

Enhancing CMB+kSZ Power Spectrum Measurements by Removing CIB and tSZ Contamination with LSS

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Grenoble mm Universe, June 2023

Work with Ola Kusiak and Colin Hill



Motivation: measuring kSZ is challenging. Can we improve it by removing tSZ and CIB?

Idea: Use the external LSS data that is correlated with both CIB and tSZ to remove those contaminants to enhance CMB+kSZ measurements using ILC methods

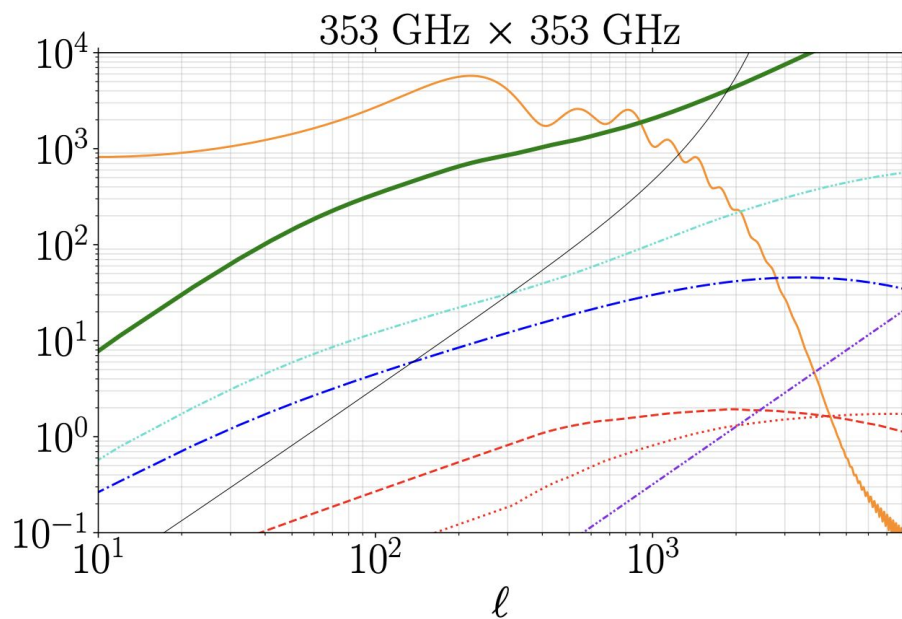
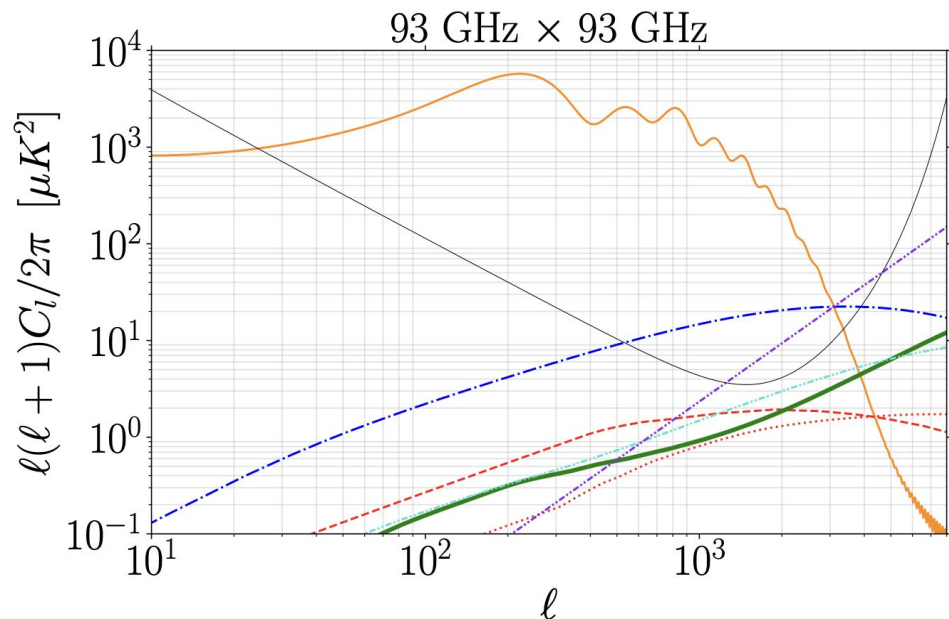
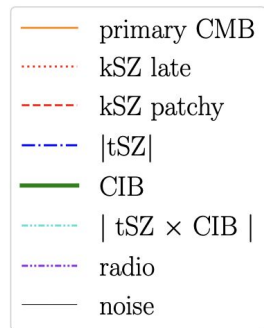
Key: Two-point correlation function of LSS with CMB + kSZ vanishes

- electron velocity as likely to be positive as to be negative, thus kSZ-LSS vanishes

Modeling choices

Modeling: Overview of Sky Models

Analytically model microwave sky as primary lensed **CMB**, **kSZ**, **tSZ**, **CIB**, and **radio** source signals, as well as **detector and atmospheric noise** for Simons Observatory (SO) and Planck-like experiment at 8 frequencies from 93 to 353 GHz.



Computed in the halo model with [class_sz](#) (Boris Bolliet)

Modeling: *unWISE* Galaxies

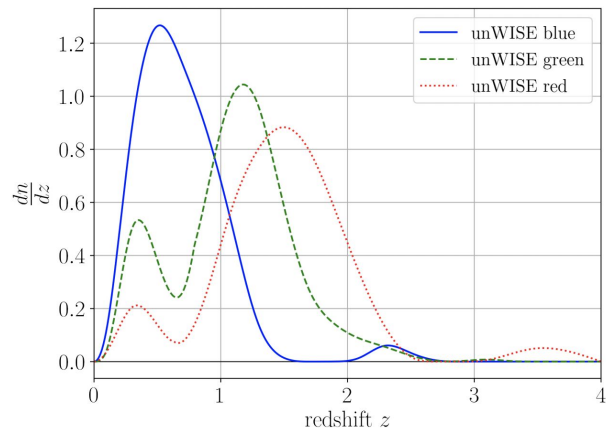
- *unWISE* galaxy catalog (Krolewski+ 2020):

- Based on *WISE* and *NEOWISE*
- Over 500 million galaxies on the full sky
- 3 subsamples: **blue** ($z=0.6$), **green** ($z=1.1$), and **red** ($z=1.5$)
- Coarse dn/dz via cross-correlation with spectroscopic surveys

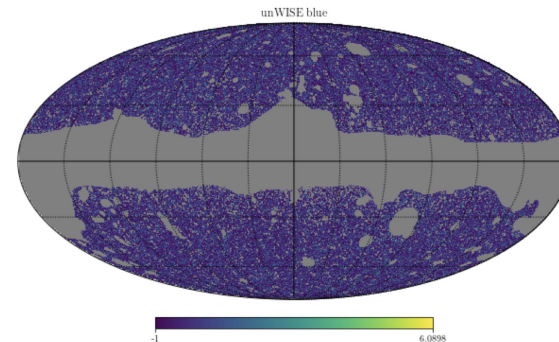
Number density of galaxies

<i>unWISE</i>	\bar{z}	δ_z	\bar{n}_g [deg ⁻²]
blue	0.6	0.3	3409
green	1.1	0.4	1846
red	1.5	0.4	144

Redshift Distribution



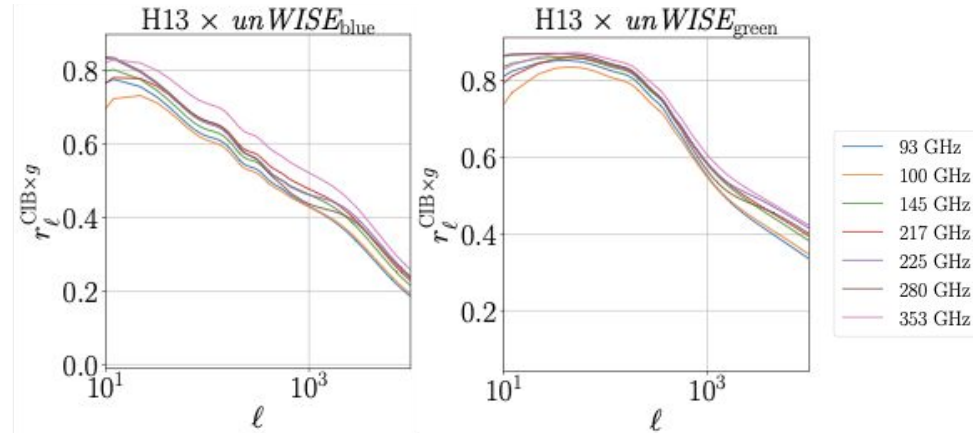
Galaxy overdensity map (*unWISE* blue)



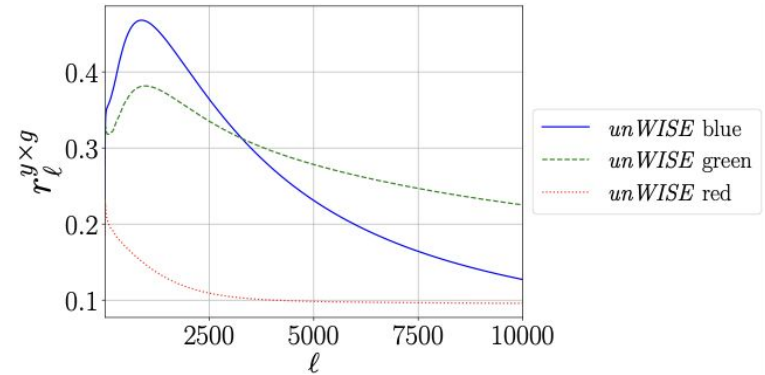
We model *unWISE* halo occupation distribution (HOD) = galaxy clustering model based on Kusiak+ 2022

unWISE correlates with CIB and tSZ fields

CIB x *unWISE*:



Compton-y x *unWISE*:



unWISE

- great tracer of the CIB
- also partially traces tSZ

Internal Linear Combination (ILC): Theory

Standard ILC

Internal Linear Combination (ILC)—standard method to construct a map of the signal of interest and optionally deproject (=remove) some other contaminant components

$$\hat{T}_{\ell m}^{\text{ILC}} = \sum_i w_{\ell}^i T_{\ell m}^i$$

temperature map at i^{th} frequency

Minimization subject to **one constraint**:

- Minimize variance

$$\sigma_{\hat{T}_{\ell m}^{\text{ILC}} \hat{T}_{\ell m}^{\text{ILC}}}^2 = \sum_{i,j} w_{\ell}^i w_{\ell}^j \left(\hat{R}_{\ell} \right)_{ij}$$

Frequency-frequency
covariance matrix from
data

- Unit response to signal of interest

$$\sum_i w_{\ell}^i a_i = 1$$

SED of signal of interest

Solution for the weights:

$$w_{\ell}^i = \frac{\left(\hat{R}_{\ell}^{-1} \right)_{ij} a_j}{\left(\hat{R}_{\ell}^{-1} \right)_{km} a_k a_m}$$

Constrained ILC

Minimization subject to **two constraints**:

- Minimize variance

$$\sigma_{\hat{T}_{\ell m}^{\text{ILC}} \hat{T}_{\ell m}^{\text{ILC}}}^2 = \sum_{i,j} w_{\ell}^i w_{\ell}^j \left(\hat{R}_{\ell} \right)_{ij}$$

- **Unit response** to signal of interest

$$\sum_i w_{\ell}^i a_i = 1$$

- Deproject one contaminant

$$\sum_i w_{\ell}^i b_i = 0$$

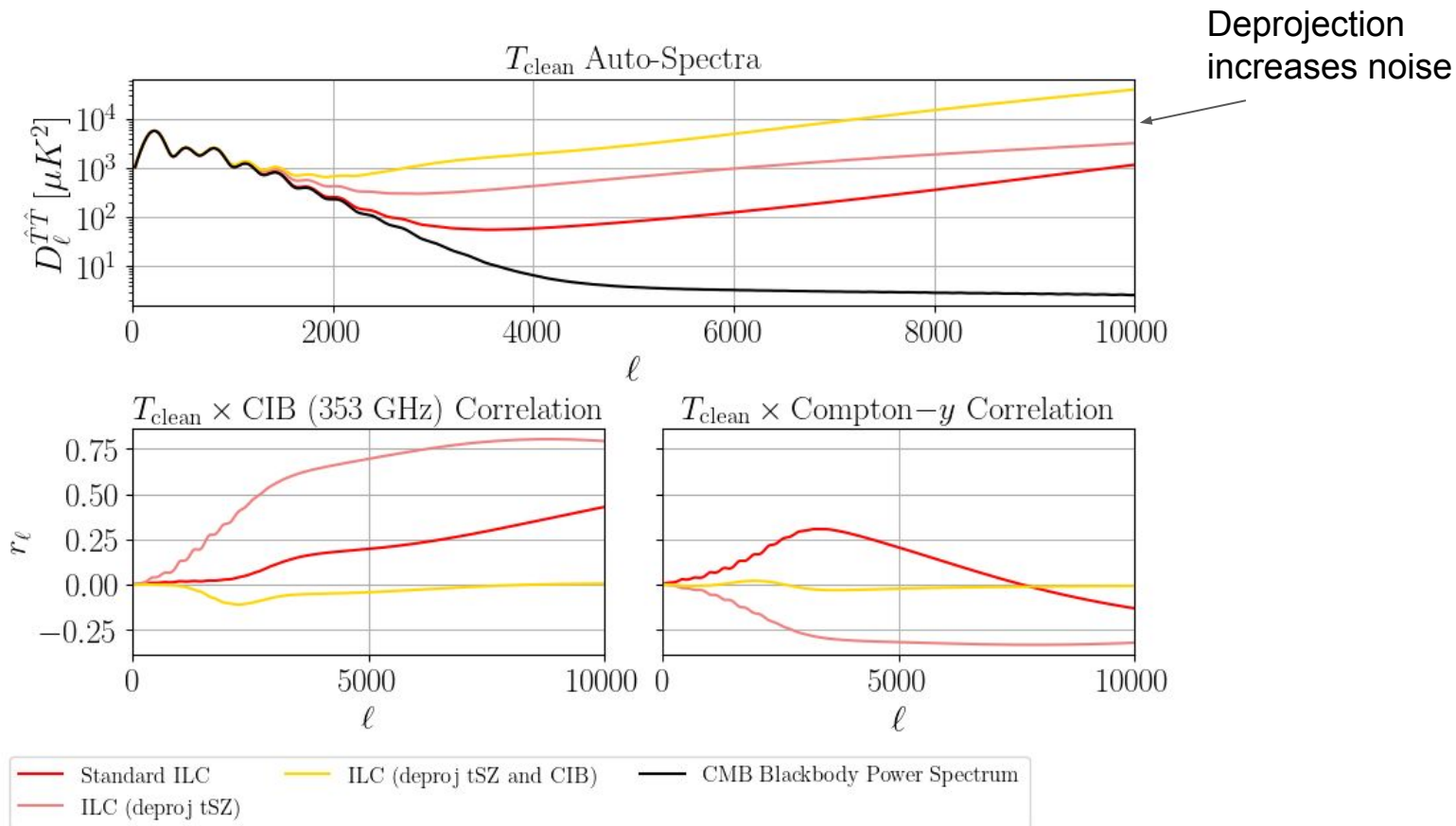
SED of component to deproject

Solution for the weights:

$$w_{\ell}^j = \frac{\left(b_k (\hat{R}_{\ell}^{-1})_{kl} b_l \right) a_i (\hat{R}_{\ell}^{-1})_{ij} - \left(a_k (\hat{R}_{\ell}^{-1})_{kl} b_l \right) b_i (\hat{R}_{\ell}^{-1})_{ij}}{\left(a_k (\hat{R}_{\ell}^{-1})_{kl} a_l \right) \left(b_m (\hat{R}_{\ell}^{-1})_{mn} b_n \right) - \left(a_k (\hat{R}_{\ell}^{-1})_{kl} b_l \right)^2}$$

- Solution can be generalized to **N** deprojected components
- Generally, (except tSZ) we don't know the exact SED of the contaminant
- No free lunch!
 - deprojecting/adding a constraint increases the noise of the final map

Standard and Constrained ILC Results



New Methods

Overview of New Methods

1. ILC with tracers as additional input “frequency” maps
2. de-CIB
3. ILC requiring tracers to have zero correlation with the cleaned map

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ILC with Tracers as Additional Input “Frequency” Maps

- Use the fact that the CMB and kSZ fields have zero “response” in the tracer maps

$$\hat{T}_{\ell m}^{\text{ILC}} = \sum_{i=1}^N w_{\ell}^i T_{\ell m}^i + \sum_{i=1}^{N_g} w_{\ell}^{N+i} g_{i,\ell m}$$

Spectral response
vector for signal of
interest

$$\mathbf{a} \leftarrow \mathbf{a} + \underbrace{[0, \dots, 0]}_{N_g}$$

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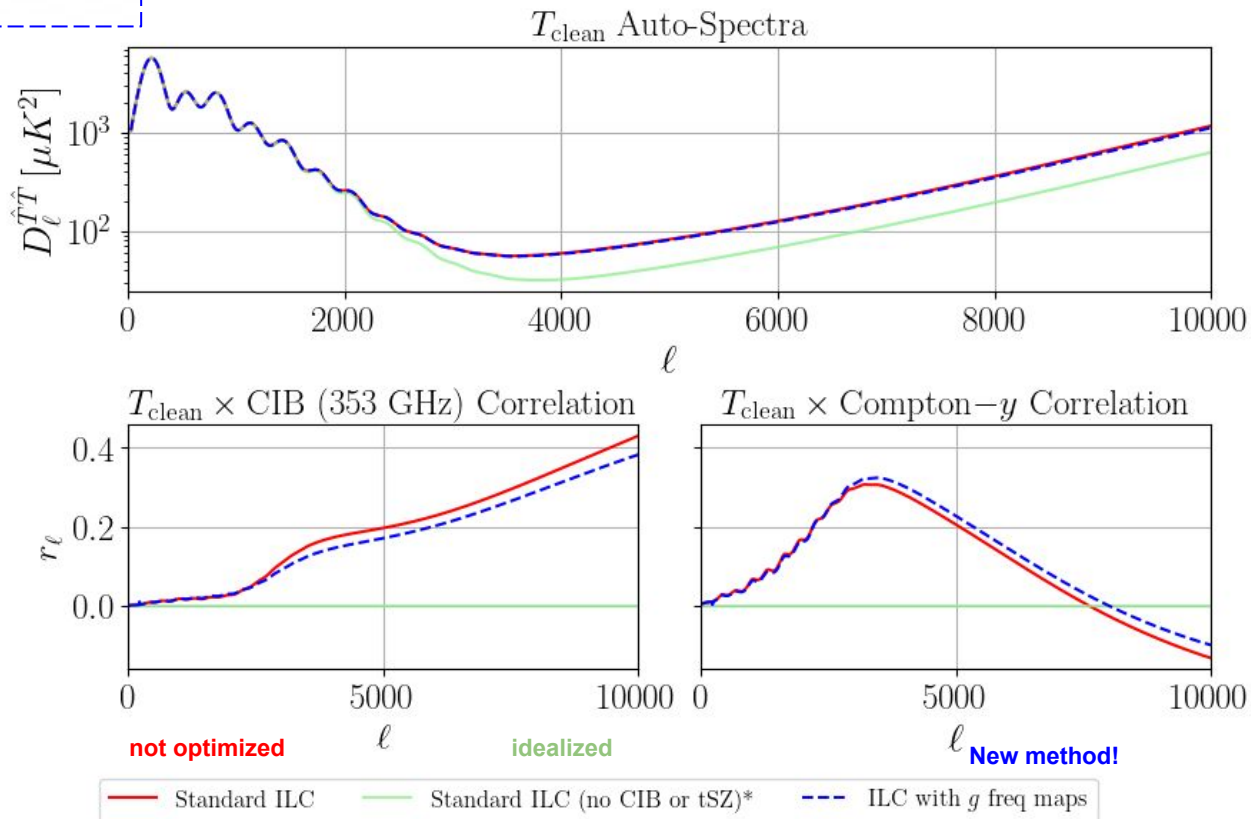
Standard ILC solution with modified covariance

$$w_{\ell}^i = \frac{\left(\hat{R}_{\ell}^{-1}\right)'_{ij} a_j}{\left(\hat{R}_{\ell}^{-1}\right)'_{km} a_k a_m}$$

$$\left(\hat{R}^{\ell}\right)' = \begin{bmatrix} \hat{R}^{\ell} & X_{\ell} \\ X_{\ell}^T & G_{\ell} \end{bmatrix}$$

ILC with Tracers as Additional Input “Frequency” Maps: Results

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de-CIB

Idea taken from delensing, following
Yu, Hill, Sherwin 2017

1. Find the linear combination of LSS tracer samples most highly correlated with CIB at i^{th} frequency:

$$g_{lm}^i = \sum_a c_{a,l}^i g_{lm}^a$$

Optimal combination of tracers at i^{th} frequency Different tracers, e.g. *unWISE* blue, green, red

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2. Starting from an ILC map, subtract the fraction of tracers that's contained in the CIB at each frequency:

$$T_{\ell m}^{\text{clean}} = \sum_i w_{\ell}^i (T_{\ell m}^i - f_{\ell}^i g_{\ell m}^i)$$

ILC

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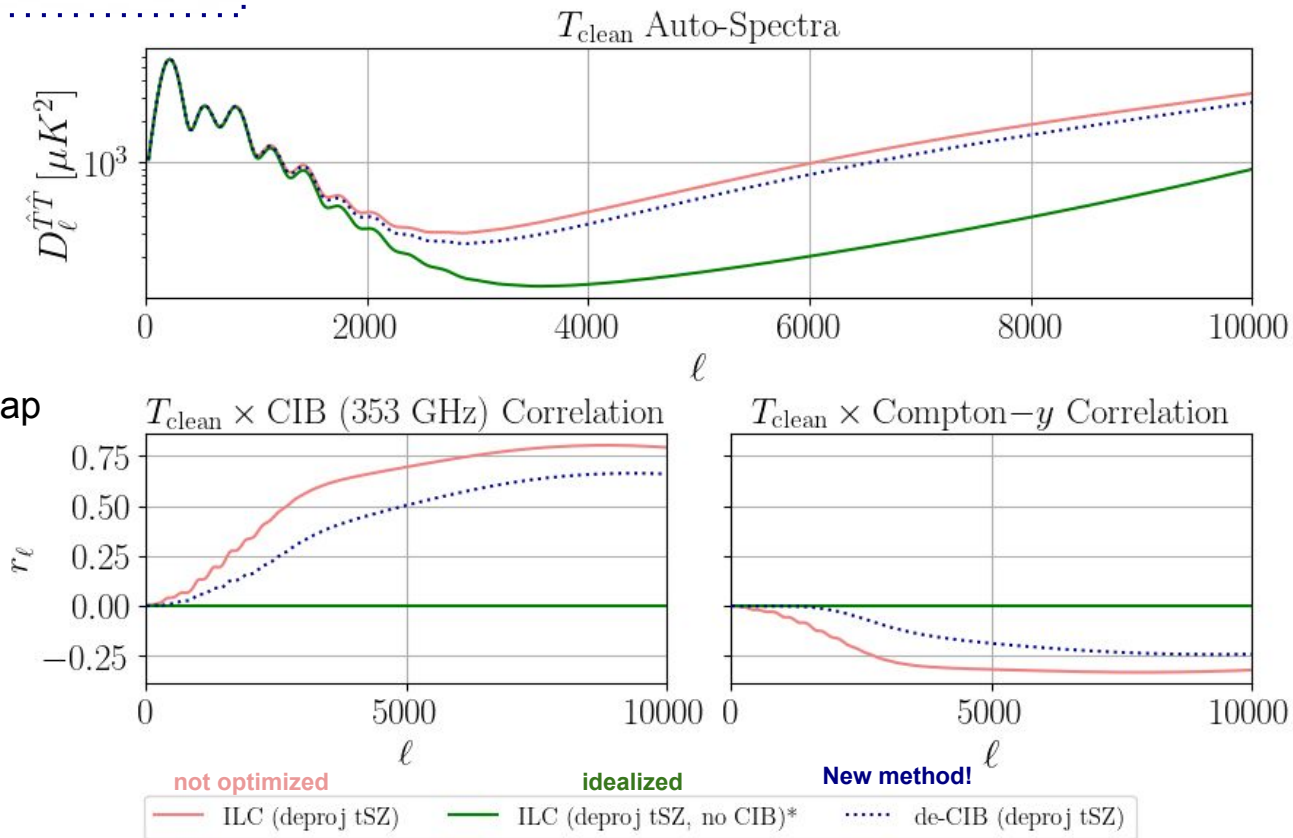
The fraction of tracers contained in the CIB at each frequency (from delensing: minimize the variance of the final map) is

$$f_{\ell}^i \equiv \frac{C_{\ell}^{\text{CIB}_i, g^i}}{C_{\ell}^{g^i, g^i}} = \frac{\sum_a c_{a,\ell}^i C_{\ell}^{\text{CIB}_i, g_a}}{\sum_{a,b} c_{a,\ell}^i c_{b,\ell}^i C_{\ell}^{g_a g_b}}$$

de-CIB: Results

$$T_{lm}^{\text{clean}} = \sum_i w_\ell^i (T_{lm}^i - f_\ell^i g_{lm}^i)$$

de-CIB is applied to
tSZ-deprojected ILC map



Overview of New Methods

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ILC with Zero-Tracer-Correlation Constraint

Start from an ILC, and explicitly require the output map to have zero cross-correlation with the tracer map:

$$\langle T_{\ell m}^{\text{ILC}} g_{\ell m} \rangle = 0 \iff \sum_i w_{\ell}^i \langle T_{\ell m}^i g_{\ell m} \rangle = 0$$

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Def. $c_i^{\ell} \equiv \langle T_{\ell m}^i g_{\ell m} \rangle = C_{\ell}^{ig}$



$$\sum_i w_{\ell}^i c_i^{\ell} = 0$$

similar to usual constrained ILC!

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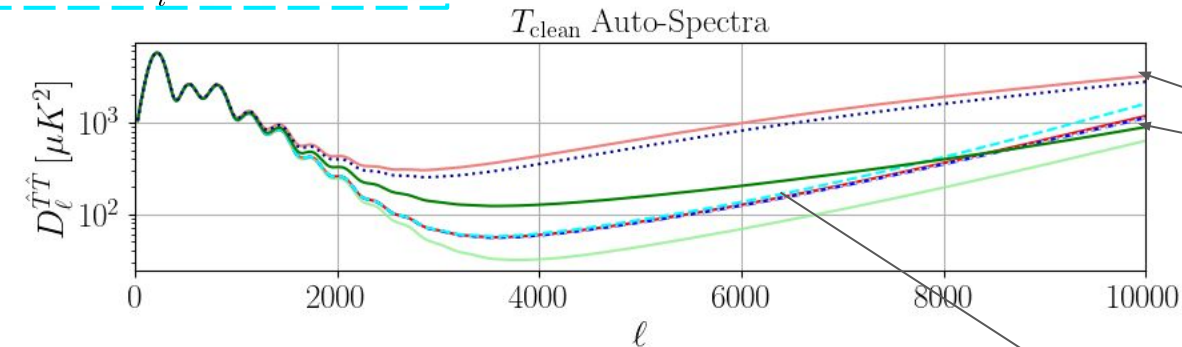
- This is a **spatial** deprojection instead of the usual **spectral** deprojection
- Can be applied to a (spectral) tSZ-deprojected ILC or can spatially deproject both tSZ and CIB, i.e., using

$$b_i^{\ell} \equiv \langle T_{\ell m}^i g_{\ell m}^{\text{tSZ}} \rangle \text{ and } c_i^{\ell} \equiv \langle T_{\ell m}^i g_{\ell m}^{\text{CIB}} \rangle$$

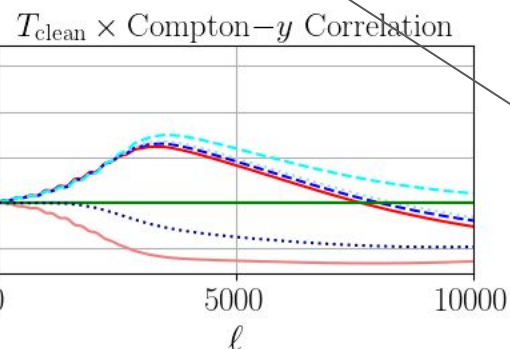
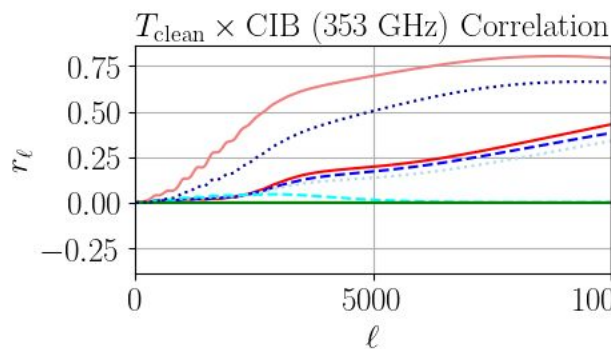
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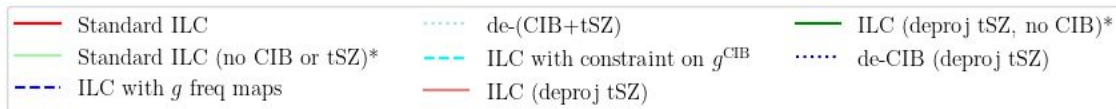
(no tSZ-deprojection)



Compare with 2 constraints:
 ILC (deproj tSZ) and ILC
 (deproj tSZ, no CIB)



Spatial deprojection –
 noise behaves
 differently!



Summary

- Three new methods to remove CIB and tSZ contamination using LSS tracers to enhance kSZ+CMB
- Methods have various trade-offs depending on metrics one cares about
- Theoretical models provided for forecasting purposes but methods can be applied to data
- Results here are just for *unWISE*—can actually combine several tracers (DES, BOSS/eBOSS, 2MASS, or CMB/galaxy lensing)!
- Future surveys (LSST, DESI, Euclid) will have more galaxies and give even better results
- Future directions: test on sims and apply to data (ACT + *unWISE*)

arXiv:2303.08121

Backup Slides

Modeling: CIB

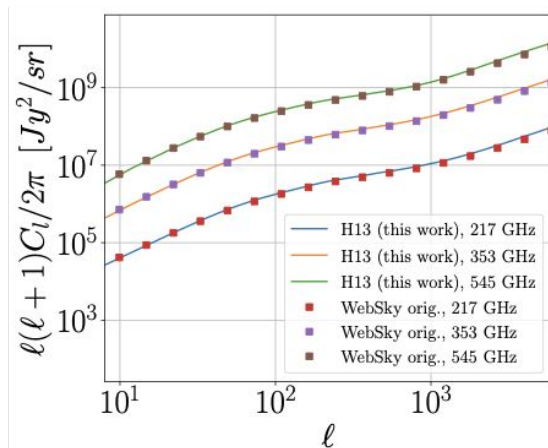
Two models based on the standard Shang et al.

CIB model:

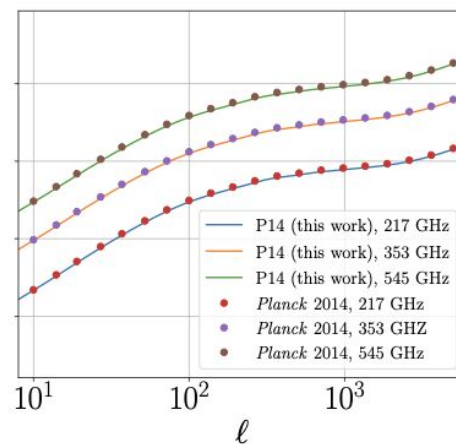
- Planck 2014 paper (P14)
- Herschel 2013 paper (H13); also in WebSky simulations

Parameter	Parameter description	Value
L_0	Normalization of $L-M$ relation	7.0×10^{-8} Jy Mpc ² /M _⊙ /Hz
α^{CIB}	Redshift evolution of dust temperature	0.36
T_0	Dust temperature at $z = 0$	24.4 K
β^{CIB}	Emissivity index of SED	1.75
γ^{CIB}	Power law index of SED at high frequency	1.7
$\log_{10}(M_{\text{eff}}^{\text{CIB}}/M_{\odot})$	Most efficient halo mass	12.6
$\log_{10}(M_{\text{min}}^{\text{CIB}}/M_{\odot})$	Minimum halo mass to host a galaxy	10
σ_{L-M}^2	Distribution of of halo masses sourcing CIB emission	0.5
δ_{CIB}	Redshift evolution of $L-M$ relation	3.6
z_p	Plateau redshift of $L-M$ relation	10 ¹⁰⁰

H13



P14



Different values of the Shang et al. CIB parameters

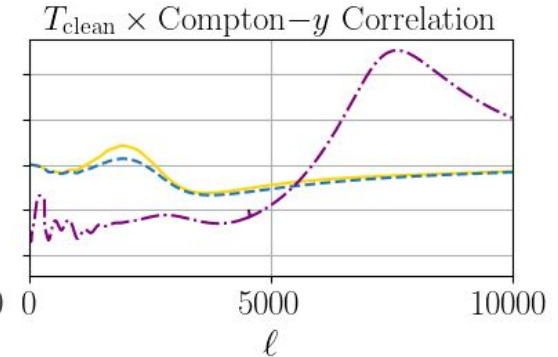
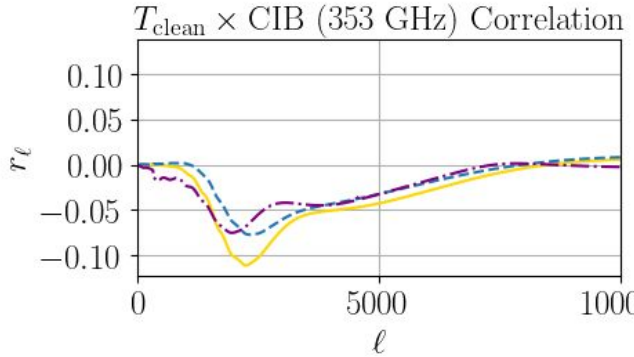
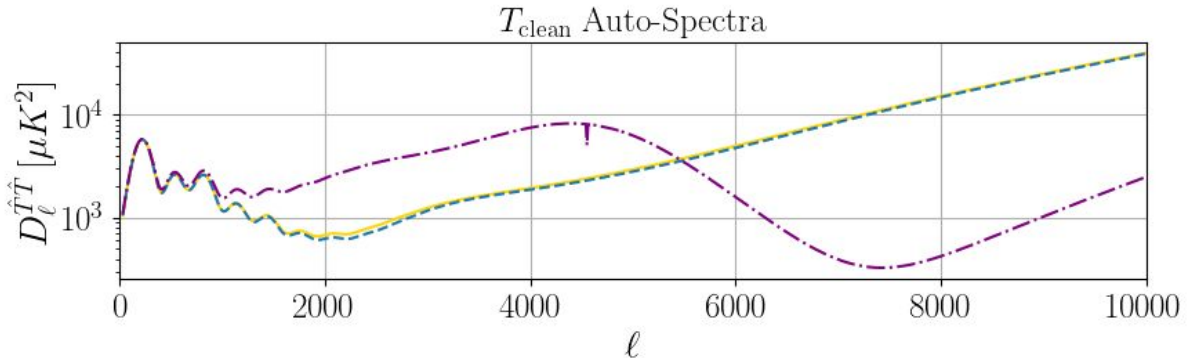
ILC w/ 2 Zero-Tracer-Corr Constraints (3 constraints): Results

$$\langle T_{\ell m}^{\text{ILC}} g_{\ell m} \rangle = 0 \iff \sum_i w_{\ell}^i \langle T_{\ell m}^i g_{\ell m} \rangle = 0.$$

The shape of the new ILC (purple)

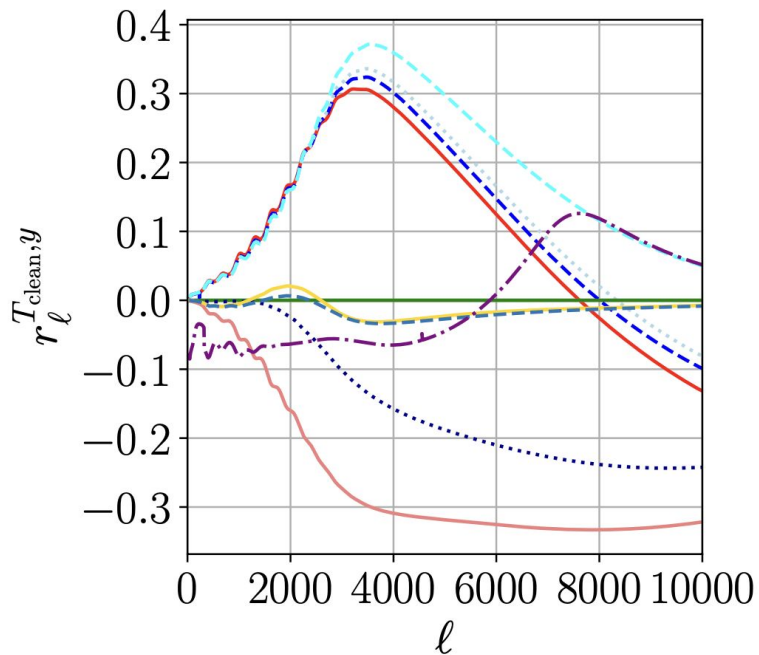
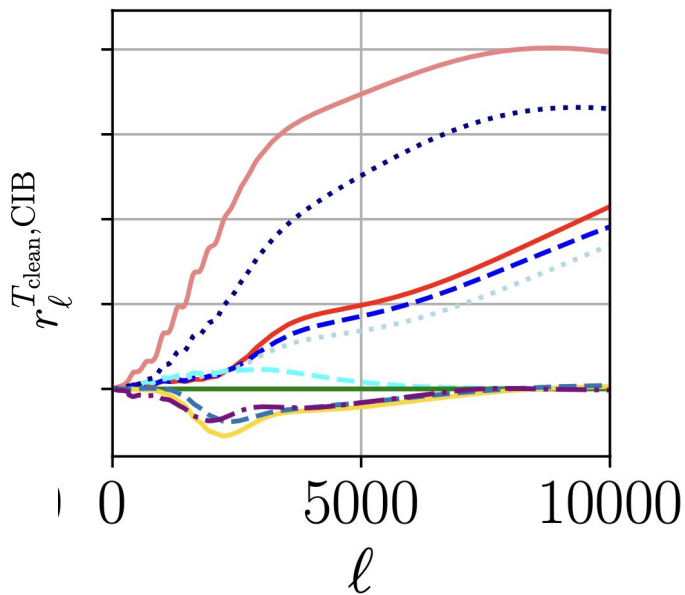
→ CIB and tSZ highly correlated at small scales

→ spatially deprojecting two highly correlated tracers allows both to be deprojected for the price of one



— ILC (deproj tSZ and CIB)
 - - - ILC with constraint on g^{CIB} (deproj tSZ)
 - · - · ILC with constraints on g^{CIB} and g^{tSZ}

353 GHz

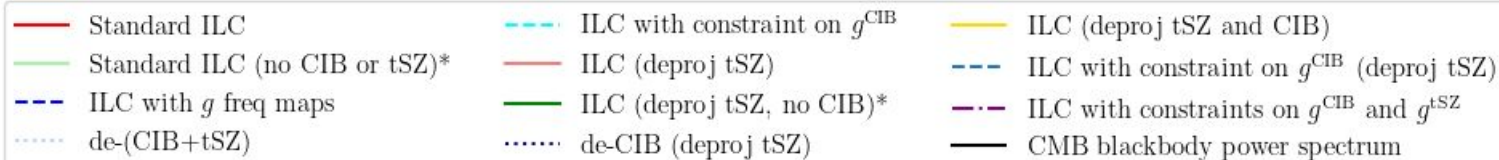
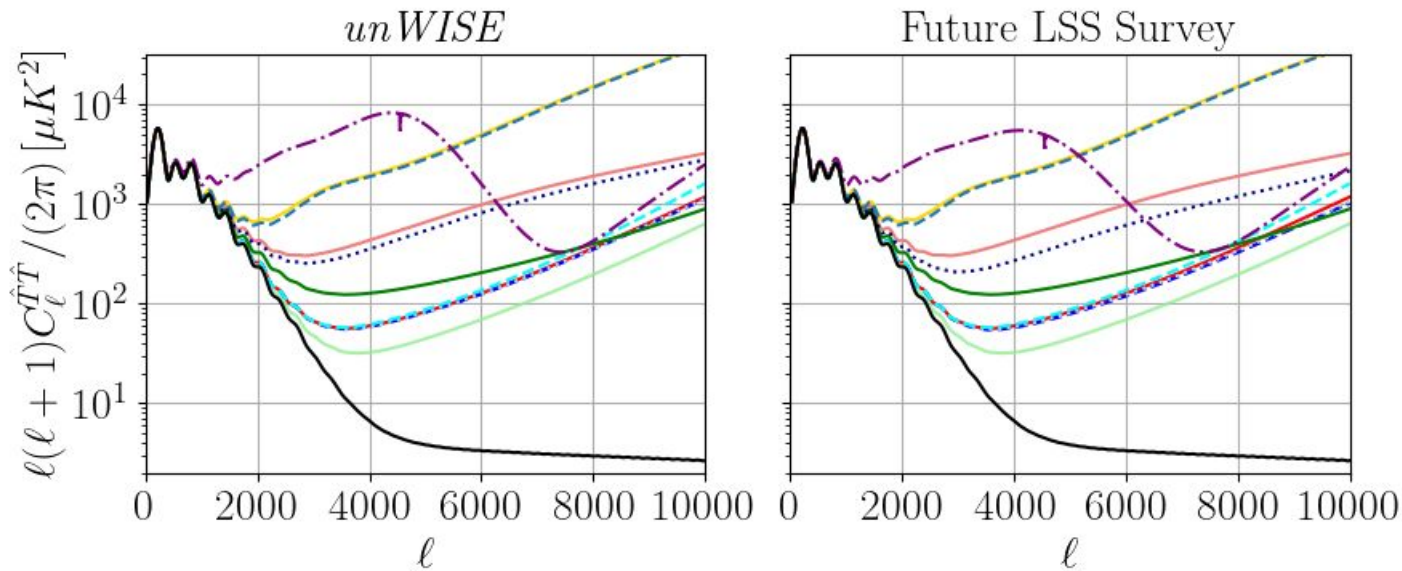


- | | | |
|--|---|---|
| — Standard ILC | - - - ILC with constraint on g^{CIB} | — ILC (deproj tSZ and CIB) |
| — Standard ILC (no CIB or tSZ)* | - - - ILC (deproj tSZ) | - - - ILC with constraint on g^{CIB} (deproj tSZ) |
| - - - ILC with g freq maps | — ILC (deproj tSZ, no CIB)* | - - - ILC with constraints on g^{CIB} and g^{tSZ} |
| ⋯ de-(CIB+tSZ) | ⋯ de-CIB (deproj tSZ) | |

SNR

	H13 CIB Model (using <i>unWISE</i>)	P14 CIB Model (using <i>unWISE</i>)	H13 CIB Model (future LSS survey)	P14 CIB Model (future LSS survey)
Standard ILC	115.32	115.25	115.32	115.25
Standard ILC (no CIB or tSZ)*	202.74	202.74	202.74	202.74
ILC with g freq maps	116.80	116.75	120.04	118.37
de-(CIB+tSZ)	116.49	116.69	117.90	118.11
ILC with constraint on g^{CIB}	110.67	95.43	110.58	96.65
ILC (deproj tSZ)	19.43	25.17	19.43	25.17
ILC (deproj tSZ, no CIB)*	62.21	62.21	62.21	62.21
de-CIB (deproj tSZ)	23.22	25.52	29.25	25.93
ILC (deproj tSZ and CIB)	5.94	7.33	5.94	7.33
ILC with constraint on g^{CIB} (deproj tSZ)	6.36	8.47	6.34	8.43
ILC with constraints on g^{CIB} and g^{tSZ}	18.74	6.66	19.72	6.60

unWISE vs. Future LSS Surveys



Ratios

