Cold nuclear effects on azimuthal decorrelation in Heavy-Ion collision

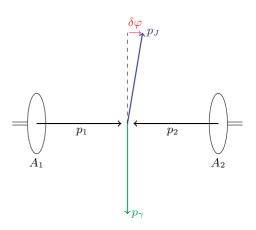
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Initial setup

Regime of interest

- Large nuclei: $A_1, A_2 \gg 1$.
- Mid-rapidity back to back pair $\{ {\sf Jet} \gamma \}$, or DY-pair. Mostly transverse momenta $p_J \sim p_\gamma \sim Q$
- **Observable**: azimuthal decorrelation (\Leftrightarrow Momentum imbalance): $\delta \varphi$
- Scaling $\delta \varphi \sim Q^y/Q \ll 1$, with Q^y semihard scale $Q^y \gg \Lambda_{QCD}$



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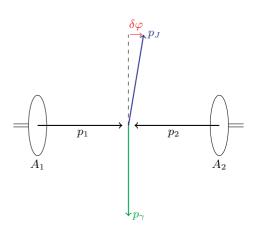
Initial setup

Goals:

- Prove to the best of our knowledge from first principle factorization for DY and jet- γ in AA collisions.
- Compute the leading order correction in α_S to the azimuthal decorrelation due to cold nuclear effects.
- Computed using single scattering, further resumed

Result:

 Factorization of the hard process with the medium-modified initial state parton distribution and medium-modified jet function



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Outline

Part 1: Baseline

- Setup
- Localization of the hard process

Part 2: Medium implementation

• Glauber gluons and implications

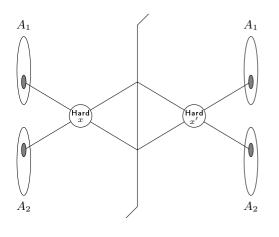
Part 3: Recovering a factorization formula

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- Initial state effect
- Final state effect
- Interference terms

Part 1: Baseline

Baseline: Factorization of the bare process



Setup:

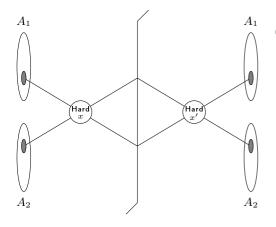
- Working in coordinate space.
- Describe colliding nuclei states by introducing the Wigner function. For the plus moving A_1 , and $q^+ \ll P^+$, we have

$$W_{A_1}(P,b) = \int \frac{dq^+ d^2\underline{q}}{(2\pi)^3 2P^+} e^{-\frac{i}{2}q^+b^- + i\underline{q}\cdot\underline{b}}$$
$$\times \langle P + q/2|A_1\rangle \langle A_1|P - q/2\rangle$$
$$= \hat{\rho}_{A_1}(b^-,\underline{b})2(2\pi)^3 \delta^3(P - P_1)$$

 \Rightarrow Introduce the color charge density $\hat{\rho}$.

 This is the McLerran-Venugopalan model. Nucleons are assumed uncorrelated, and partons within are recoilless classical sources of the gluon field.

Baseline: Factorization of the bare process



Observable:

- Hard process: inner propagators far off-shell
 - ⇒ Those propagators shrink to points (wrt to any other separation involved)
- ullet Sum over all possible nucleon yields the impact parameter dependent cross section for the observable ${\cal O}$
- Involve the parton distributions of nucleons within the nuclei, and the partonic cross section $\hat{\sigma}_{ij \to k\ell}$

$$\frac{d\sigma^{(0)}}{d^2\underline{b}d\mathcal{O}} = 2\int d^4X \rho_{A_1}(X^-, \underline{X})\rho_{A_2}(X^+, \underline{X} - \underline{b}) \sum_{ij} \int d\xi_1 d\xi_2 f_i(\xi_1) f_j(\xi_2) \frac{d\hat{\sigma}_{ij \to k\ell}}{d\mathcal{O}}$$
(1)

using $X = \frac{1}{2}(x + x')$

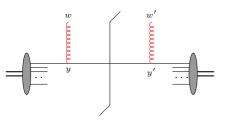
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Part 2: Medium implementation

Medium Implementation: Glauber gluons

Add and extra scattering center / nucleon

Real diagram



- ullet Background field A^μ obtained following the MV model
- ullet Sources: fast parton moving in the direction of either A_1 or A_2 . Focus only on A_2 in this presentation
- Expectation value of the 2 points correlator $\langle A(w)A(w') \rangle$ in coordinate space, over all possible nucleon configurations
- Glauber gluon: propagator dominated by the transverse momenta $k^2 \longrightarrow -k^2$

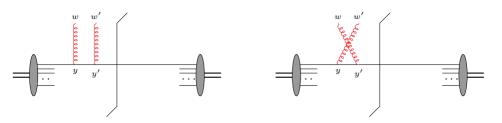
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Features of interest:

- Integrate out the final state parton after emission of the red gluons.
 - \longrightarrow Sets the time at y to be the same as y'
- Glauber gluon are instantaneous:
 - \longrightarrow Sets the time at w to be the same as $w' \Longrightarrow \mathsf{Locality} \ \mathsf{w.r.t.}$ time
- Remains the dependence on the transverse separations. \implies Encoded by model dependent (dim-reg / IR cutoff / dipoles, ...) functions, which we denote by $F_1(\underline{w}-y)F_1(\underline{w}'-y')$.

Medium Implementation: Glauber gluons

Virtual diagrams



Done many times before, some key points:

- See BDMPS-Z / GLV / ...
- Result similar to the real diagram
- Only two changes: (1) Relative sign wrt the real diagram, and (2) the transverse phase changes

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Part 3: Recovering a factorization formula

Recovering a factorization formula: basic tools

Feynman propagator in coordinate space at high energy

$$D_F(x) \longrightarrow \int \frac{dp^{n_j}}{2\pi} \frac{1}{2p^{n_j}} e^{-\frac{i}{n_i \cdot n_j} p^{n_j} x^{n_i}} \theta(x^{n_j}) \times \frac{n_i \cdot n_j}{4\pi i \ x^{n_j}} exp\left(\frac{i}{2} \frac{n_i \cdot n_j}{2} \frac{p^{n_j}}{x^{n_j}} \underline{x}_{ij}^2\right)$$
(2)

using
$$g^{\mu\nu}=rac{n_i^\mu n_j^\nu + n_j^\mu n_i^\nu}{n_i \cdot n_j} - g_{\perp ij}$$
 and $v^{n_i}\equiv v \cdot n_i.$

Remarks:

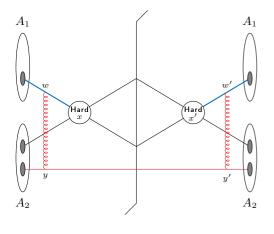
- For usual plus-minus movers: usual light cone coordinates.
- ullet Starts being interesting when the mover are not parallel. Think of an initial parton in the e_z direction and a produced parton moving in the e_x direction.

The transverse part is wrt the two directions n_i and n_j .

- In the high energy limit: Gaussian in transverse space becomes a Dirac delta.
- \bullet x^{n_j} plays the role of time. Notice: $\theta(x^{n_j})$

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Recovering a factorization formula: Initial state



Real diagram

- ullet Need to integrate over w and w' positions.
- Use the time ordering from the Feynman propagator:

$$(X-w)^+ > 0$$
 and $(X-w')^+ > 0$

• Average the bg fields $\langle A(w)A(w')\rangle$, use the locality in time of the 2pts function.

Iterating

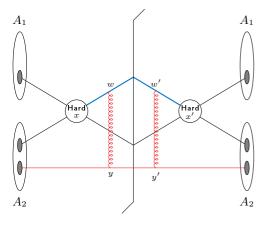
 Recurssive use of the locality of the 2pts function and the time ordering in the Feynman propagator (eikonal limit).

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• Building up a Wilson line along the plus-direction.

Alternatively Recast into a medium-modified initial distribution

Recovering a factorization formula: Final state



Metric

- Use $n_i = n_3$, in the plane transverse to the beam direction.
- ullet For the interaction with the minus moving nucleon, use $n_j=n_1$
- Feynman propagator is proportional to $\theta(x^+)$. With this choice, we use the same time as the previous case.

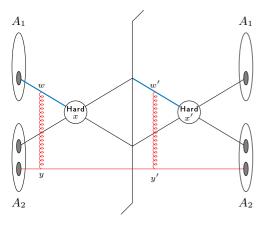
Real diagram

- \bullet Need to integrate over w and w^\prime positions.
- From Feynman propagators: $w^+ X^+ > 0$ and $w'^+ X^+ > 0$
- Use the properties of the bg field average, and iterate to build-up a Wilson line in the n_3 direction.

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Alternatively Recast into a medium-modified jet function:

Recovering a factorization formula: Interference



Short answer

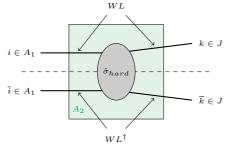
- Using the same metric, $n_i=n_3$ and $n_j=n_1$, same time v^+
- ullet Integrate over w and w'
- From the propagators $D_F(X-w)$, \Rightarrow we have $X^+ > w^+$
- From The propagator $D_F(w'-X)$, \Rightarrow we have $w'^+ > X^+$
- Perform the average over the bg fields. It sets $w^+ = w'^+ \longrightarrow {\sf No}$ support

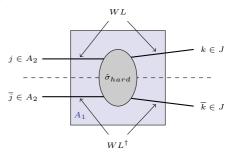
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Due to the lack of support, interference diagrams vanish

Summary and Prospects

Summary: consider $A_1A_2 \rightarrow J + \gamma$. Schematic result





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Alternatively

- ullet Initial distribution of parton $i\in A_1$ modified by the cold medium A_2
- ullet Initial distribution of parton $j\in A_2$ modified by the cold medium A_1
- ullet Jet function of parton k is modified by cold medium A_1 and cold medium A_2

Summary and Prospects

One very last equation:

$$\frac{d\sigma_{A_{1}A_{2}\to J\gamma}^{(1)}}{d^{2}\underline{b}d\eta_{3}d\eta_{4}dp_{T}dQ_{T}^{y}} = \frac{-1}{2} \int d^{4}X \rho_{A_{1}}(X^{-},\underline{X}) \rho_{A_{2}}(X^{+},\underline{X}-\underline{b}) \int \frac{dy}{2\pi} e^{-iyQ_{T}^{y}} y^{2}$$

$$\times \sum_{ijk} \left[\int_{X^{+}}^{X^{+}} dW^{+} \frac{d\sigma_{ij\to k\gamma}^{(0)}}{d\eta_{3}d\eta_{4}dp_{T}} \bigg|_{X^{+}-W^{+}} \hat{q}_{i/A_{2}}(W^{+},\underline{X}-\underline{b},|y|) \right]$$

$$+ \int_{X^{-}}^{X^{-}} dW^{-} \frac{d\sigma_{ij\to k\gamma}^{(0)}}{d\eta_{3}d\eta_{4}dp_{T}} \bigg|_{X^{-}-W^{-}} \hat{q}_{j/A_{1}}(W^{+},\underline{X},|y|)'$$

$$+ \int_{X^{+}} dW^{+} \frac{d\sigma_{ij\to k\gamma}^{(0)}}{d\eta_{3}d\eta_{4}dp_{T}} \bigg|_{W^{+}-X^{+}} \hat{q}_{k/A_{2}}(W^{+},\underline{X}\frac{W^{+}-X^{+}}{n_{3}^{+}},|y|)$$

$$+ \int_{X^{-}} dW^{-} \frac{d\sigma_{ij\to k\gamma}^{(0)}}{d\eta_{3}d\eta_{4}dp_{T}} \bigg|_{W^{+}-X^{+}} \hat{q}_{k/A_{1}}(W^{-},\underline{X}\frac{W^{-}-X^{-}}{n_{3}^{-}},|y|) \bigg|$$

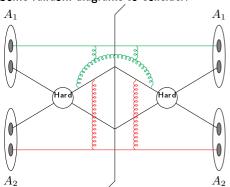
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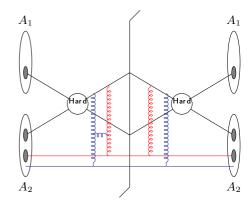
Summary and Prospects

Prospects: - higher order in α_S (add more gluons for more fun!)

- Introduce radiation. Where is the Sudakov? What about the medium modification?
- Finding small-x evolution?
- Can we distinguish? Is there overlap?

Some random diagrams to consider:





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