

Beyond the Standard Model: phenomenology and LHC constraints of exotic dark matter

Gaël Alguero

LPSC Grenoble & LAPTh Annecy

Mid-thesis presentation, March 25, 2021

Supervised by Sabine Kraml and Genevieve Belanger



- From dark matter to LHC
- Setting constraints with SModelS and Madanalysis5
- Probing exotic dark matter (work in progress)

G.A., Araz, Fuks, Kraml, Waltenberger, Les Houches 2019 proceedings, contribution 15 [arXiv] G.A., Kraml, Waltenberger, CPC 264 (2021) 107909 [arXiv] G.A., Heisig, Khosa, Kraml, Kulkarni, Lessa, Neuhuber,

Reyes-González, Waltenberger, Wongel, TOOLS 2020 proceedings [arXiv]

Evidences for dark matter (DM)



From dark matter to LHC

- Many evidences hinting for a missing mass
 - Galaxy rotation curves
 - Galaxy clusters
 - Gravitational lensing
 - And others
- Must be massive, weakly interacting, dark



The DM theory landscape



From dark matter to LHC

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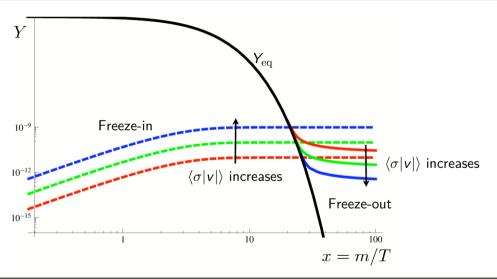


Nature 562, 5156 (2018)

Dark matter production in the universe



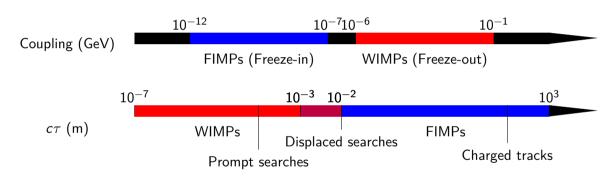
From dark matter to LHC



WIMPs, FIMPs and beyond



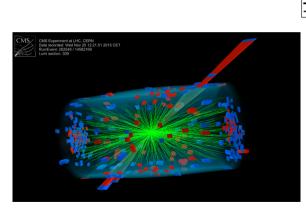
From dark matter to LHC

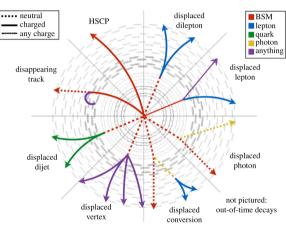


At the LHC: Missing energy and long lived signatures



From dark matter to LHC





How to use LHC results?



Setting constraints with SModelS and Madanalysis5

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- $\blacksquare \ \, \mathsf{Monte} \ \, \mathsf{Carlo} \ \, \mathsf{event} \ \, \mathsf{generation} \, + \, \mathsf{detector} \ \, \mathsf{simulation} \ \, \mathsf{(MadGraph} \, + \, \mathsf{MadAnalysis)}$
 - simulate the whole collision, detector effects and selection cutflow
 - captures differences in kinematical distributions
 - more generally applicable
 - very CPU-time consuming
- Using simplified model results (SModelS)
 - uses the efficiencies provided by experimentalists
 - assumes the cut acceptances are approx. the same as in the simplified models
 - much faster
 - suitable for large scans

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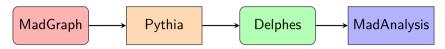
Reyes-González, Waltenberger, Wongel, TOOLS 2020 proceedings [arXiv]

Recasting workflow



Setting constraints with SModelS and Madanalysis5

 $\label{eq:mid-thesis} \mbox{ Mid-thesis presentation, March 25, $2021}$



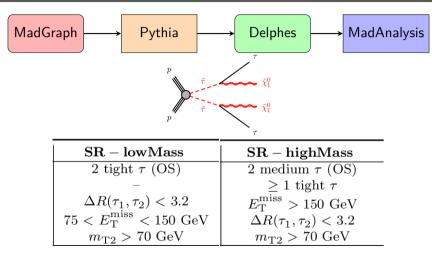
ATLAS-SUSY-2018-04

Recasting workflow



Setting constraints with SModelS and Madanalysis5

Mid-thesis presentation, March 25, 2021



ATLAS-SUSY-2018-04

Re-using simplified model results: SModelS

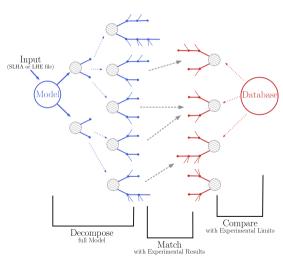


Setting constraints with SModelS and Madanalysis5

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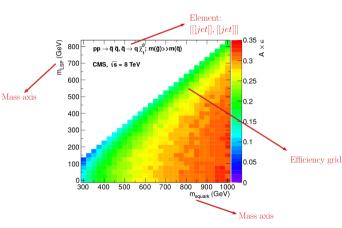
- Based on a general procedure to decompose BSM collider signatures featuring a \mathbb{Z}_2 -like symmetry into simplified-model topologies
- Large database of simplified-model results (currently \sim 100 ATLAS ans CMS searches)
- New generic treatment of width-dependent results with a variety of LLP analyses

See online SModelS documentation at smodels.readthedocs.io



Setting constraints with SModelS and Madanalysis5 Mid-thesis presental

- Upper limit (UL)
 - Constrains $\sigma \times \prod_i BR_i$
 - Only binary decision
- Efficiency map (EM)
 - Allows to sum contributions from several topologies
 - Can compute a likelihood (confidence level)



Combining signal regions (SRs)



Setting constraints with SModelS and Madanalysis5

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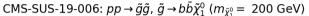
- For a proper statistical evaluation
 - need to compute a likelihood
 - without specific likelihood information : can only compute for each signal region (SR)

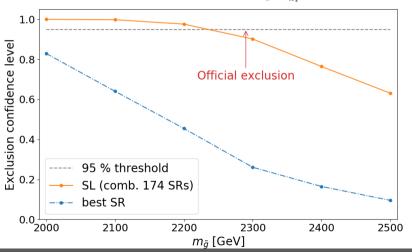
$$\mathcal{L}(\mu, \theta | D) = \frac{(\mu s + b + \theta)^{n_{obs}} e^{-(\mu s + b + \theta)}}{n_{obs}!} \exp\left(-\frac{\theta^2}{2\delta^2}\right)$$

 \Rightarrow use the "most sensitive" SR

- Simplified likelihoods : global background uncertainty [CMS-NOTE-2017-001]
 - summarized into a covariance matrix
 - one Gaussian error for each SR
- Full likelihoods : full statistical descriptions of analyses [ATL-PHYS-PUB-2019-029]
 - encapsulate the detailed information about the analysis (detailed systematic uncertainties)
 - reproduces exactly the experimental analysis and allows for a more precise reinterpretation

Setting constraints with SModelS and Madanalysis5





Les Houches 2019 proceedings, contribution 15 [arXiv]

- CMS-SUS-16-039
- CMS-SUS-17-001
- CMS-SUS-19-006
- CMS-SUS-16-048

Full ATLAS likelihoods



Setting constraints with SModelS and Madanalysis5

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- Serialization of HistFactory workspaces under the JSON format
- Provides observed signals, expected backgrounds and systematic uncertainties as in the experimental analysis
- Can be used in RooFit or pyhf
 (a pure-python implementation of HistFactory)

ATL-PHYS-PUB-2019-029

```
"channels": [
   "name": "SR1cut cuts".
   "samples": [=
                                  Signal regions
   "name": "SR2cut cuts",
   "samples": [
       "data": [
         2.570836067199707
                                  Background
       "modifiers": [-
       "name": "Boson Staus"
                                  contributions
       "data": [
         0.044047050178050995
       "modifiers": [=
       "name": "Z Staus"
```

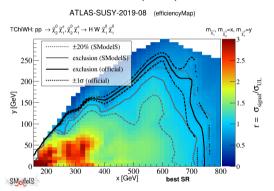
Implementation of a SModelS interface to pyhf

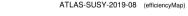


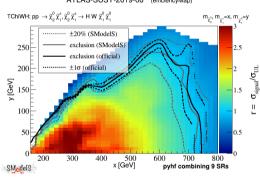
Setting constraints with SModelS and Madanalysis5

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CPC 264 (2021) 107909 [arXiv] and TOOLS2020 proceedings [arXiv]







Standard "best SR" procedure

pyhf likelihoods

⇒ ATLAS-SUSY-2018-04, ATLAS-SUSY-2018-31 and more to come

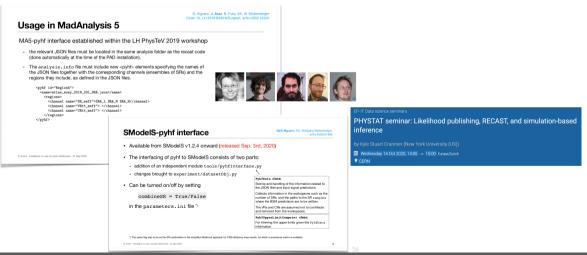
Cited in CERN seminar by Kyle Cranmer



Setting constraints with SModelS and Madanalysis5

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USAGE IN PHENO RECASTING TOOLS



Outline

Probing exotic dark matter (work in progress)

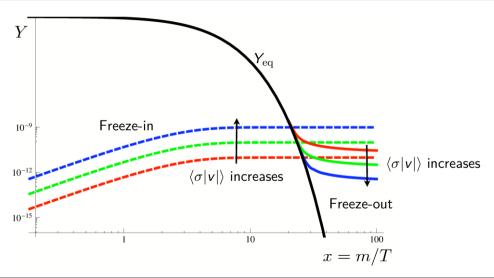
1 From dark matter to LHC

2 Setting constraints with SModelS and Madanalysis5

3 Probing exotic dark matter (work in progress)



Probing exotic dark matter (work in progress)



Probing exotic dark matter (work in progress)

- Standard freeze-out equation $\dot{n}_{\chi} + 3Hn_{\chi} = -\langle \sigma | v | \rangle \left(n_{\chi}^2 n_{\chi, \mathrm{eq}}^2 \right)$
- Freeze-in
- $$\begin{split} & \text{coscattering} \\ & \dot{n}_{\psi} + 3Hn_{\psi} = -\langle \sigma | v | \rangle \left(n_{\psi}^2 n_{\psi, \text{eq}}^2 \right) \\ & \Gamma_{\psi \to \chi} \left(n_{\chi} n_{\psi} \frac{n_{\chi}^{\text{eq}}}{n_{\psi}^{\text{eq}}} \right) \\ & \dot{n}_{\chi} + 3Hn_{\chi} = \Gamma_{\psi \to \chi} \left(n_{\chi} n_{\psi} \frac{n_{\chi}^{\text{eq}}}{n_{\psi}^{\text{eq}}} \right) \end{aligned}$$

$$\chi$$
 SM SM χ SM χ SM

SM

SM

 χ : dark matter

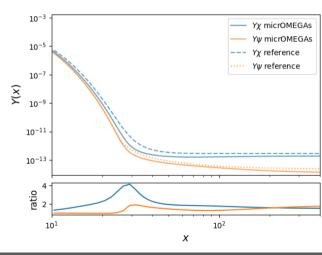
 ψ : heavier odd particle

R.T.D'agnolo et al., Phys. Rev. Lett. 119, 061102 (2017) [arXiv] M.Garny et al., Phys. Rev. D 96, 103521 (2017) [arXiv]

 $M = 500 \text{ GeV}, m = 490 \text{ GeV}. \theta = 1.22 \times 10^{-6}$

- Testing coscattering with a Singlet-Triplet model
 - Standard Model $+\chi^0 + \{\psi^0, \psi^{\pm}\}\$
- with a compressed spectrum
 - small mass splitting
 - small couplings

F.Brümmer, JHEP 2001 (2020) 113



[arXiv]

Scotogenic model



Probing exotic dark matter (work in progress)

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Ernest Ma, Phys.Rev.D73:077301,2006 [arXiv]

- Standard Model extended with
 - an inert Higgs doublet
 - and right-handed neutrinos

$$\begin{pmatrix} H_1^{\pm} \\ H_1^{0} \end{pmatrix}, \begin{pmatrix} H_2^{\pm} \\ H_2^{0} \end{pmatrix}, \begin{pmatrix} N_1 \\ N_2 \\ N_3 \end{pmatrix}$$

• odd under a \mathbb{Z}_2 -symmetry

Field	Generations	$SU(3)_c$	$\mathrm{SU}(2)_{\mathrm{L}}$	$U(1)_{Y}$	\mathbb{Z}_2
ℓ_{L}	3	1	2	-1/2	+
e_R	3	1	1	-1	+
H_1	1	1	2	1/2	+
H_2	1	1	2	1/2	_
N	3	1	1	0	_

$$H_1 \xrightarrow{\mathbb{Z}_2} H_1$$

$$H_2 \xrightarrow{\mathbb{Z}_2} -H_2$$

- radiative neutrino masses
- provides different DM candidates
 - different production mechanisms (freeze-in, freeze-out, ...)
 - probe with the appropriate LHC signatures (prompt, long-lived, ...)

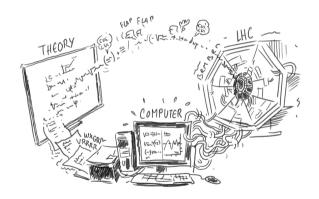
Conclusions

Probing exotic dark matter (work in progress)

- simplified likelihoods implemented in MadAnalysis5
 - ability to use CMS covariance matrices
 - database extended with all usable covariance matrices
- SModelS/pyhf interface from SModelS v1.2.4 onwards
 - ability to use full likelihoods from ATLAS
 - lacktriangle added EMs + JSON files for three 139 fb⁻¹ analyses
- more accurate reinterpretations, more reliable statistical evaluation e.g. for fits
- work in progress
 - coscattering in micrOMEGAs
 - scotogenic model



Thanks for your attention!

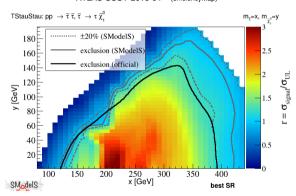


©Lison Bernet

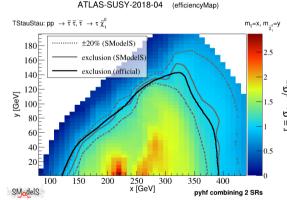
Backups

arXiv:2009.01809 (pyhf interface)

ATLAS-SUSY-2018-04 (efficiencyMap)



ATI AS-SUSY-2018-04



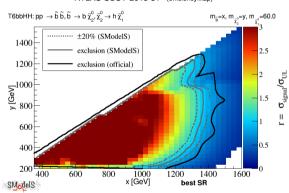
Standard "best SR" procedure

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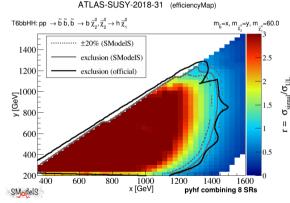
Backups

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ATLAS-SUSY-2018-31 (efficiencyMap)



ATLAS-SUSY-2018-31



Standard "best SR" procedure

pyhf likelihoods



$$p(n, a|\eta, \chi) = \prod_{\substack{c \in channels \ b \in bins}} \operatorname{Pois}\left(n_{cb}|\nu_{cb}(\eta, \chi)\right) \underbrace{\prod_{\chi} c_{\chi}(a_{\chi}|\chi)}_{\text{Constraints}}$$
(1)

relates the observed events and auxiliary data (n, a) to the free and constrained parameters (η, χ)

$$\nu_{cb}(\eta,\chi) = \sum_{s \in samples} \nu_{scb}(\eta,\chi) = \sum_{s \in samples} \underbrace{\prod_{\kappa} \kappa_{scb}(\eta,\chi)}_{\text{Mujtiplicative modifiers}} \left(\nu_{scb}^{0}(\eta,\chi) + \underbrace{\sum_{\Delta} \Delta_{scb}(\eta,\chi)}_{\text{Additive modifiers}} \right)$$

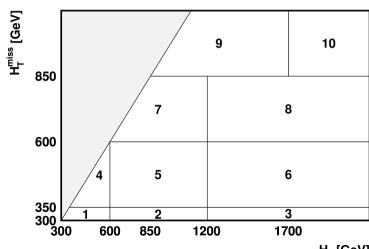
ATL-PHYS-PUB-2019-029

Backups

$$\mathcal{L}(\mu,\theta|D) = \prod_{i=1}^{N} \frac{(\mu s_i + b_i + \theta_i)^{n_{obs}^i} e^{-(\mu s_i + b_i + \theta_i)}}{n_{obs}^i!} e^{-(\mu s_i + b_i + \theta_i)} e^{-(\mu s_i + b_i + \theta_i)}$$

with a product over all N SRs where V is the covariance matrix

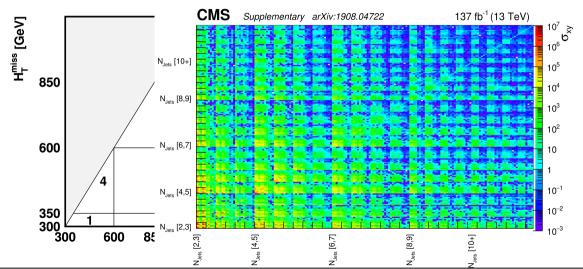
Backups



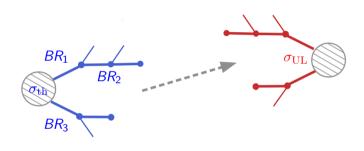
CMS-SUS-19-006 covariance matrix



Backups Mid-thesis presentation, March 25, 2021

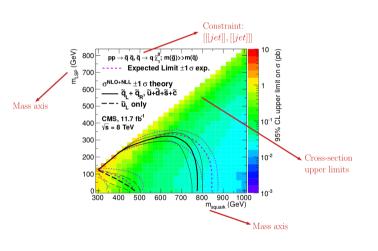


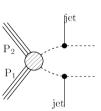




$$\sigma_{\rm th} \times \textit{BR}_1 \times \textit{BR}_2 \times \textit{BR}_3$$
 to compare with $\sigma_{\rm UL}$

Backups Mid-thesis presentation, March 25, 2021

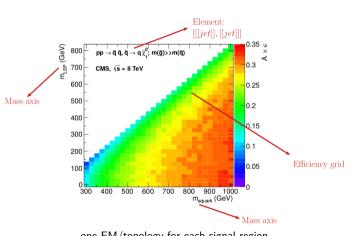




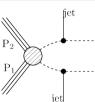
- Constrains $\sigma \times \prod_j BR_j$ per topology
- Only binary decision : excluded or not



Backups Mid-thesis presentation, March 25, 2021



one EM/topology for each signal region



- Constrains $\sum_{i} A_{i} \epsilon_{i} \times \sigma_{i} \times \prod_{i} BR_{i}$ per signal region
- Can sum contributions from several topologies
- Can compute a likelihood (exclusion confidence level) Efficieny map

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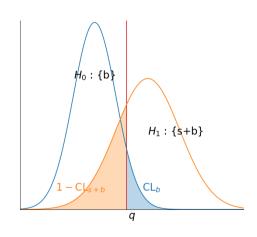
Backups

- Testing hypothesis H_0 : {only background} against H_1 : {background + new signal}
- Definition of a log likelihood ratio

$$q = -2\log\left(rac{\mathcal{L}(\mathrm{data}|H_1)}{\mathcal{L}(\mathrm{data}|H_0)}
ight)$$

■ Computing a robust p-value ratio

$$CL_s = \frac{CL_{s+b}}{CL_b}$$



$$N = A \epsilon \sigma f$$

- lacksquare A and ϵ : acceptance and efficiency
- lacksquare σ and $\mathcal L$: cross section and integrated luminosity