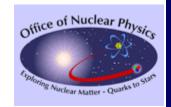


Poincaré Invariant Three-Body Scattering

Ch. Elster

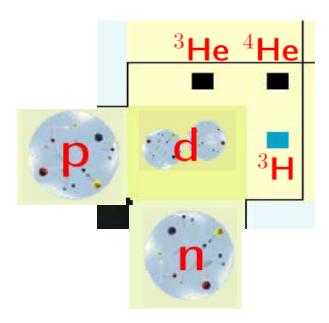
T. Lin

W. Polyzou, W. Glöckle



A Few-Body Theorist's view of the Nuclear Chart



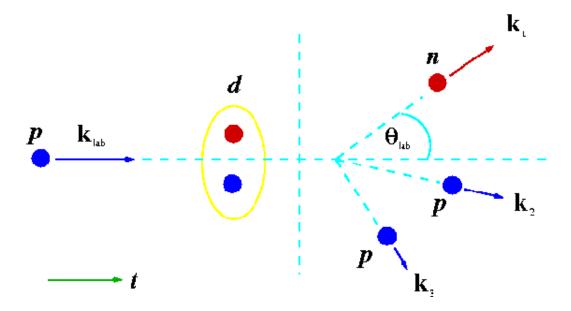


3 Nucleons

•Bound State: ³H - ³He

•Scattering: Elastic – Inelastic (Breakup)

•Energy Scale: $keV \rightarrow MeV \rightarrow GeV$



Challenges in 3N Physics

- Test of nuclear forces in the simplest nuclear environment (over a large energy range!)
 - Two-body forces
 - Genuine three-body forces
- Reaction mechanisms
 - Example: deuteron breakup, (p,n) charge exchange
 - Higher Energy: Lorentz vs. Galilean Invariance
 - Check of commonly used approximations (Glauber)
 - Three-body decays of particles (e.g. η or η ')

Relativistic Effects at Higher Energies Computational Challenge:

3N and 4N systems:

- standard treatment based on pw projected momentum space successful (3N scattering up to ≈250 MeV) but rather tedious
- 2N: j_{max} =5, 3N: J_{max} =25/2 \rightarrow 200 `channels'
- Computational maximum today:
- $2N: j_{max} = 7$, $3N: J_{max} = 31/2$

⇒ Solution:

⇒ NO partial wave decomposition of basis states

Roadmap for 3N problem without PW Scalar NN model | Realistic NN Model

- NN scattering + bound state
- 3N bound state
- 3N bound state + 3NF
- 3N scattering:
- Full Faddeev Calculation



- Elastic scattering
- Below and above break-up
- Break-up

Poincarė Invariant (Faddeev Calculations

- NN scattering + deuteron
 - Potentials AV18 and Bonn-B
- Break-up in first order:
 - (p,n) charge exchange



Lorentz kinematics



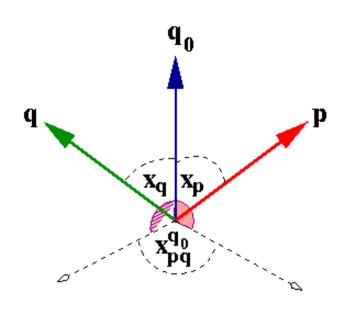
NN interactions

High energy limits



Variables for 3D Scattering Calculation

3 distinct vectors in the problem: $\mathbf{q_0}$ \mathbf{q} \mathbf{p}



5 independent variables:

$$p = |\mathbf{p}|, \ q = |\mathbf{q}|$$

$$x_p = \hat{\mathbf{p}} \cdot \hat{\mathbf{q}}_0, \ x_q = \hat{\mathbf{q}} \cdot \hat{\mathbf{q}}_0$$

$$x_{pq}^{q_0} = (\mathbf{q}_0 \times \mathbf{q}) \cdot (\mathbf{q}_0 \times \mathbf{p})$$

q system: z || q

 \mathbf{q}_0 system : $\mathbf{z} \parallel \mathbf{q}_0$

Numerical calculation:

Three-body transition amplitude is a function of 5 variables

Relativistic Three-Body Problem

Context: Poincarė Invariant Quantum Mechanics

- Poincarė invariance is exact symmetry, realized by a unitary representation of the Poincarė group on a fewparticle Hilbert space
- Instant form
- Faddeev equations same operator form but different ingredients

Kinematics

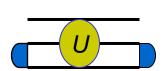
Lorentz transformations between frames

Dynamics

- Bakamjian-Thomas Scheme: Mass Operator M=M₀+V replaces Hamiltonian H=H₀+v
- Connect Galilean two-body v with Poincarė two-body v
- Construct V := $\sqrt{M^2 + q^2} \sqrt{M_0^2 + q^2}$

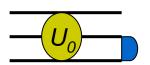
Three-Body Scattering

 Transition operator for elastic scattering $U = PG_0^{-1} + PT$



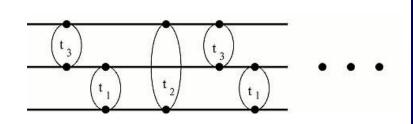
Transition operator for breakup scattering

$$U_0 = (1 + P)T$$



Faddeev equation (Multiple Scattering Series)

$$T = tP + tG_0 PtP + \cdots$$
1st Order in tP



$$t = v + vg_0t =: NN t$$
-matrix

$$t = v + vg_0t =: NN t$$
-matrix $P = P_{12} P_{23} + P_{13} P_{23} \equiv Permutation Operator$

Kinematic Relativistic Ingredients:

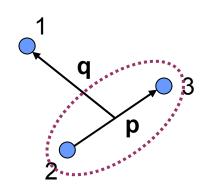
- Lorentz transformation Lab → c.m. frame (3-body)
- Phase space factors in cross sections
- Poincarė-Jacobi momenta
- Permutations for identical particles

Kinematics: Poincaré-Jacobi momenta

Nonrelativistic (Galilei)

$$\mathbf{p} = \frac{1}{2}(\mathbf{k}_2 - \mathbf{k}_3)$$

$$\mathbf{q} = \frac{2}{3}(\mathbf{k}_1 - \frac{1}{2}(\mathbf{k}_2 + \mathbf{k}_3))$$



Relativistic (Lorentz)

$$\mathbf{p} = \frac{1}{2}(\mathbf{k}_{2} - \mathbf{k}_{3}) + \frac{\mathbf{k}_{2} + \mathbf{k}_{3}}{2m_{23}} \left(\frac{(\mathbf{k}_{2} - \mathbf{k}_{3}) \cdot (\mathbf{k}_{2} + \mathbf{k}_{3})}{(E_{2} + E_{3}) + m_{23}} - (E_{2} - E_{3}) \right)$$

$$\mathbf{q} = \mathbf{k}_{1} + \frac{\mathbf{K}}{M} \left(\frac{\mathbf{k}_{1} \cdot \mathbf{K}}{E + M} - E_{1} \right)$$

$$\begin{vmatrix} \mathbf{k}_1 \mathbf{k}_2 \mathbf{k}_3 \rangle = \begin{vmatrix} \frac{\partial (Kpq)}{\partial (\mathbf{k}_2 \mathbf{k}_3)} \end{vmatrix}^{1/2} |Kpq\rangle \neq 0$$

$$K = \mathbf{k}_1 + \mathbf{k}_2 + \mathbf{k}_3$$

$$M = \sqrt{E^2 - \mathbf{K}^2}$$

$$m_{23} = \sqrt{(E_2 + E_3)}$$

$$E = E_1 + E_2 + E_3$$

$$\mathbf{K} = \mathbf{k}_1 + \mathbf{k}_2 + \mathbf{k}_3$$

$$M = \sqrt{E^2 - \mathbf{K}^2}$$

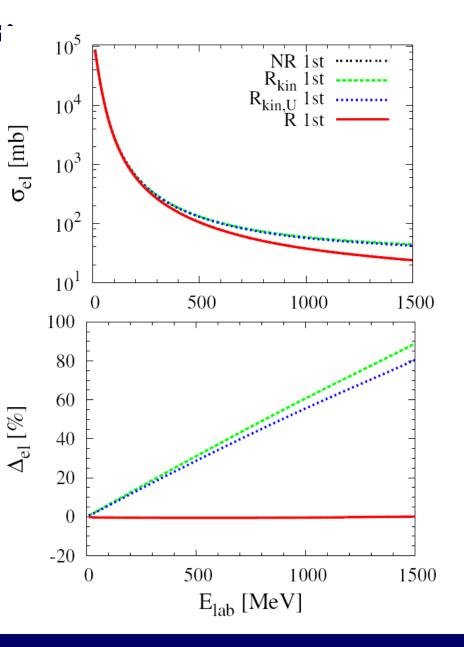
$$m_{23} = \sqrt{(E_2 + E_3)^2 - (\mathbf{k}_2 + \mathbf{k}_3)^2}$$

Relativistic kinematics

IA (1st order)

$$T = tP$$

- Lorentz transformation
 Lab → c.m. frame) (3-body)
- Phase space factors in cross sections
- Poincarė-Jacobi momenta
- Permutations



Quantum Mechanics

Galilei Invariant:
$$H = \frac{K^2}{2M_g} + h$$
 ; $h = h_0 + v_{12}^{NR} + v_{13}^{NR} + v_{23}^{NR}$

Poincaré Invariant:
$$H = \sqrt{K^2 + M^2}$$
 ; $M = M_0 + V_{12} + V_{23} + V_{31}$

$$V_{ij} = M_{ij} - M_0 = \sqrt{(m_{0,ij} + v_{ij})^2 + q_k^2} - \sqrt{m_{0,ij}^2 + q_k^2}$$

Two-body interaction embedded in the 3-particle Hilbert space

$$m_{0,ij} = \sqrt{m_i^2 + p_{ij}^2} + \sqrt{m_j^2 + p_{ij}^2}$$

$$M_0 = \sqrt{m_{0,ij}^2 + q_k^2} + \sqrt{m_k^2 + q_k^2}$$

V_{ij} embedded in the 3-particle Hilbert space

$$V_{ij} = M_{ij} - M_0 = \sqrt{(m_{0,ij} + v_{ij})^2 + q_k^2} - \sqrt{m_{0,ij}^2 + q_k^2}$$

need matrix elements: $\langle \vec{k} | V(\vec{p}) | \vec{k}' \rangle$

$$\begin{split} &= v(\vec{k}, \vec{k}') + \psi_b(\vec{k}) (\sqrt{M_b^2 + p^2} - M_b) \psi_b(\vec{k}') + \frac{1}{\omega - \omega'} \bigg[(\sqrt{\omega^2 + p^2} - \omega) \Re[t(\vec{k}', \vec{k}; \omega)] \\ &- (\sqrt{\omega'^2 + p^2} - \omega') \Re[t(\vec{k}, \vec{k}'; \omega')] \bigg] + \frac{1}{\omega - \omega'} \bigg[\mathcal{P} \int d^3k'' \frac{(\sqrt{\omega''^2 + p^2} - \omega'')}{\omega'' - \omega} t(\vec{k}, \vec{k}''; \omega'') t^*(\vec{k}', \vec{k}''; \omega'') \\ &- \mathcal{P} \int d^3k'' \frac{(\sqrt{\omega''^2 + p^2} - \omega'')}{\omega'' - \omega'} t(\vec{k}, \vec{k}''; \omega'') t^*(\vec{k}', \vec{k}''; \omega'') \bigg]. \end{split}$$

H. Kamada,^{1,*} W. Glöckle,^{2,†} J. Golak,^{2,3,‡} and Ch. Elster^{4,§} PHYSICAL REVIEW C **66**, 044010 (2002)

Two-Body Input: T1-operator embedded in 3-body system

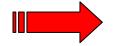
$$T_{1}(p', p; q) = V(p', p; q) + \int d^{3}k'' \frac{V(p', k''; q) T_{1}(k'', p; q)}{\sqrt{(2E(p'))^{2} + q^{2}} - \sqrt{(2E(k''))^{2} + q^{2}} + i\varepsilon}$$
The property of th

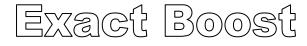
 Obtain fully off-shell matrix elements T₁(k,k',q) from half shell transition matrix elements by

Solving a 1st resolvent type equation:

$$T_1(q) = T_1(q') + T_1(q) [g_0(q) - g_0(q')] T_1(q')$$

- For every single off-shell momentum point
- Proposed in
 - Keister & Polyzou, PRC 73, 014005 (2006)
- Carried out for the first time here [PRC 76, 1014010 (2007)]







Obtain embedded 2N t-matrix $T_1(k,k',z')$ half-shell in 2-body c.m. frame first:

$$\langle \mathbf{k} | T_1(\mathbf{q}; z') | \mathbf{k}' \rangle = \langle \mathbf{k} | V(\mathbf{q}) | \mathbf{p}'^{(-)} \rangle$$

$$= \frac{2(E_{k'} + E_k)}{\sqrt{4E_{k'}^2 + \mathbf{q}^2} + \sqrt{4E_k^2 + \mathbf{q}^2}} t(\mathbf{k}, \mathbf{k}'; 2E_{k'})$$

$$t(\mathbf{k}, \mathbf{k}'; 2E_{k'}) = v(\mathbf{k}, \mathbf{k}') + \int d\mathbf{k}'' \frac{v(\mathbf{k}, \mathbf{k}'')t(\mathbf{k}'', \mathbf{k}'; 2E_{k'})}{E_{k'} - 2\sqrt{m^2 + k''^2} + i\epsilon}$$

Solution of the relativistic 2N LS equation with 2-body potential

Consideration for two-body t-matrix

- Relativistic and non-relativistic t-matrix should give identical observables for determining relativistic effects
- Or two-body t-matrices should be phase-shift equivalent
- Four options:
 - Start from relativistic LS equation
 - natural option employed for NN interactions fit to 1 GeV +
 - If non-relativistic LS equation is used:
 - Refit of parameters (maybe time consuming in practice)
 - Transformation of Kamada-Glöckle PRL 80, 2547 (1998)
 - Transformation of Coester-Piper-Serduke as given in Polyzou PRC 58, 91 (1998)

Phase equivalent 2-body t-matrices: Coester-Pieper-Serduke (CPS) (PRC11, 1 (1975))

Add interaction to square of non-interacting mass operator

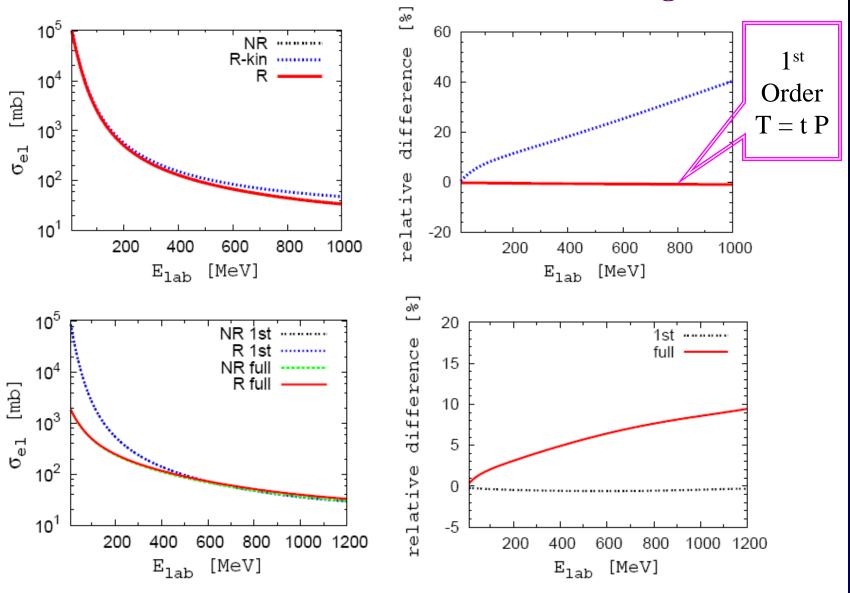
$$M^{2} = M_{0}^{2} + u = 4mh$$
 with $h = \frac{k^{2}}{m} + \frac{u}{4m} + m$
 $u = v^{2} + \{M_{0}^{2}, v\}$

- NO need to evaluate v directly, since M, M², h have the same eigenstates
- Relation between half-shell t-matrices

$$\langle k' | t_R(e(k)) | k \rangle = \frac{4m}{e(k) + e(k')} \langle k' | t_{NR}(k^2/m) | k \rangle$$

 Relativistic and nonrelativistic cross sections are identical functions of the invariant momentum k

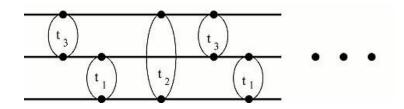
Total Cross Section for Elastic Scattering:



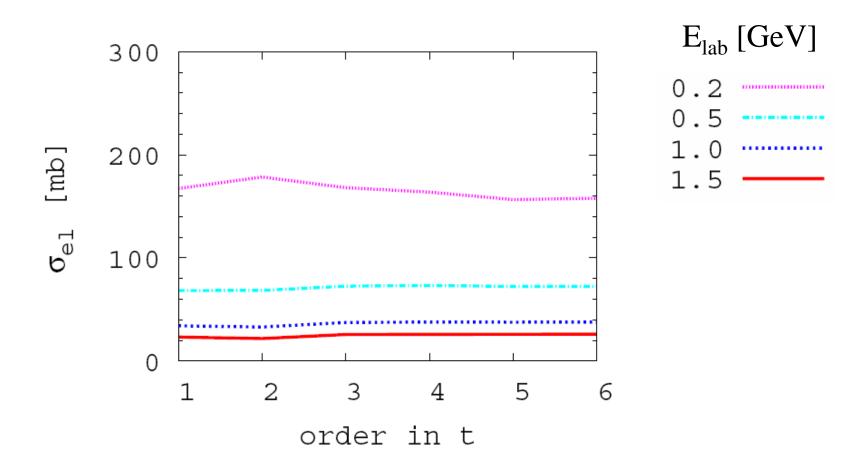
Faddeev Equation as Multiple Scattering Series

$$T = tP + tG_0PT$$

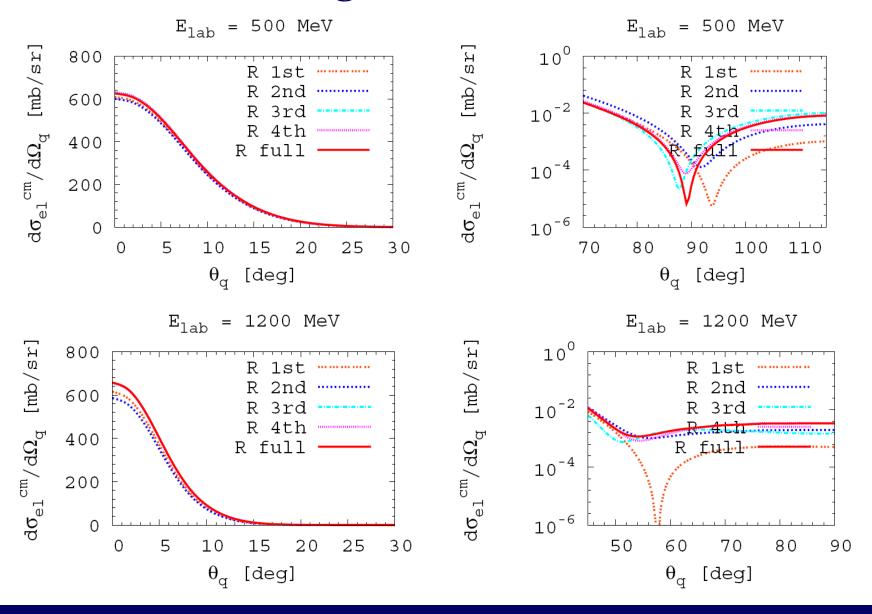
$$T = tP + tG_0PtP + \cdots$$
1st Order or IA

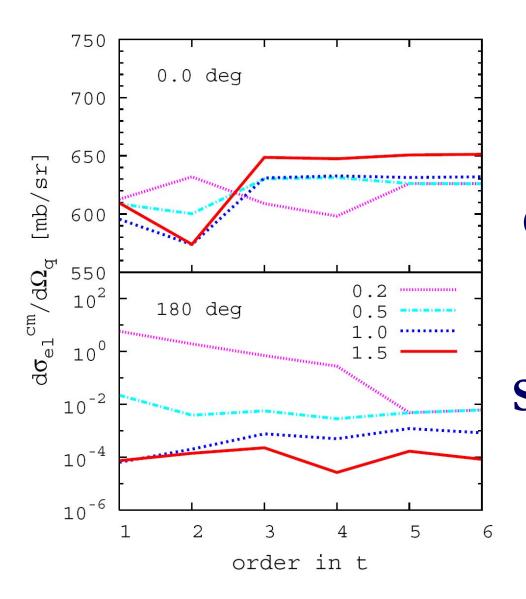


Convergence of the Faddeev Multiple Scattering Series



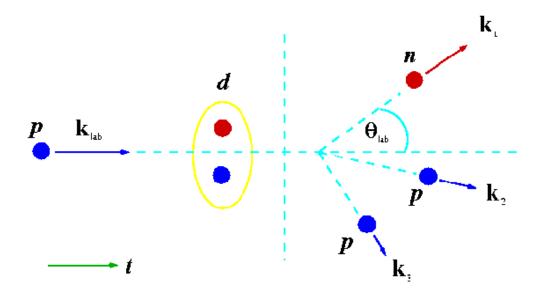
Elastic Scattering: Differential Cross Section





Differential
Cross Section:
Convergence of
the Faddeev
Multiple
Scattering Series

Breakup Scattering



Exclusive: Measure energy & angles of two ejected particles

V.Punjabi et al. PRC 38, 2728 (1998) – TRIUMF p+d @ 508 MeV

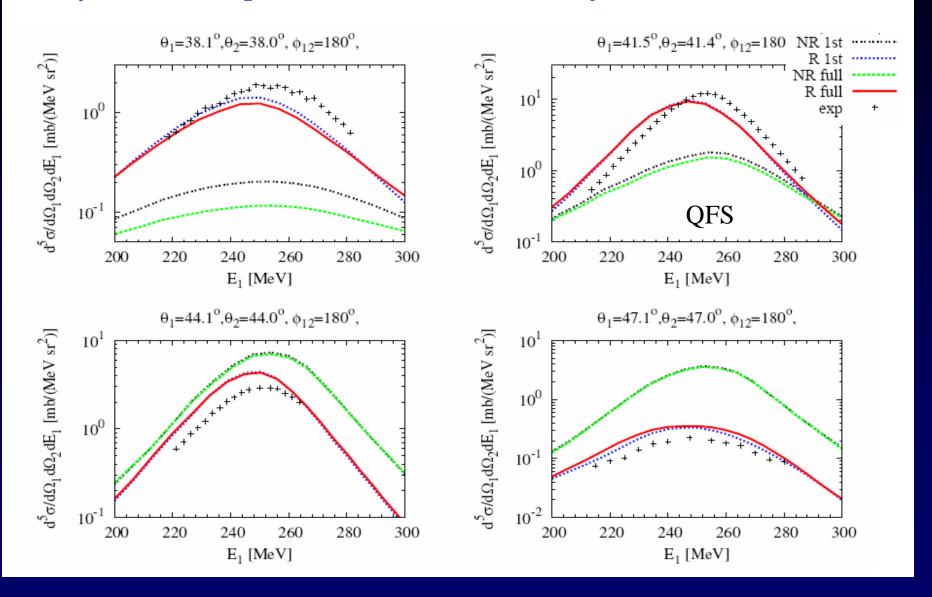
Outgoing protons are measured in the scattering plane

Exclusive Breakup Scattering

 $E_{lab} = 508 \text{ MeV}$

(symmetric configuration)

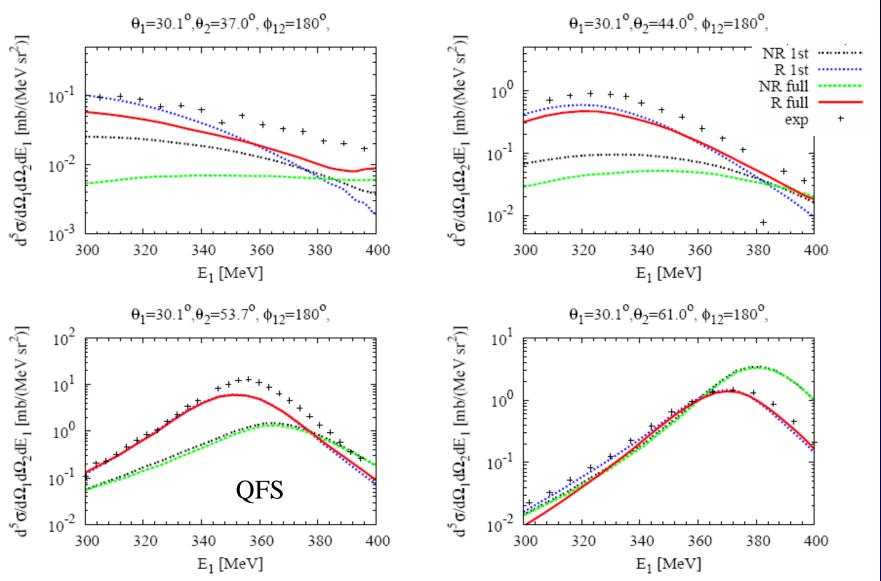
(V.Punjabi et al. PRC 38, 2728 (1998)



Exclusive Breakup Scattering

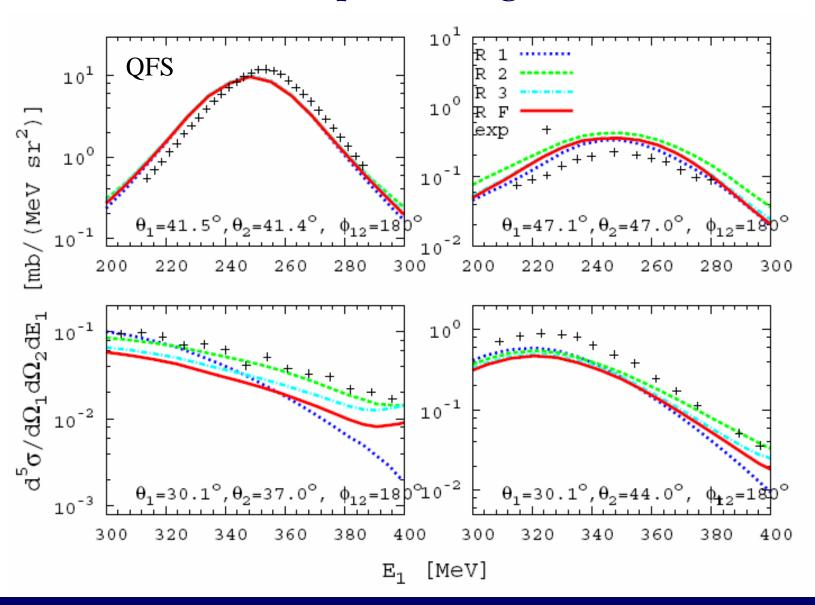
 $E_{lab} = 508 \text{ MeV}$

(asymmetric configuration)

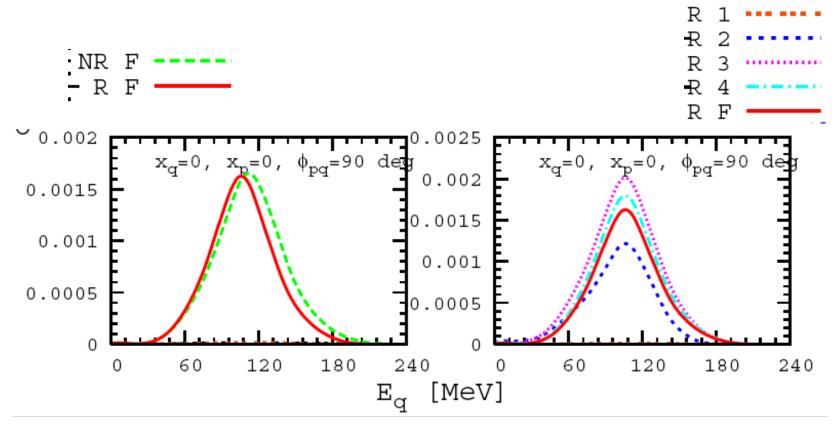


Exclusive Breakup Scattering

 $E_{lab} = 508 \text{ MeV}$



Exclusive Breakup Scattering Space-Star



 $E_{lab} = 508 \text{ MeV}$

Poincaré Invariant Faddeev Calculations

Kinematics

- Phase space factors
- Lorentz Transformation from Lab to c.m. frame
- Lorentz Transformation of Jacobi Coordinates
 - Always reduces effects of phase-space factors
- Kinematics determines peak positions in break-up observables

Dynamics

- Exact calculation of the two-body interaction embedded in the three-particle Hilbert space
- The dynamic effects act in general opposite kinematic effects

Poincaré Invariant Faddeev Calculations

- Carried out up to 2 GeV for elastic and breakup scattering
 - Solved Faddeev equation in vector variables = NO partial waves
- Relativistic effects are important at 500 MeV and higher
 - Relativistic total elastic cross section increases up to 10% compared to the non-relativistic
 - Relativistic kinematics determines QFS peak positions in inclusive and exclusive breakup
 - Breakup: Relativistic effects very large dependent on configuration
- Above 800 MeV projectile energy:
 - multiple scattering series converges after ~2 iterations
- Future
 - Systematic studies of selected cross sections & high energy limits
 - Long term: include Spin