A PERSONAL JOURNEY THROUGH HADRONIC EXOTICA

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EXOTICA

Exotica and **Erotica** differ only in one letter.

They are equally addictive.

Like all addictions, they have consequences.

- They consume a lot of the resources.
- They make you often do things you should not do.
- **BUT**, they are exciting, and they give you a great surge of adrenaline.

I have to confess that over the years I have fallen for exotica, and often. So, let me take you on a personal journey through exotica.

So, what is Exotic?

Exotic has to be unexpected. Exotic has to have the nature of the "forbidden fruit".

- Exotic in hadronic physics often begins with provocative suggestions by theorists, which drives experimentalists to search for it, often at **exotic cost** (think Higgs).
- At other times, it begins with an unexpected experimental observation for which theorists come up with exotic explanations (think J/ψ).

I want to tell the story of the hadronic exotica, necessarily from a personal point of view.

Chasing Exotica in Nuclear Physics

I began my career as a nuclear physicist. So, my first run in with exotica was in the search for **exotic nuclei**. Nuclei are exotic if they are very rich in neutrons, i.e., have an exceptionally large value of (N - Z)/A, or if they are just very heavy, $A \gg 240$.

- In the 1970's, there were no easy ways of making a nucleus which was very rich in neutrons, like ¹⁸C with 6 protons and 12 neutrons. And so we went for it by the very exotic pion double charge exchange (DCX) reaction (π⁺, π⁻).
- We discovered $^{18}\mathrm{C}$ by the reaction $^{18}\mathrm{O}(\pi^-,\pi^+)^{18}\mathrm{C}$ [1]. That was exciting.
- As I said before, exotica is addictive. So, after ¹⁸C we went for ⁹He, 2 protons+7 neutrons, (N – Z)/A = 5/9, by means of the reaction ⁹Be(π⁻, π⁺)⁹He [2]. We found it, and Bethe called it "a drop of neutron star". How much more exotic can you get?
- Well, how about ⁶H by ⁶Li(π⁻, π⁺)⁶H. We tried and failed to find it, bound or unbound [3].

So, running after exotica can lead to disappointments.



• The other end of exotic nuclei is the **superheavy nuclei**. I have never worked in this field. But Berkeley, Dubna, and GSI have crossed swords in claims about who has the heaviest of the superheavy. After some embarrassing incidents, the current winner is ²⁹⁴X₁₁₄ with 114 protons and 180 neutrons [4].

Chasing Exotica in Quark Physics

Quarks carry color, and **only color-neutral hadrons**, $q\bar{q}$ mesons or qqq, baryons exist in nature. In the quark bag model [5] hadrons with other color-neutral combinations, such as (qqq)(qqq) dibaryons, or $qq\bar{q}\bar{q}$ four-quark state can exist. de Swart and colleagues calculated the masses of scores of dibaryons [6] and started a stampede for the search of dibaryons.

- Lots of people started looking for dibaryons in their old experiments, analyzing old bubble chamber pictures and claiming observation of scores of dibaryons. As many as 40 dibaryon states were claimed in the mass range 1900 – 2300 MeV.
- We thought we could become famous by pinning these dibaryons down since we had orders of magnitude greater luminosity and energy resolution available at the Los Alamos Meson Factory. Instead of becoming **famous** for discovering dibaryons, we became **infamous** for killing all of them.





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Pentaquark

But that is not the end of this story. If not two baryons making a dibaryon, how about a baryon+a meson, or a color-neutral pentaquark? It surfaced a few years ago by the claim by Nakano et al. of a narrow peak, called Θ^+ , with a mass of $M(\Theta^+) = 1540 \pm 10$ MeV, $\Gamma(\Theta^+) < 25$ MeV, in the invariant mass of K^+n in the reaction $\gamma n \rightarrow K^-(K^+n)$ [7]. If true, it would have strangeness +1 and at least five quarks/antiquarks. The object was so exotic that a stampede of **confirming** claims flooded the literature. In a little more time, an equal number of non-observations were reported. If you go to Google, you find 99,800 entries for **Pentaquark**, and it will be difficult to decide whether the pentaquark is alive or not.

In a high-statistics repeat of their own measurement, JLab found that their own earlier observation of Θ^+ was false and no evidence for the existence of the pentaquark exists [8]. However, rumor has it that Nakano et al. claim that they still see the pentaquark in a high-statistics remeasurement.



So, once claimed, an exotic is difficult to kill!

I end with a quote from PDG08 summarizing the saga of the pentaquark

"The whole story — the discoveries themselves, the tidal wave of papers by theorists and phenomenologists that followed, and the eventual **'undiscovery'** — is a curious episode in the history of science."

The H Dibaryon

The uuddss H dibaryon was predicted by Jaffe [9], but it became so exotic that it was even considered a candidate for dark matter. Stubborn searches for the H were made for years at Brookhaven and KEK. The u, d quark dibaryons died a long time ago, but the H dibaryon lived longer. By now, however, by common consensus it is also considered dead. For a detailed history see [10].

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Exotica in QCD

In Dec. 1974, a large narrow peak was disocvered at ~ 3.1 GeV mass at Brookhaven and SLAC [11] in e^+e^- formation and $\mu^+\mu^-$ decay. It was the J/ψ which launched the era of modern Quantum Chromodynamics (QCD).

- It is amusing to note that barely four weeks later eight papers by theoretical physicists (including four Nobel laureates) appeared in the Jan. 6, 1974 issue of Phys. Rev. Letters [12], offering explanations of what J/ψ might be. Several of them were truely exotic explanations, like J/ψ was a bound state of a baryon/antibaryon, or two spin-one mesons, or it was a member of a 15 ⊕ 1 dimensional representation of SU(4).
- Tells you that nobody is immune to the lore of exotica.
- I have been talking too much about the exotics which failed to materialize. Let me now, for awhile, focus on exciting physics which is **not exotica, but exitica** (my construct for something very exciting).

QCD versus QED

The QCD potential which arises due to the exchange of a massless vector **photon** is $V(r) \propto \alpha_{em}/r$. The QCD potential due to the exchange of a massless vector **gluon** is $V(r) \propto \alpha_{\rm strong}/r$. Because free quarks do not exist, in QCD there is an additional confinement term proportional to r.

With such close analogy to QED, it is interesting to compare the QCD spectrum of charmonium with the QCD spectrum of positronium, with masses and interactions miles apart. It is nothing short of fantastic that **Nature repeats herself!** with energy scales different by a factor $\sim 10^{10}$.



Hyperfine Interaction in QCD

The Coulombic $(\propto 1/r)$ part of the QCD interaction gives rise to the usual spin dependence in the potential, with spin-orbit, tensor, and spin-spin components, in addition to the central part. Of these three, arguably the most important is the spin-spin interaction. For example, the ground state masses of $q\bar{q}$ mesons are:

$$M(q_1\bar{q}_2) = m_1(q_1) + m_2(q_2) + A_{hf} \left[\frac{\vec{s}_1 \cdot \vec{s}_2}{m_1 m_2}\right]$$

In order to determine the hyperfine interaction A_{hf} it is necessary to measure the **hyperfine splitting** between the spin-singlet and spin-triplet states it causes. This means identifying and measuring the masses of ${}^{3}L_{J}$ and ${}^{1}L_{J}$ states.

- The masses of **spin-triplet** ${}^{3}L_{J}$ states, ${}^{3}S_{1}$ and ${}^{3}P_{J}$ states are well-determined because either they are directly populated in $e^{+}e^{-}$ annihilation $(|{}^{3}S_{1}\rangle)$ or they are reached by strong E1 transitions from the $|{}^{3}S_{1}\rangle$ states $(|{}^{3}S_{1}\rangle \rightarrow \gamma_{E1} |{}^{3}P_{J}\rangle)$.
- The **spin-singlet** states ${}^{1}L_{J=L}$ can not be directly formed, and radiative transitions to them from spin-triplet states are either forbidden or weak M1.
- The net result is that our knowledge of the spin-singlet states, and therefore of the hyperfine interaction, has been very poor in the past. Very recently this has changed.

Hyperfine Interaction in QCD

For heavy quark systems, $c\bar{c}$ charmonium, and $b\bar{b}$ bottomonium, we would like to know how the hyperfine interaction changes as we move from the Coulomb dominated region of the $q\bar{q}$ potential to the confinement dominated region. We would like to study the change in the hyperfine interaction

- 1. between $c\bar{c}(1S)$ and $c\bar{c}(2S)$
- 2. between $c\bar{c}(1S)$ and $b\bar{b}(1S)$
- 3. between $c\bar{c}(1S)$ and $c\bar{c}(1P)$



• Until recently, the only hyperfine splitting known was

 $\Delta M_{hf}(1S) \equiv M(J/\psi(1S)) - M(\eta_c(1S)) = 116.7 \pm 1.2 \text{ MeV}$

• $\eta'_c(1^1S_0)$, $h_c(1^1P_1)$, and $\eta_b(1^1S_0)$ were not even identified. In the last five years, all this has changed due to the measurements at Belle, BaBar, and CLEO.



Back to Exotica



I now return to the domain of Exotica.

Recently, a number of new states have been claimed in the mass region 3800–4700 MeV, above the $D\overline{D}$ breakup of charmonium at 3730 MeV. It is usually called the "charmonium energy region".



- This narrow state with M(X) = 3872.2 ± 0.8 MeV, and Γ(X) = 3.0^{+1.9}_{-1.4} ± 0.9 MeV, has been observed by Belle, BaBar, CDF, DØ, and it definitely exists. [PDG08]
- CDF angular correlation studies show that its $J^{PC} = 1^{++}$ or 2^{-+} .
- X(3872) does not easily fit in the charmonium spectrum. Since its mass is very close to M(D) + M(D*), the most popular conjecture is that it is a weakly bound molecule of D and D*. If so, our recent precision measurement of D⁰ mass at CLEO gives M(D⁰D^{0*}) = 3871.81 ± 0.36 MeV. This corresponds to X(3872) being unbound by 0.4 ± 0.8 MeV. If X(3872) were bound by ~ 0.4 MeV, the branching fraction for the molecule's breakup into DDπ is predicted to be factor 400 smaller than observed. These two observations raise serious doubts about the molecular model for X(3872).
- To avoid the $D\overline{D}\pi$ problem it is speculated that there is another resonance nearby. There are no convincing observations of it so far. So what is X(3872) ?
- We need higher precision M(X) and $M(D^0)$, and $B(X \rightarrow D^0 D^0 \pi^0)$ to throw some fresh light on the nature of X(3872).



Y(4260)

 The Y(4260) has been observed in ISR production by BaBar, CLEO and Belle, and in direct production by CLEO. Y(4260) is clearly a vector with J^{PC} = 1⁻⁻, but a very strange one, since it sits at a very deep minimum in R, with

 $M(Y(4260)) = 4263^{+8}_{-9} \text{ MeV}, \quad \Gamma(Y(4260)) = 95 \pm 14 \text{ MeV}$ (PDG08)

So it is not likely to be a charmonium vector, which are all spoken for, anyway. So what is Y(4260)?

- It is suggested that Y(4260) is a ccg charmonium hybrid. If so, there ought to be 0⁻⁺ and 1⁻⁺ hybrids companions nearby. The exciting challenge for experimentalists is to find them.
- There are new problems. Belle has revived the question whether there is actually one resonance or two. Further, Belle reports that M(Y) in $Y \rightarrow J/\psi \pi \pi$ and $Y \rightarrow \psi' \pi \pi$ is different by almost 120 MeV.
- It is a real experimental challenge to clarify this situation before taking any theoretical conjecture seriously.



The Saga of X,Y,Z(\sim 3940)

- These three states, reported so far by Belle only, all have same masses within ± 7 MeV. All decay into states which contain a c and a \bar{c} quark; hence the designation charmonium-like. Each is produced in a different formation channel and each decays into a different decay channel. Even with e^+e^- luminosities of up to $\sim 700 \text{ fb}^{-1}$ thrown at them none has more than 75 counts in their favorite decay. If all that makes you slightly skeptical you are not alone. I summarize them in a table.
- The X(3943) is produced in e⁺e⁻ →double charmonium, and since only J = 0 states, η_c, χ_{c0}, and η'_c appear to be produced in the same spectrum, it is conjectured that its spin is also J = 0, and it is most likely η''_c(3¹S₀).
- The Z(3929) is produced in $\gamma\gamma$ fusion and decays to $D\overline{D}$. Its angular distribution suggests J = 2, and it is conjectured to be $\chi'_{c2}(2^3P_2)$.
- The Y(3943) is produced in B → KY and decays to ωJ/ψ. It is speculated that it might be a hybrid. It appears least convincing of the three.



Three Newer States from Belle

	Source	Mass (MeV)	Width (MeV)	Events	Reaction
Χ′	Belle	4160	$139(^{113}_{65})$	$24\binom{12}{8}$	$e^+e^- \rightarrow J/\psi + D^*D^*$
X″	BaBar	4324	172(33)	65(10)	$e^+e^- \rightarrow \psi(2S)\pi^+\pi^-$
	Belle	4360	74(18)	~ 50	$e^+e^- \rightarrow \psi(2S)\pi^+\pi^-$
X′′′′	Belle	4660	48(15)	~ 36	$e^+e^- \rightarrow \psi(2S)\pi^+\pi^-$



More From Belle

At this point you probably wish that these "discoveries" of exotic states by Belle would stop. But it is not to be. In the last six months, Belle has produced at least three more, and they are more exotic than all the previous ones. Because they are charged. They are:



Epilogue

The sum total of the experiences in this journey through hadronic exotica is that

- The journey is certainly worth it. It is unquestionably exciting.
- But the road is full of pitfalls and disappointments.

Only the brave should enter!

They should be proud of their successes, and humble enough to admit their failures.

- [1] H. Nann et al., PRL 41, 1589 (1978)
- [2] K. K. Seth et al., PRL 58, 1936 (1987)
- [3] B. Parker et al., Phys. Lett. B 251, 483 (1999)
- [4] Yu. Ts. Oganessian et al. PRC 74, 044602 (2006)
- [5] A. Chodos, R. L. Jaffe, et al. PRD 9, 3471 (1974)
- [6] P. J. Mulders, A. T. Aerts and J. J. de Swart, PRD 19, 2635 (1979); PRD 21, 1370 (1980); PRD 21, 2653 (1981)
- [7] T. Nakano et al. PRL 91, 012002 (2003)
- [8] B. McKinnon et al., PRL 96, 212001 (2006)
- [9] R. L. Jaffe, PRL 38, 195 (1977); PRL 38, 617E (1977)
- [10] A. L. Tralfner, Ph. D. dissertation, UC Berkeley, 2006; also PRC 77, 044908 (2008)
- [11] J. J. Aubert et al., PRL 33, 1404 (1974); J. E. Augustin et al. PRL 33, 1406 (1974)
- [12] Various authors, PRL 34, 36-56 (1975)
- [13] D. M. Asner et al., PRL 92, 142001 (2004)
- [14] J. L. Rosner et al., PRL 95, 102003 (2005); S. Dobbs et al., PRL (2008, accepted)
- [15] B. Aubert et al., PRL 101, 071801 (2008)