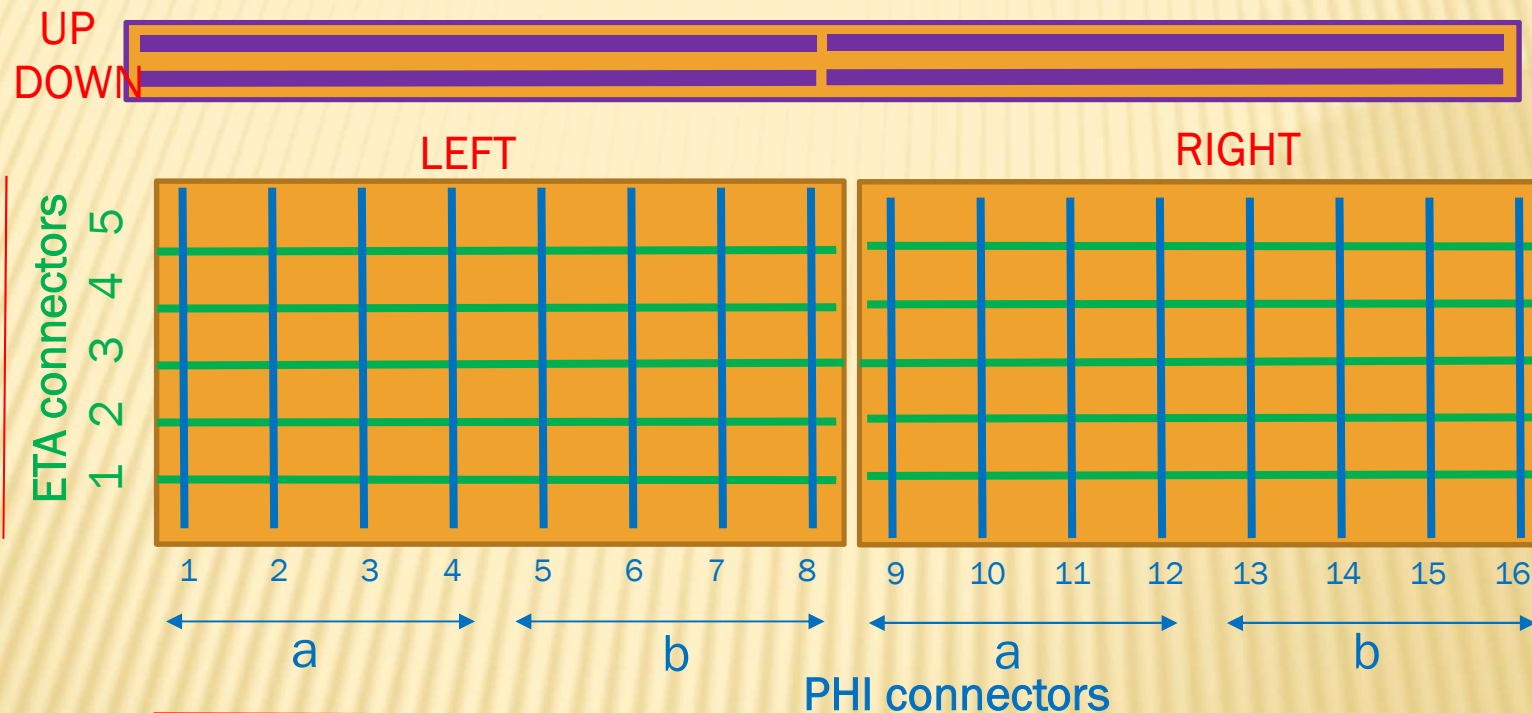
- Test facility in Lecce for testing all the RPC modules assembled in Lecce during ATLAS construction.
 - modules tested
- 2005 last run performed
 - RPC test stand was dismissed
 - 12 years of inactivity
- We started the reworking in autumn 2016
 - Self-confidence: *We will have all the RPC working again as well as DAQ, and everything...*
- Some gas leakage problems - ordinary like in the ATLAS experiment
 - Some parts more dismantled then inactive.
 - Electronic missing or broken.
 - DAQ too old.
 - Gas System in pieces.

RPC TEST STAND TRIGGER



- Trigger based on 4 RPC
- SSD under test (up to 10) can be located on a mobile structure
- 3 SSDs connected to PMTs (up to 8 DAQ channel)

RPC-CHAMBER LAYOUT



- Gap gas = 2 mm
 - $C_2H_2F_4$ (94.7%), C_4H_{10} (5%), SF_6 (0.3%)
- HV~10 KV Saturated avalanche regime
- Time response ~1 ns
- Readout panels: ϕ and η copper strip (pitch ~3 cm).

TEST STAND OVERVIEW



The stand is built with 4 RPC units triggering and tracking cosmic rays

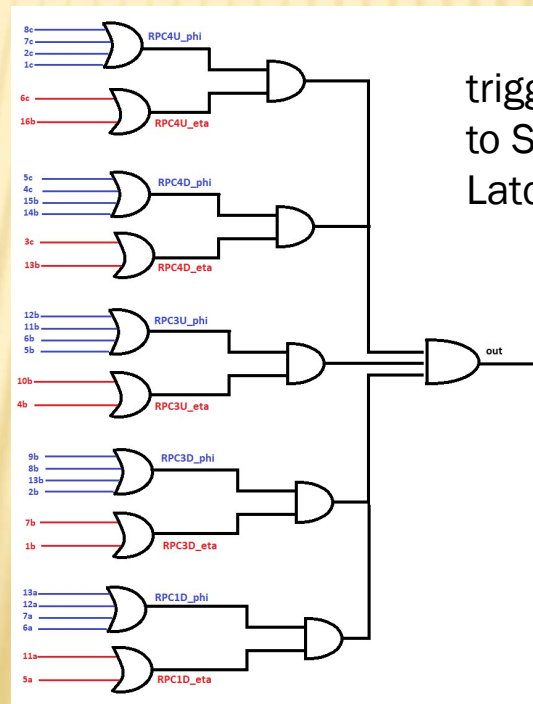
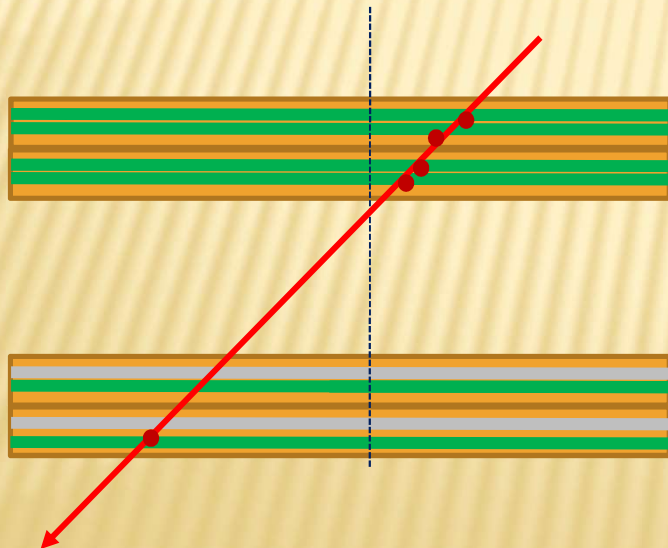
It can host up to 10 SSD.

Only 3 equipped PMTs

ACQUISITION SYSTEM



- VME Modules (VXI System)
 - Receiver (TTL signals)
 - latch
- NIM modules for the trigger logic
- CAEN SY5527 (0-15kV)
- VME Caen Digitizer V1730D (14-bit @ 500 MS/s)
- DAQ-DCS developed using Labview

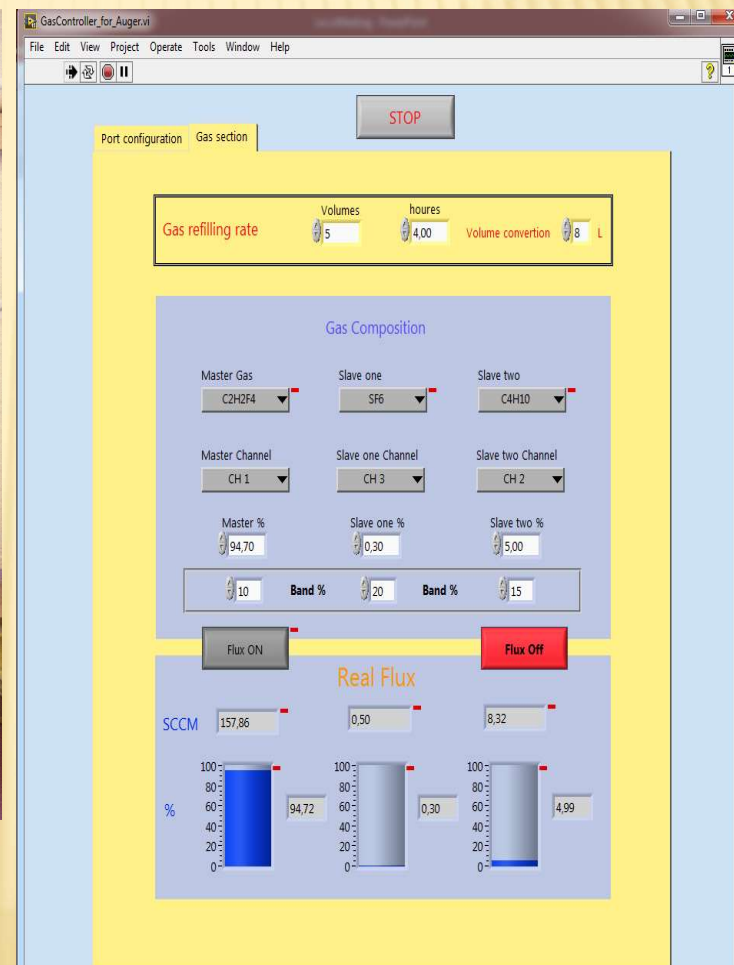


trigger-out is used
to STOP the RPCs
Latch

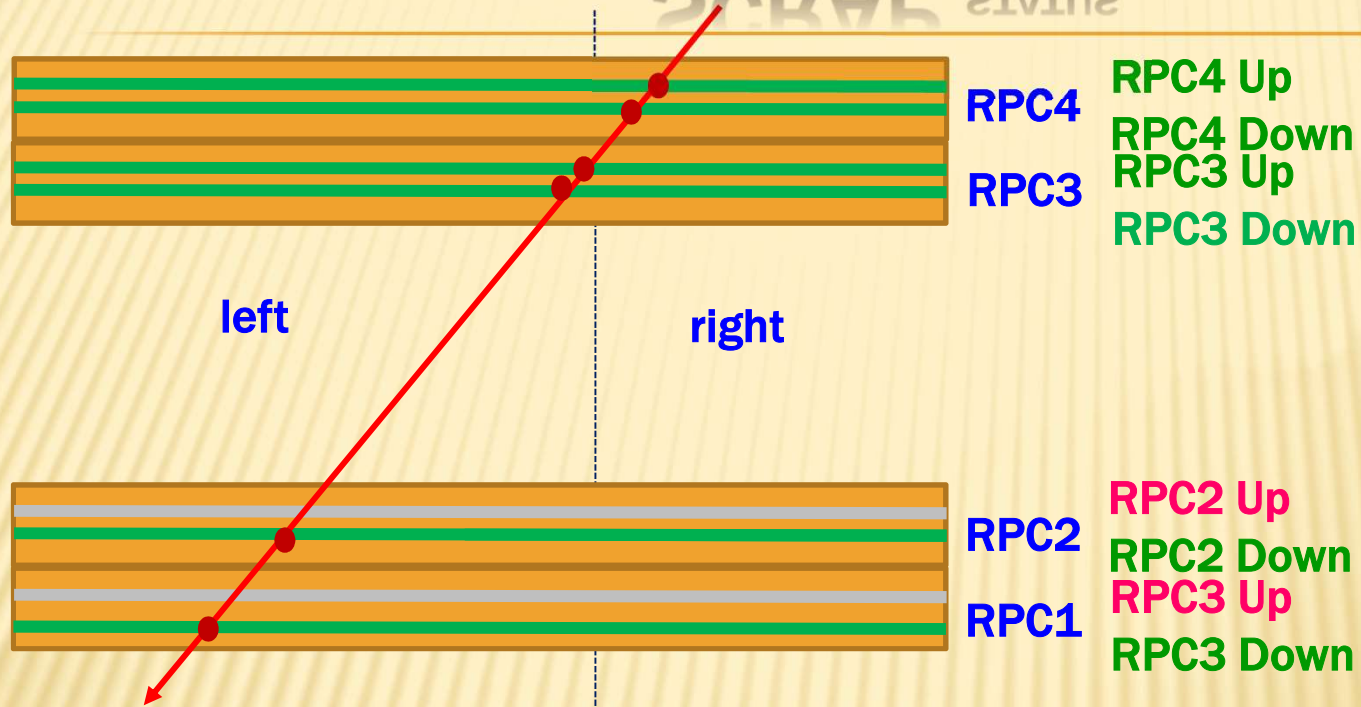
trigger-out is used to
acquire the digitizer



GAS SYSTEM



SCRAP STATUS

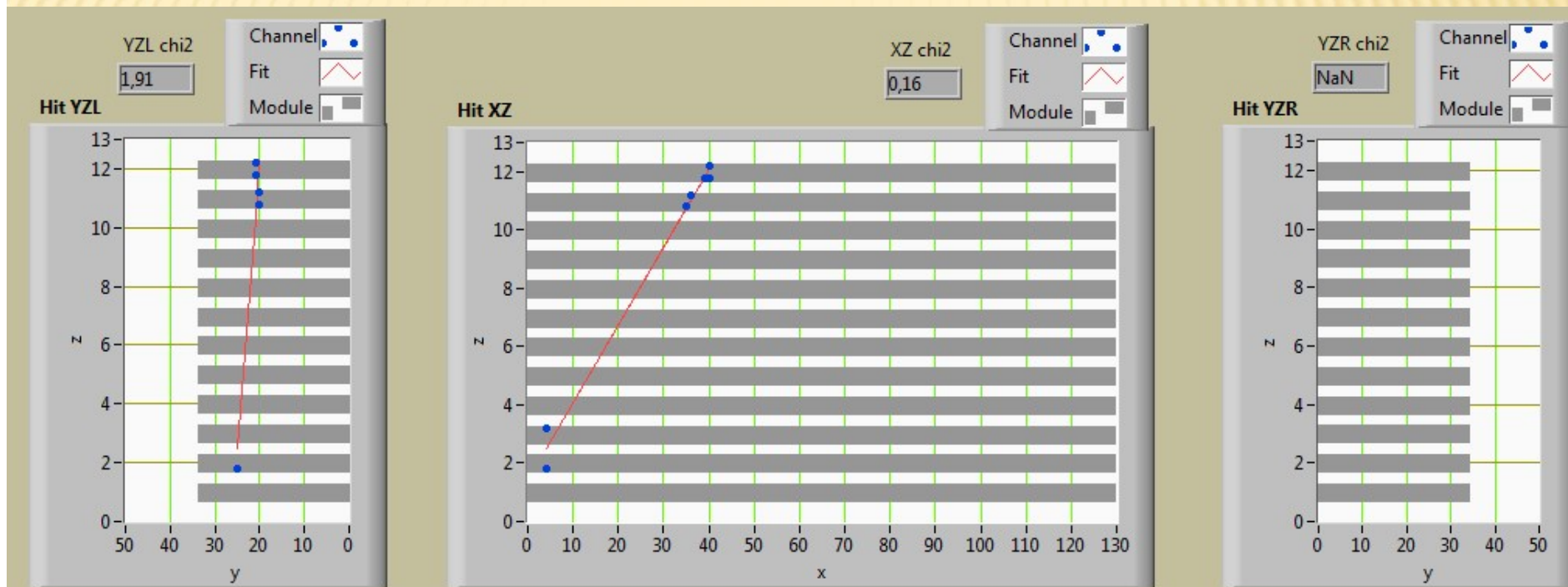


5 gas volumes
used for trigger.
6 gas volumes
used for tracks

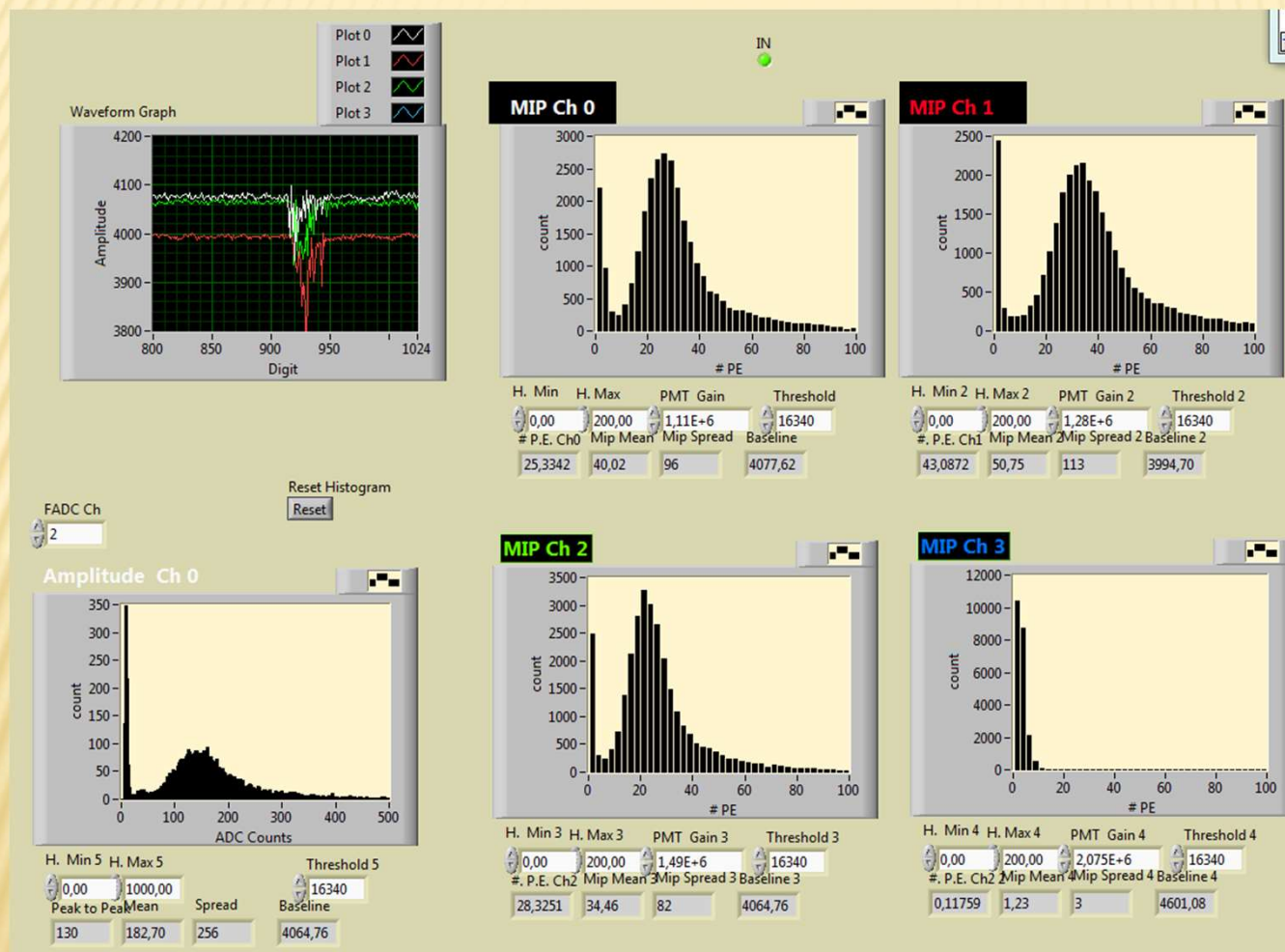
After 12 years of inactivity we have turned on the ATLAS RPC's Test Stand and renamed it **SCRAP** (Setup Cosmic Rays Auger Production)

- Labview for DAQ and Monitoring
- Reconstruction software (C++ under Unix)
- Digitizer data from the SSD module merged with the RPC data.
- Analysis and Production of ROOT tree to test MIP, light tightness and Homogeneity of the SSD detectors.

SCRAP- EVENT DISPLAY FOR RPCs



SCRAP-EVENT DISPLAY FOR SSD



From the digitizer signal to the MIP passing from

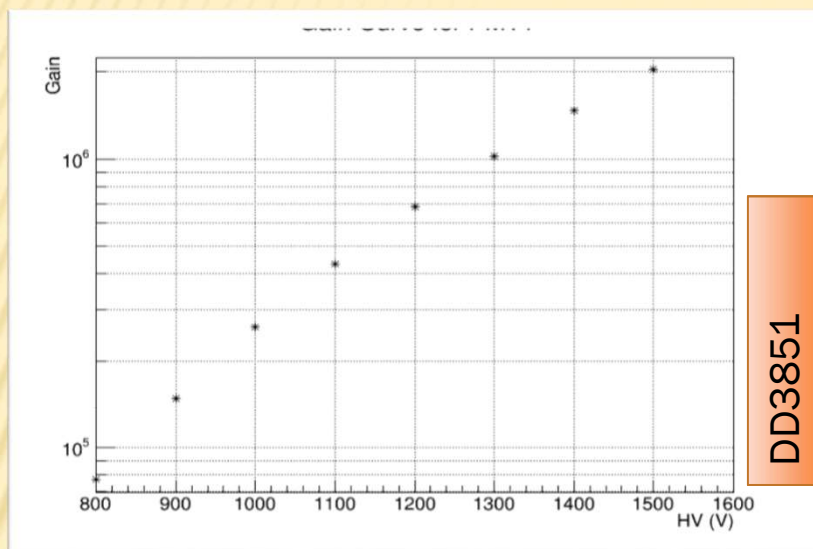
THE CALIBRATION OF PMTs

- ✗ Every PMT used for the test is calibrated in a dark box in order to establish the gain at different HV, respect to SPE.
- ✗ 3 PMTs with passive base and selected range DD3851, DD4151, DD4251.
- ✗ All PMTs have mu-metal shield.

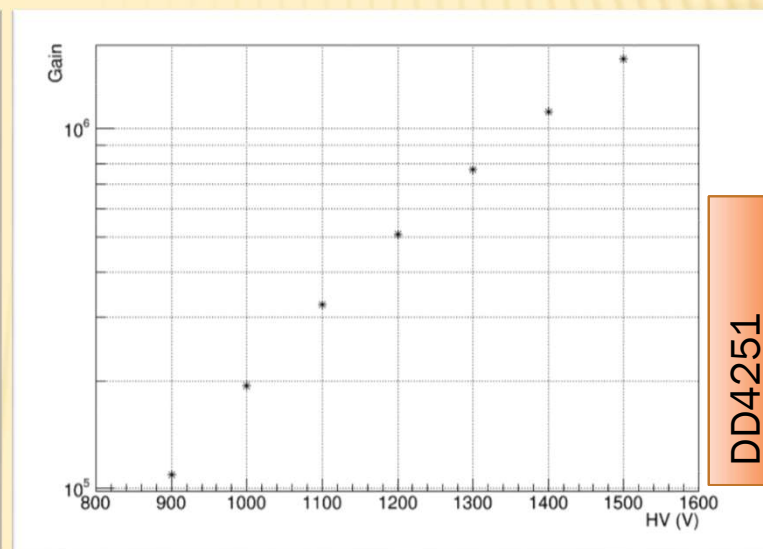
CALIBRATION METHOD

- + A blue LED (470 nm, 45 deg viewing angle) pulsed light source is used.
- ✗ SPE search (usually @1400V or @1500V)
 - + The pulsed light source is tuned (via a transmittance filter) in order to have 90% pedestal, and 10% signal. For Poisson distribution, the absence of photoelectron 90% of the time ensures a very low contamination of events with >1 PE.
- ✗ GAIN Vs HV
 - + To get the gain as a function of Voltage, the LED intensity is fixed while varying the PMTs HV.
 - + The curve is then normalized to the SPE value.

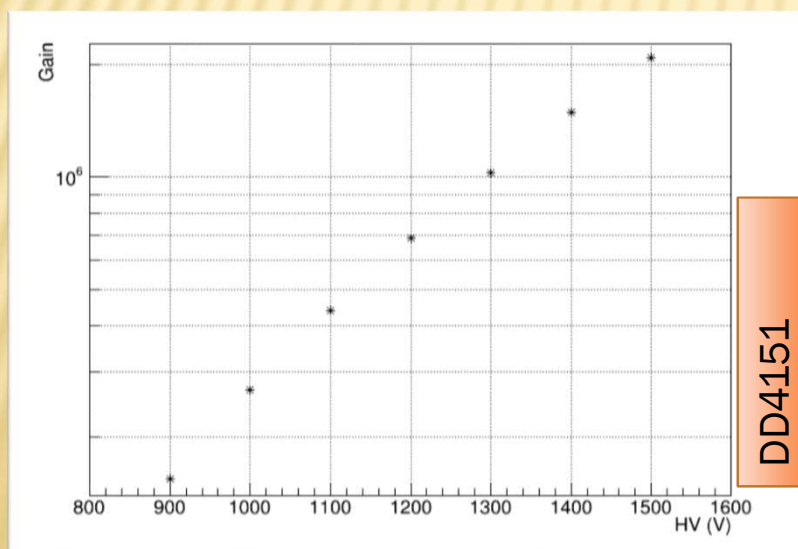
GAIN V_s HV CURVES OF PMT_s



DD3851

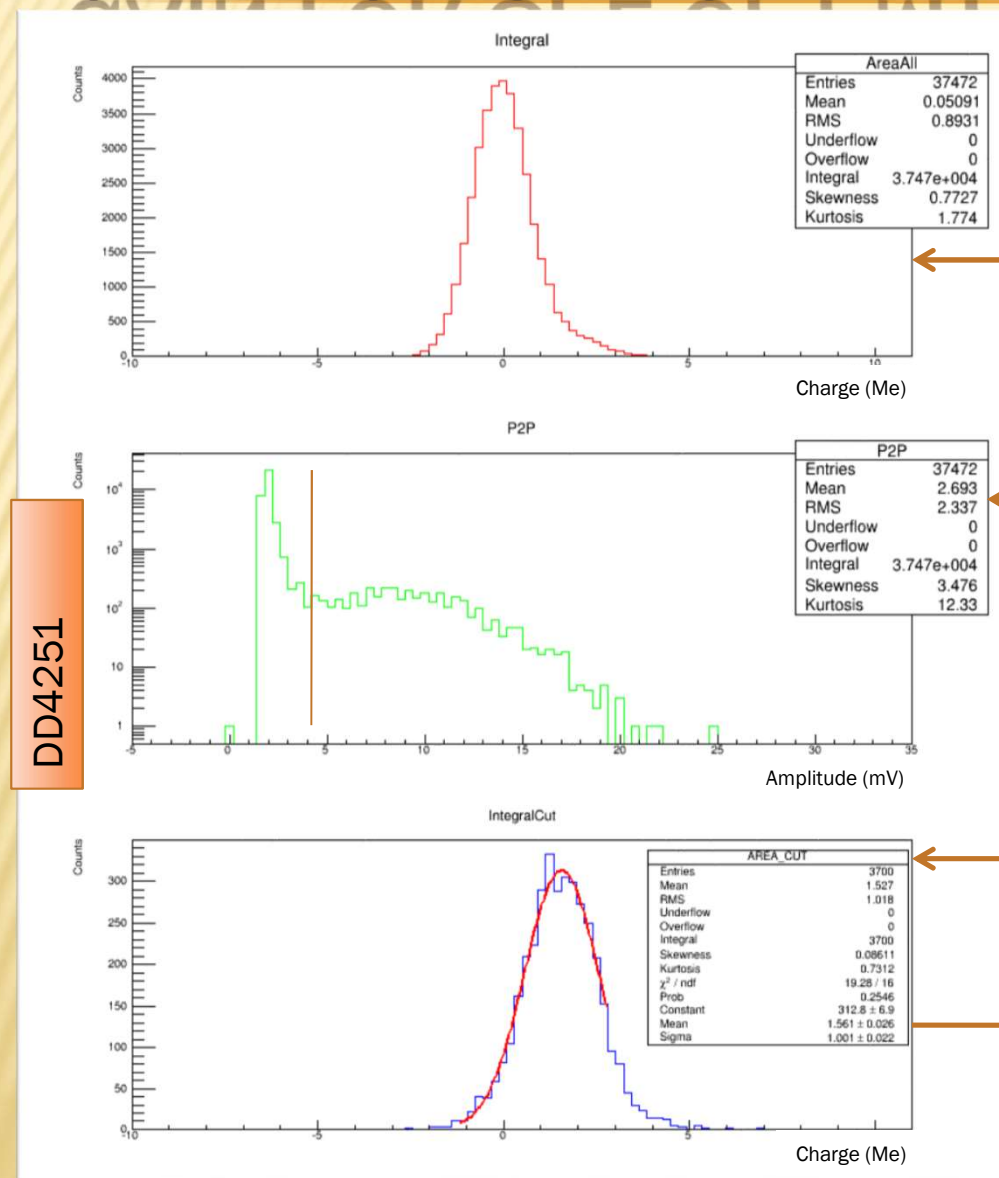


DD4251



DD4151

GAIN FOR SPE OF PMT



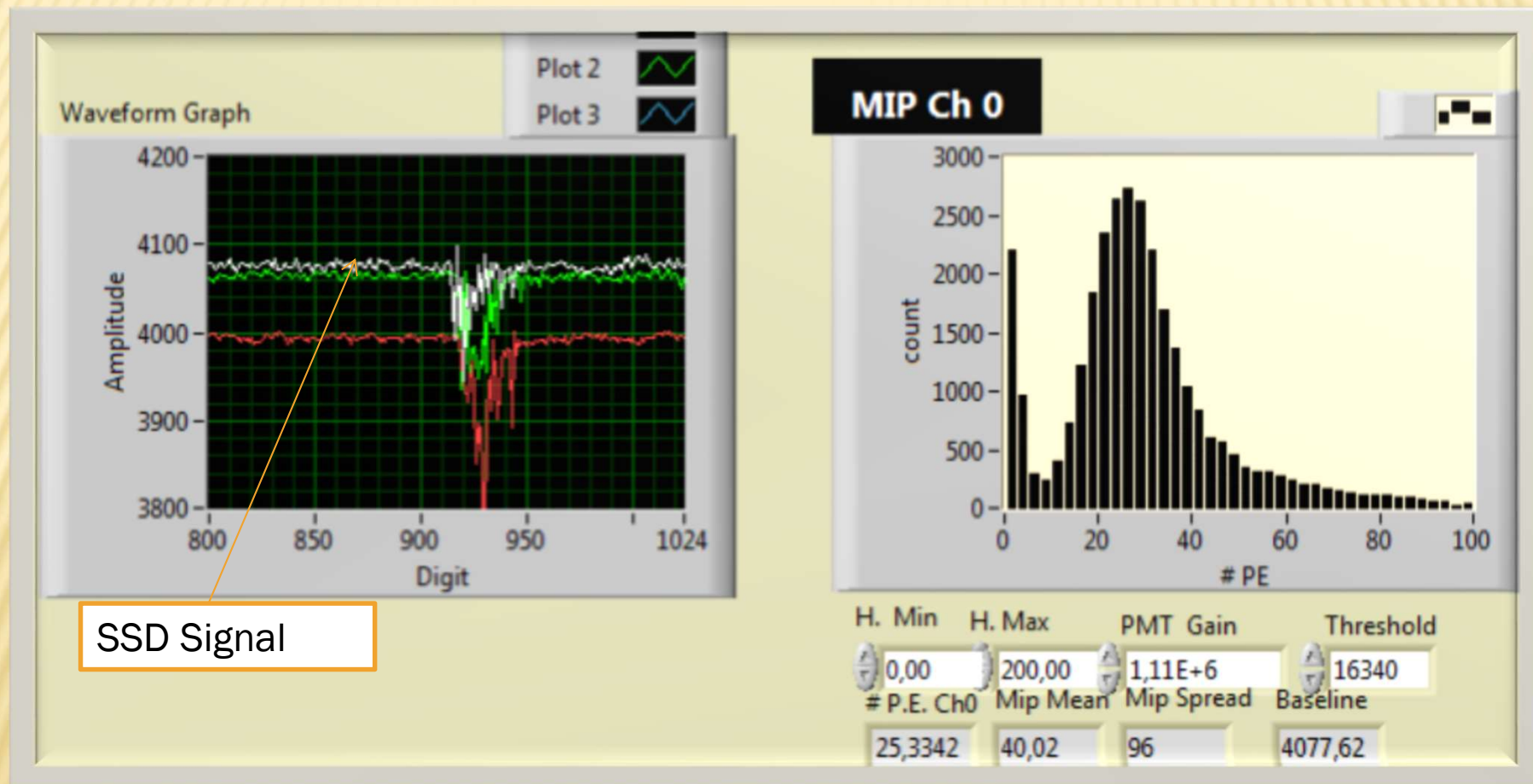
Charge integral in 90% pedestal, and 10% signal

Peak to Peak in 90% pedestal, and 10% signal

Charge integral after a cut on "Peak to Peak".

Gain of spe @1500V

NUMBER OF PHOTOELECTRONS IN A MIP-MONITORING

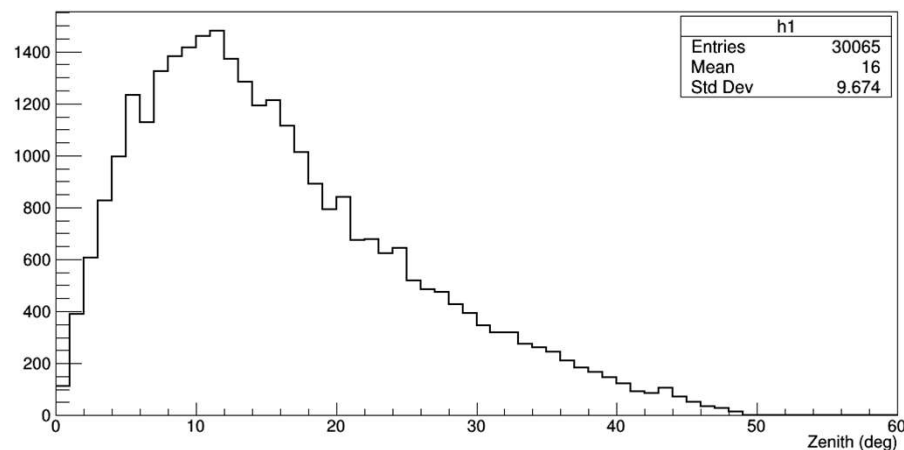


- ✓ Find the minimum
- ✓ Integrate in a window of 140 nsec subtracting the baseline

- ✓ Use the gain to determine a number of photoelectrons per MIP

NUMBER OF PHOTOELECTRONS IN A MIP- GEOMETRY

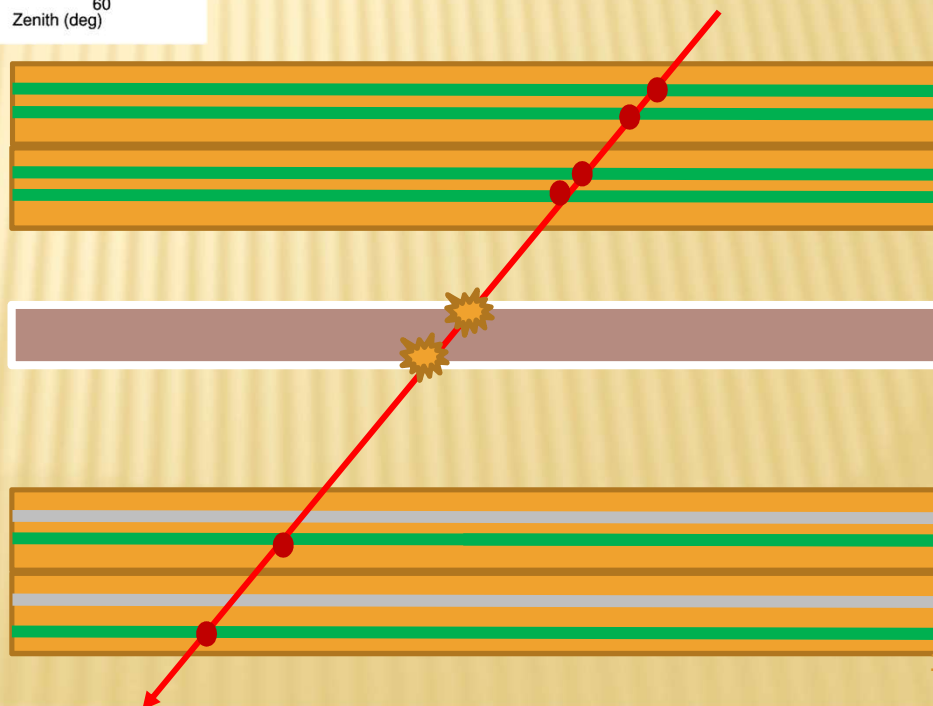
Reconstructed zenith angle.



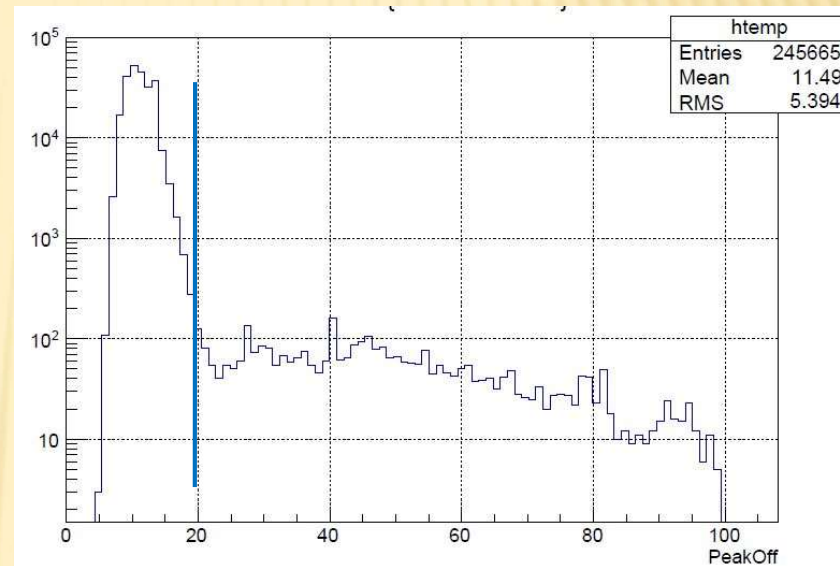
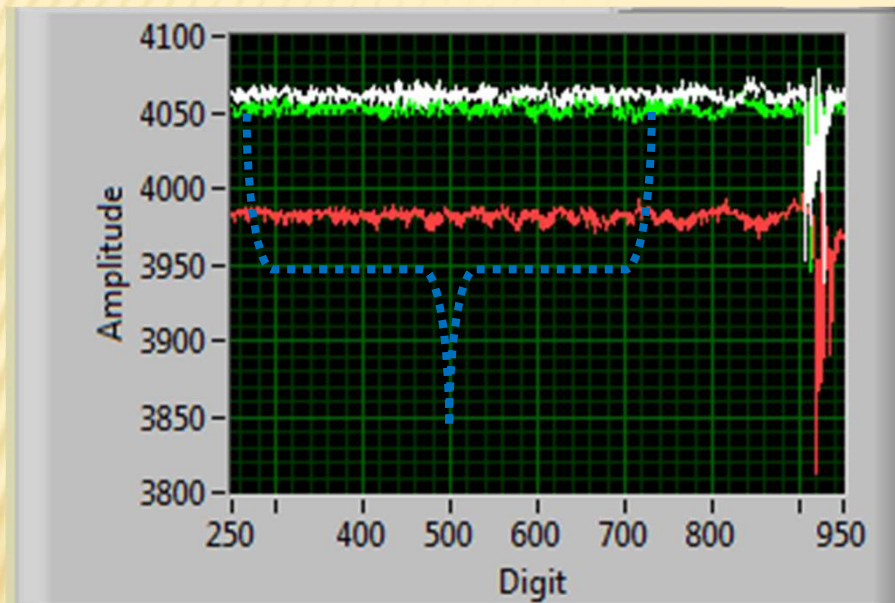
- ✓ Offline- Reconstruction of the track of the RPCs
- ✓ Calculation of the zenith angle and track length

The reconstructed zenith angle is used to correct the track length, in order to refer to a “Vertical MIP”

.. see the results in Daniele’s talk



LIGHT TIGHTNESS OF SSD EVERYDAY CHECK!



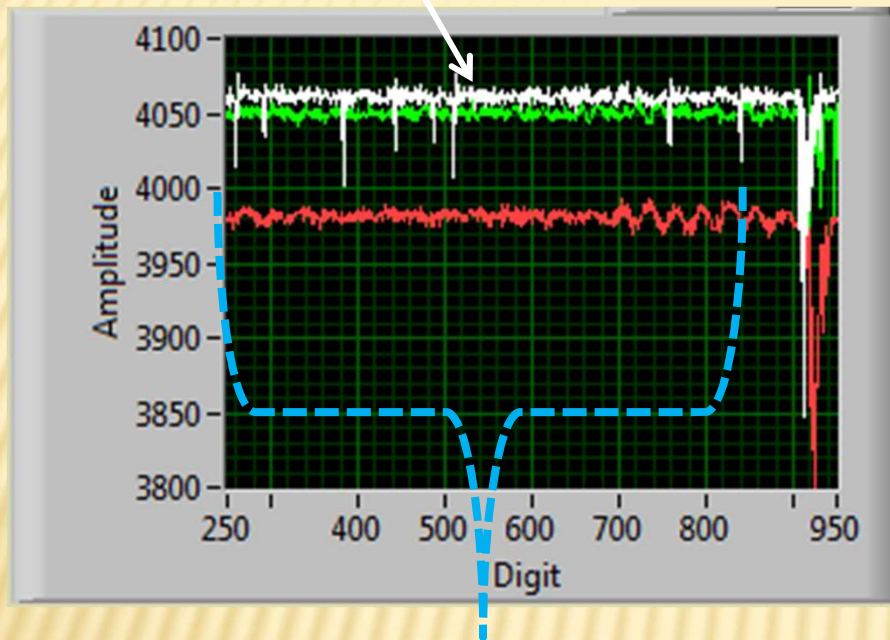
- ✓ Find the minimum “off signal” (pre-signal)
- ✓ Cut on ADC-count distribution
- ✓ Integrate in a window of 140 nsec subtracting the baseline
- ✓ Calculate the rate of “#pre-signal/baseline”

- ✓ If the ratio is higher then 2%.

DO THE LIGHT CHECK!

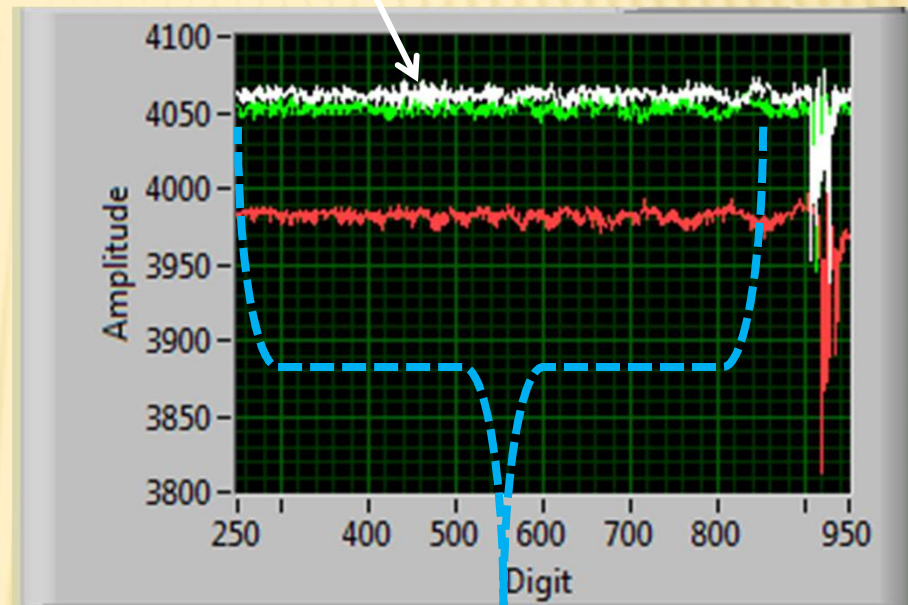
LIGHT TIGHTNESS OF SSD

✓ CIPOLLA- before light care



- ✓ Go with a flash light around the detector until you catch “THE POINT”.
- ✓ This was in correspondence of a part of “panel-frame” contour.

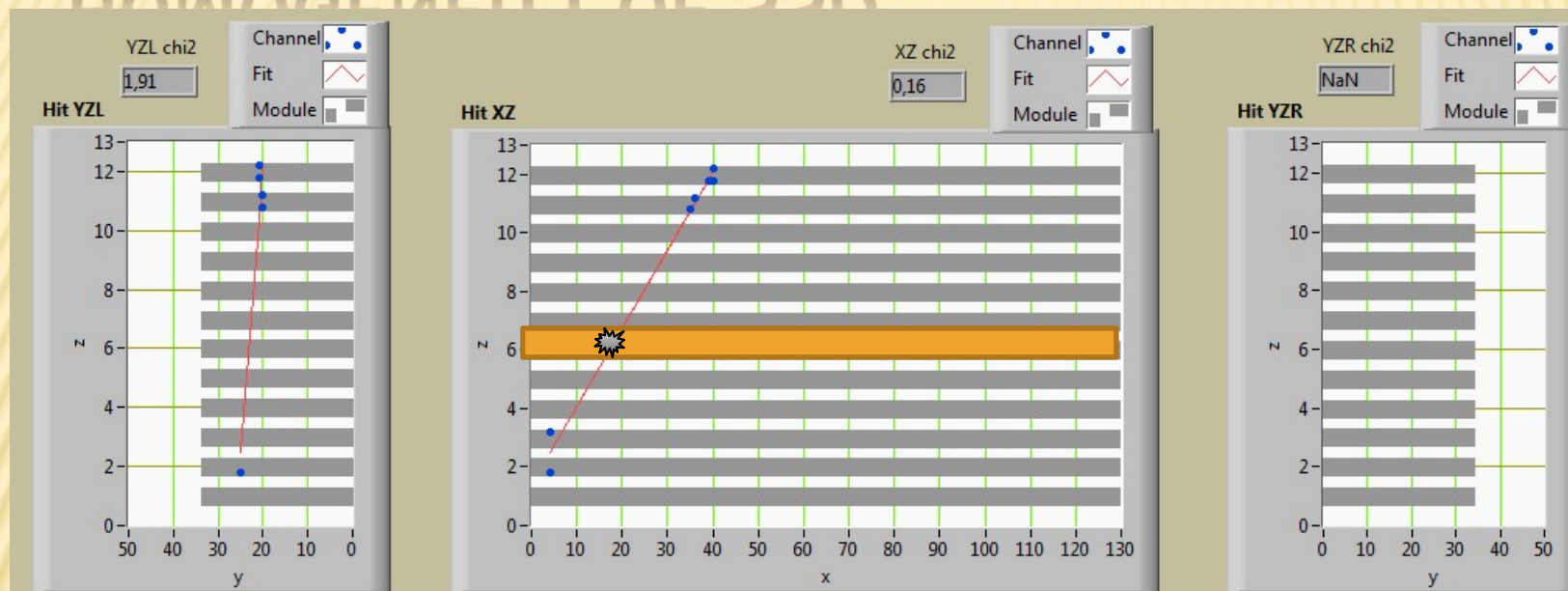
✓ CIPOLLA- after light care



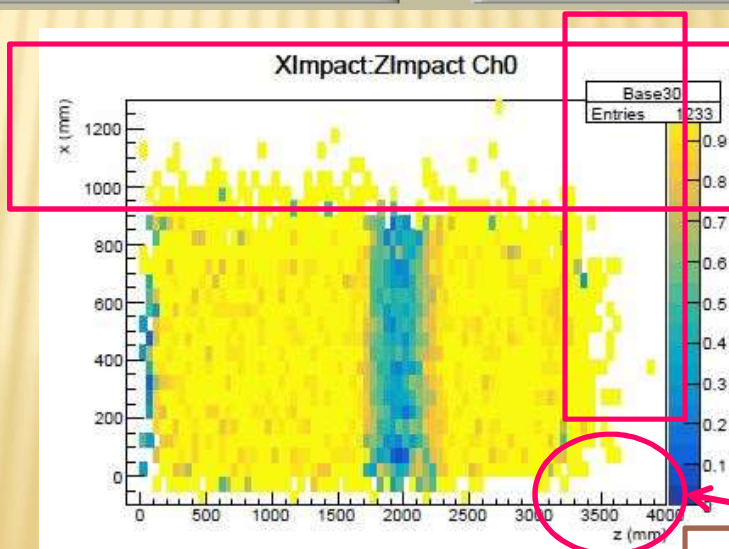
✓ “THE POINT” After sealing.

✓ Do again the “Peak-off analysis”

HOMOGENEITY OF SSD.



- ✓ Use SCRAP-RPCs to determine the impact point on SSD.
- ✓ Map ON/OFF the SSD.



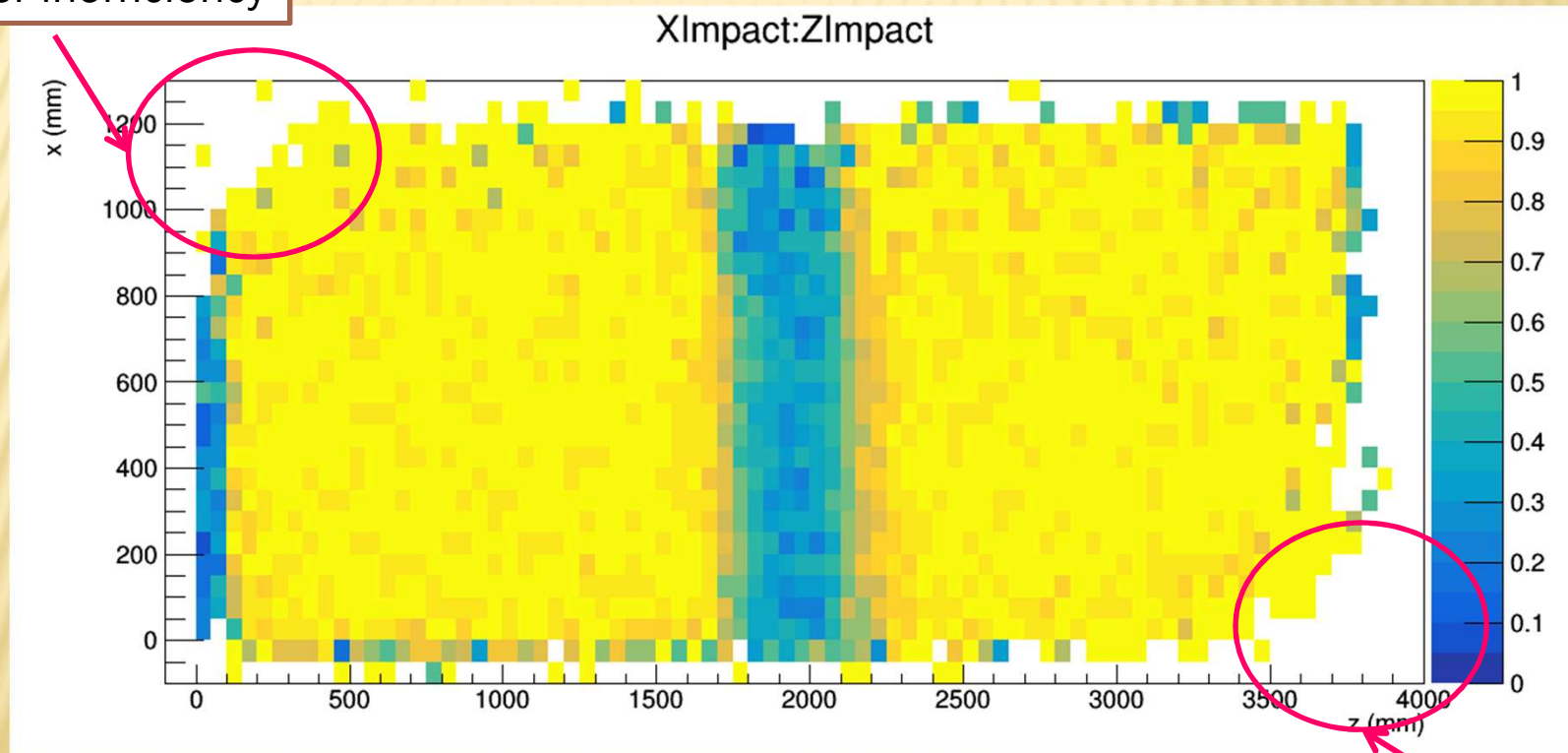
RPC tower too small

Trigger inefficiency

HOMOGENEITY OF CHAMBER.

- ✓ Rotate the chamber to have a better map (THE RPCs are smaller then SSDs)
- ✓ Merge the detectors

Trigger inefficiency



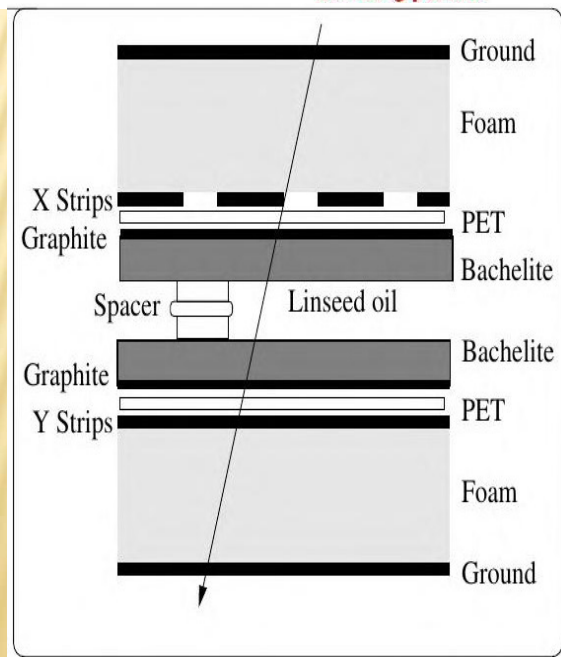
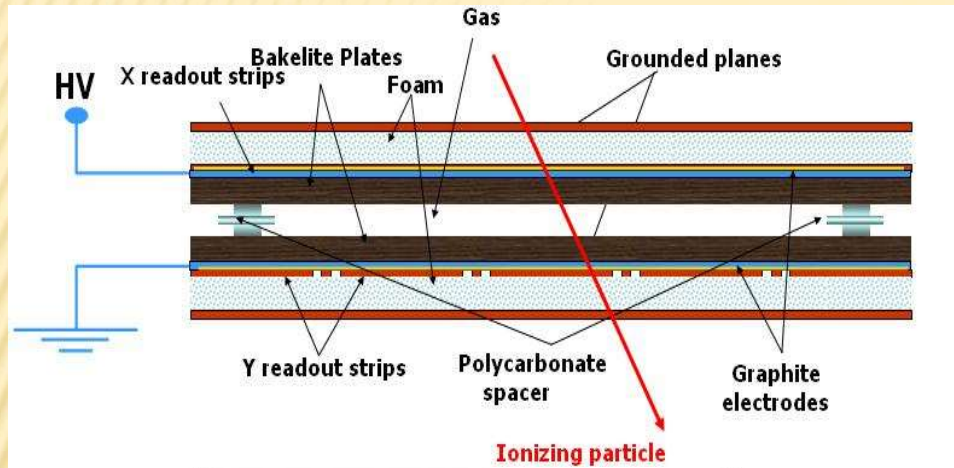
Trigger inefficiency

CONCLUSIONS

- ✗ SCRAP is in *routine* mode.
- ✗ 3PMTs (passive base) have been equipped for SSD tests.
- ✗ The SSD checks performed are:
 - + Number of Photoelectrons per MIP
 - + Light Tightness
 - + Homogeneity

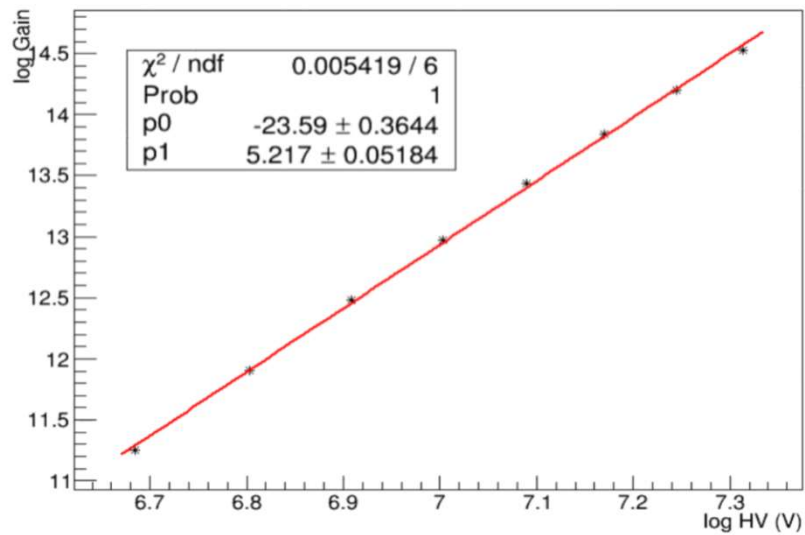
BACKUP

ATLAS RPC (RESISTIVE PLATE CHAMBER)

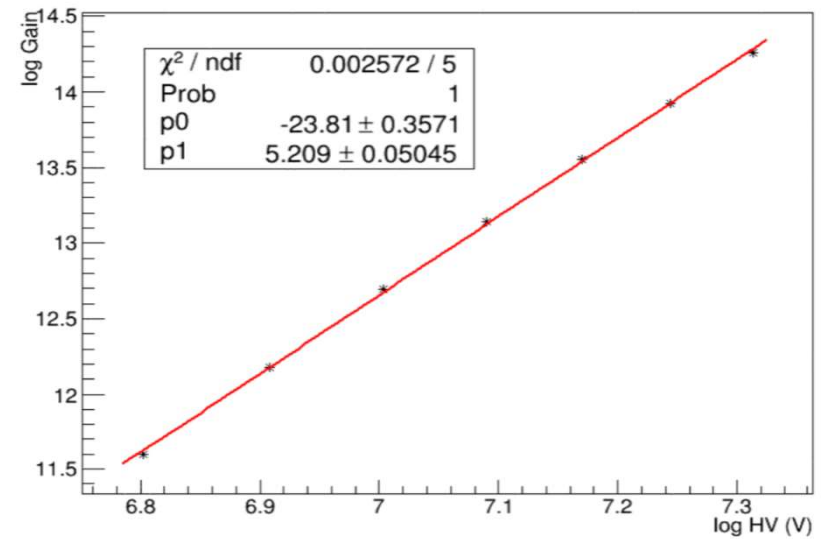


- Gap gas = 2 mm
 - $C_2H_2F_4$ (94.7%)
 - C_4H_{10} (5%)
 - SF_6 (0.3%)
- HV~10 KV Saturated avalanche regime
- Time response ~1 ns
- Readout panels: ϕ and η copper strip (pitch ~3 cm) .
- Spatial resolution (~ 1 cm) .
- Rate Capability ~1 kHz/cm²@Eff.>98%
- Gap volume size 4×1 m
- Electrodes: graphite layer + 2 mm thick bakelite ($\rho \sim 1 \div 4 \times 10^{10} \Omega cm$) .

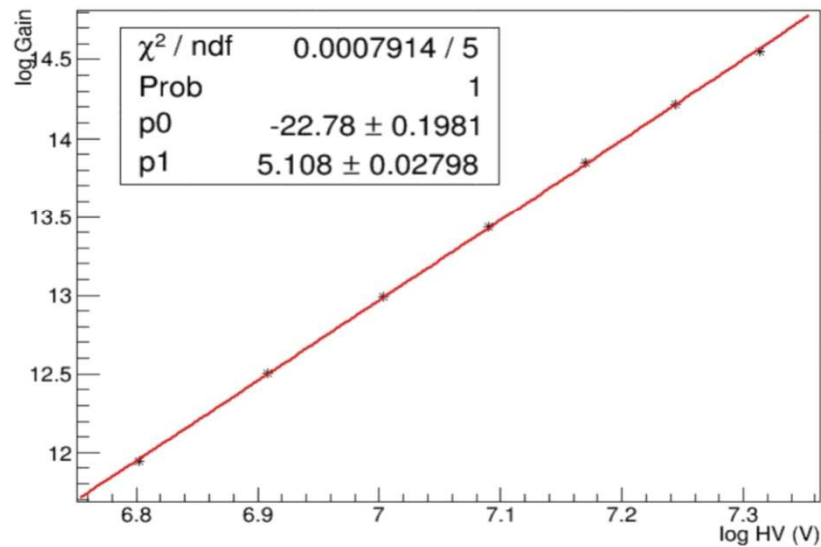
GAIN V_s HV CURVES OF PMT_s FIT



DD3851



DD4251



DD4151