



GRENOBLE | MODANE



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**euclid**

# Correcting galaxy systematics using self-organizing maps

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Gaël Alguero

ML coffee, March 29, 2024

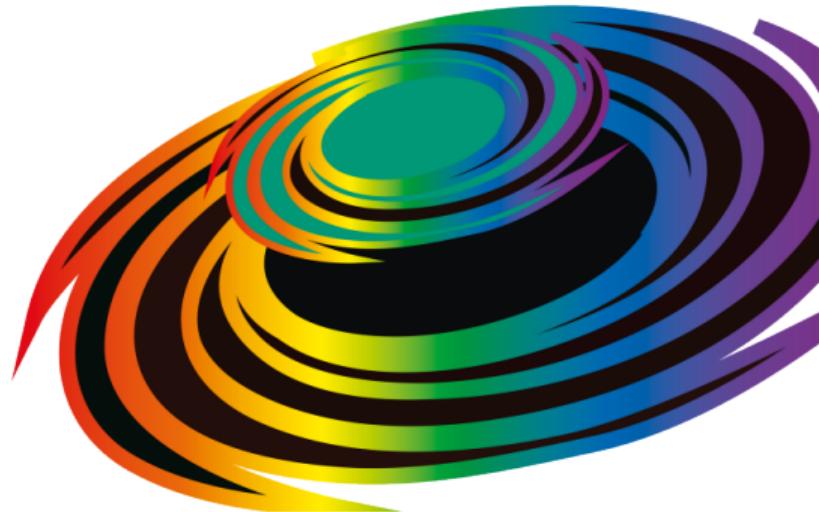
LPSC Grenoble

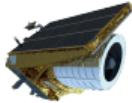
For the VMPZ-ID team: Jérôme Odier and Juan F. Macías-Pérez

Introduction systematic decontamination

Decontaminating DES systematics

Scientific validation



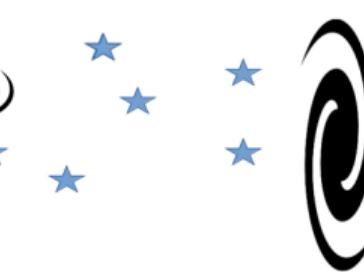


# Systematics on galaxy density

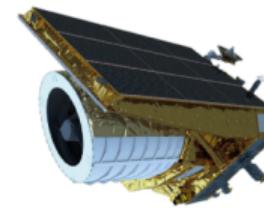
- Spatial variations of survey and sky properties as well as of instrumental performance will imprint on the measured galaxy density



Real galaxy distribution



Astrophysical contamination  
and extinction

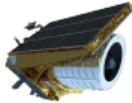


Observational conditions and  
instruments performances



Distorted galaxy distribution

- ⇒ Can affect galaxy clustering and the detection of clusters of galaxies
- ⇒ Need to correct this distortion



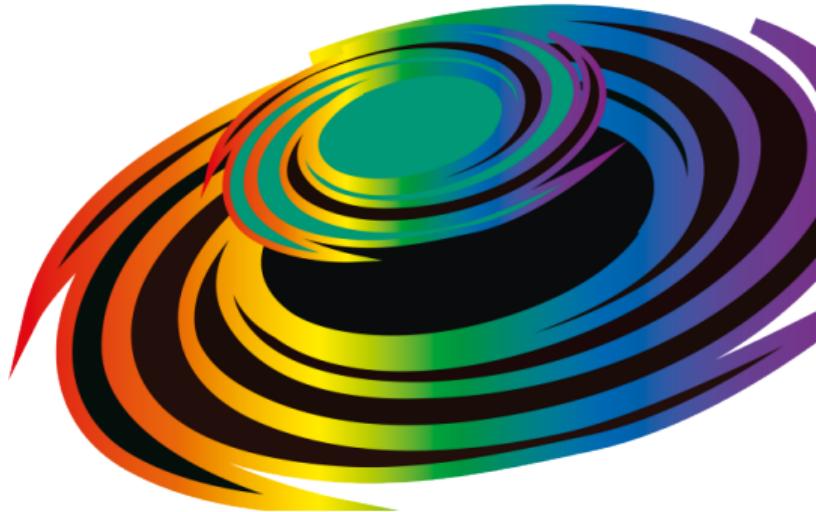
# Decontamination methods

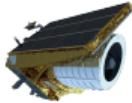
- Linear multiplicative methods in DES
  - Iterative Systematics Decontamination (ISD): Elvin-Poole et al. [[arXiv:1708.01536](#)], Rodriguez-Monroy et al. [[arXiv:2105.13540](#)]
  - Elastic Net Regularisation (ENet): Weaverdyck, Huterer [[arXiv:2007.14499](#)]
- Linear multiplicative and additive models
  - Theoretical framework: Weaverdyck, Huterer [[arXiv:2007.14499](#)]
  - Other non-linear methods
- Non-linear method in KIDS
  - Organised randoms from Self-Organizing Maps (SOM): Johnston et al. [[arXiv:2012.08467](#)]
- Other ideas?

Introduction systematic decontamination

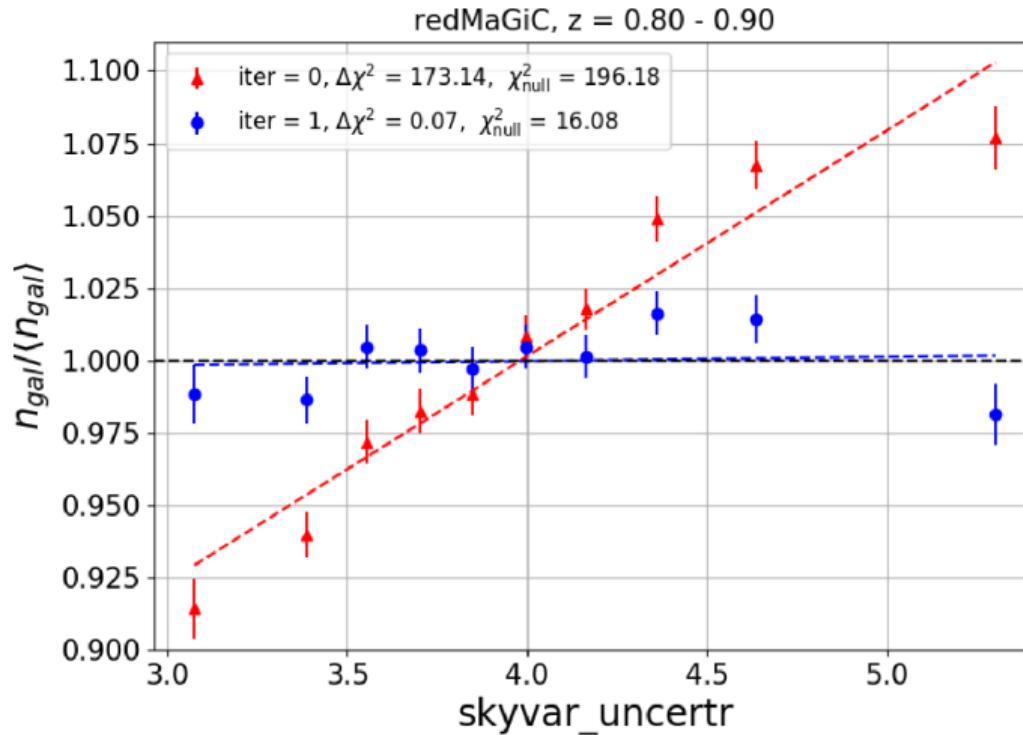
Decontaminating DES systematics

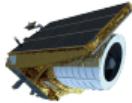
Scientific validation





# Iterative systematic decontamination



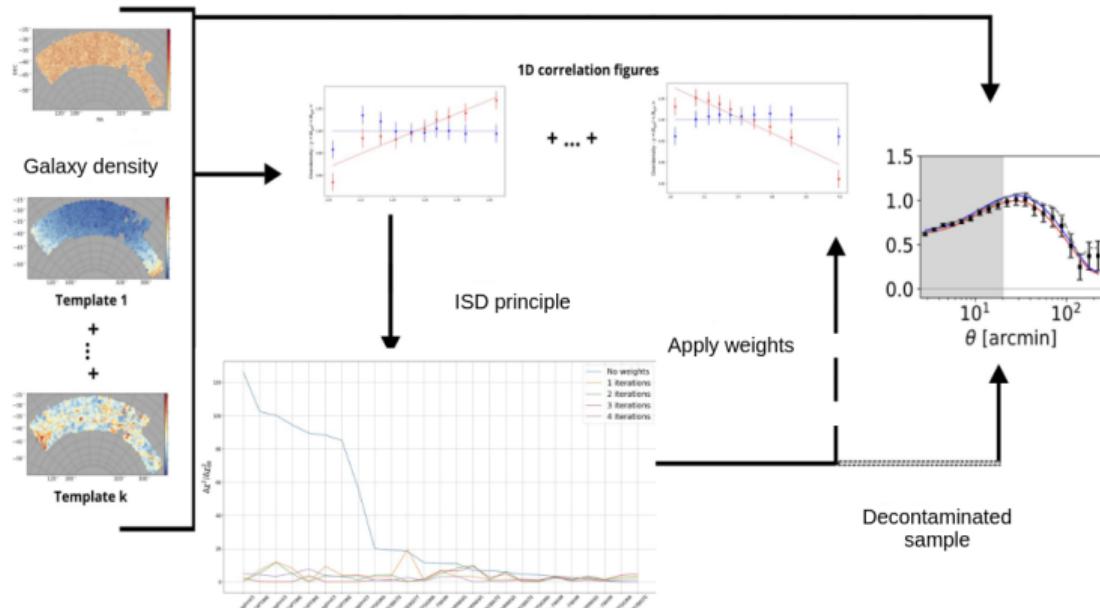


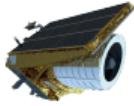
# Iterative systematic decontamination

- Assuming multiplicative bias
- Compute linear weights for correction

$$n^{\text{true}}(\theta, \phi) = \frac{n^{\text{obs}}(\theta, \phi)}{ax(\theta, \phi) + b}$$

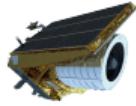
- Caveats
  - Needs fine-tuning
  - No correlations between systematics
  - Only linear effects





# Example dataset

	●	●	●	●	●	●	●	●	●	●	●	●	●	●	●	
R	0	255	255	0	0	255	0	255	128	0	128	0	255	255	255	240
G	0	255	0	255	0	255	255	0	128	128	0	0	69	165	215	230
B	0	255	0	0	255	0	255	255	0	0	128	128	0	0	0	140



# SOM principle



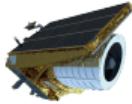
# Two implementations of the SOM

## On-line SOM

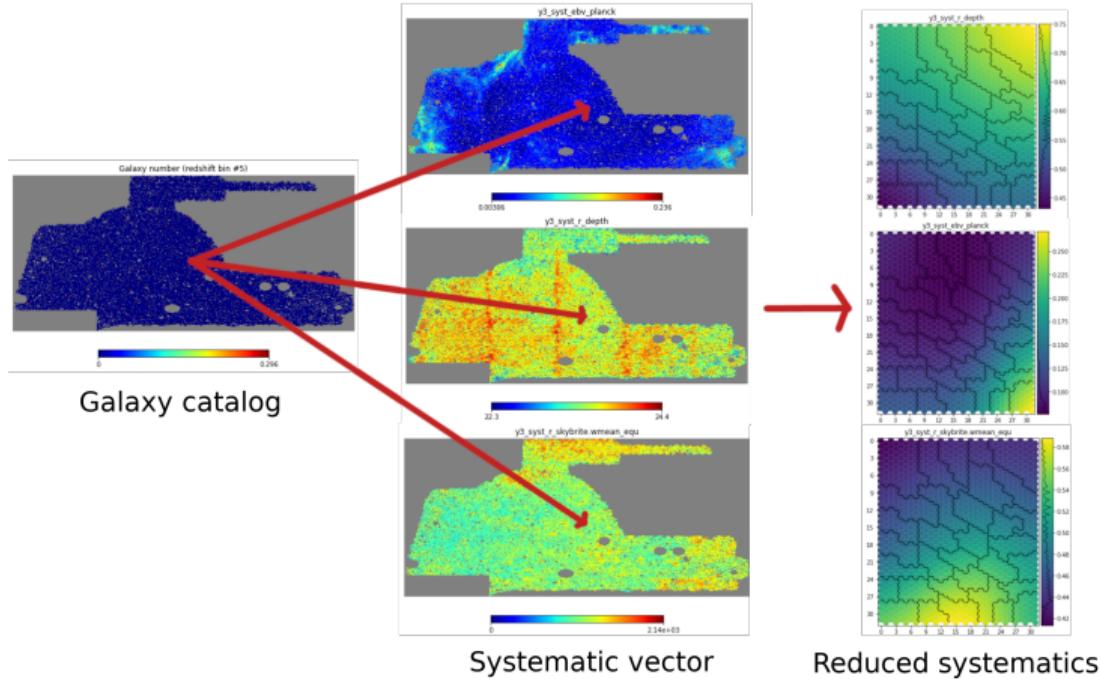
- Input vectors are presented sequentially
- The SOM latent space is updated for each new vector
- Can't be parallelized

## Batch SOM

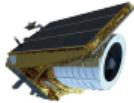
- Input vectors are all presented simultaneously
- The SOM latent space is updated once per epoch
- Highly parallelizable
- Ideal for a large amount of data, but do we lose information?



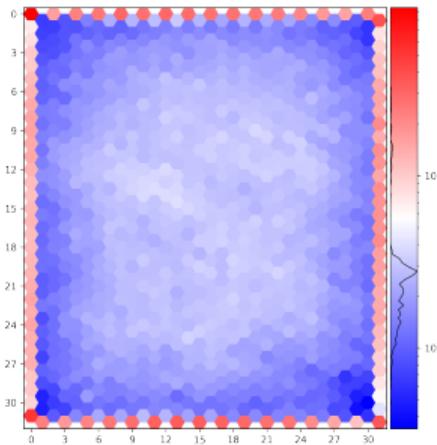
# 1. Training with galaxies



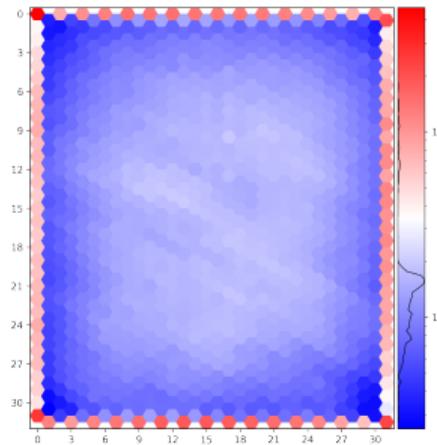
Reducing the observed sky  
into a small area while  
keeping (most of) the  
contamination information



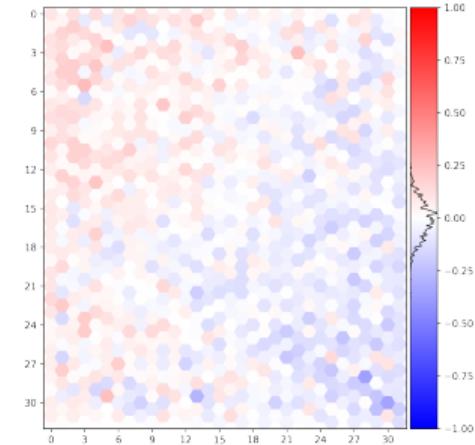
## 2. Activation maps



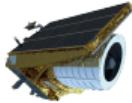
Galaxies activation map



Footprint activation map



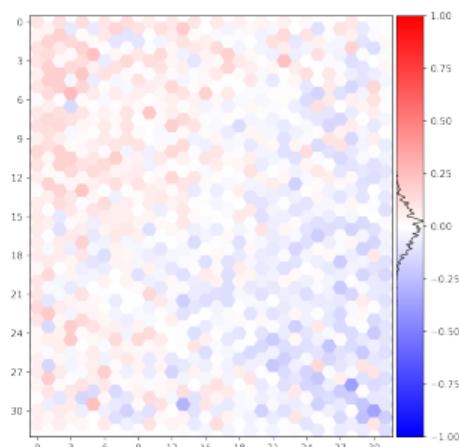
Reduced galaxy density



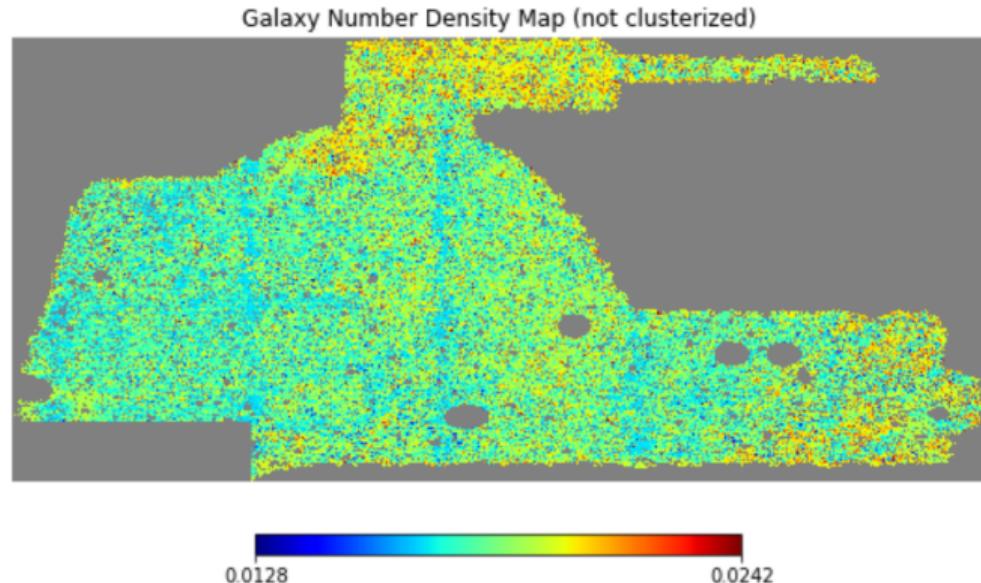
### 3. Decontaminated galaxy density

From reduced systematics

⇒ we reconstruct a galaxy density map containing the imprint of contamination effects



Reduced galaxy density

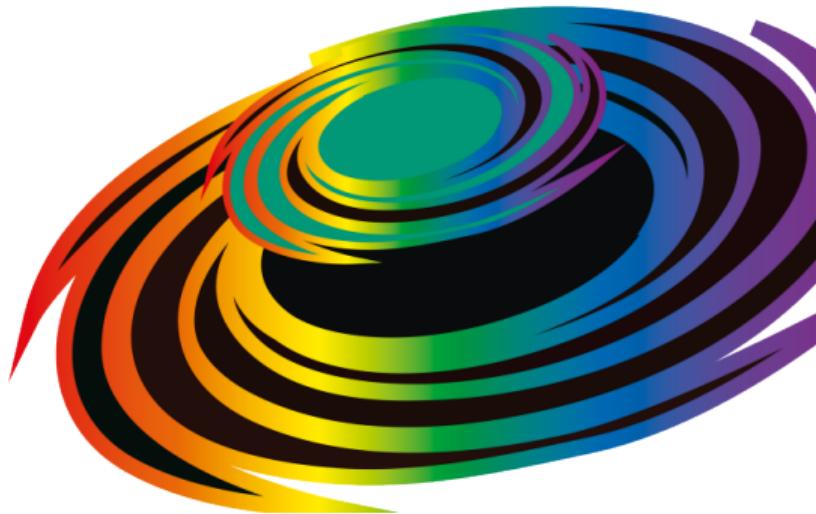


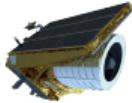
Projected back into physical space

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Decontaminating DES systematics

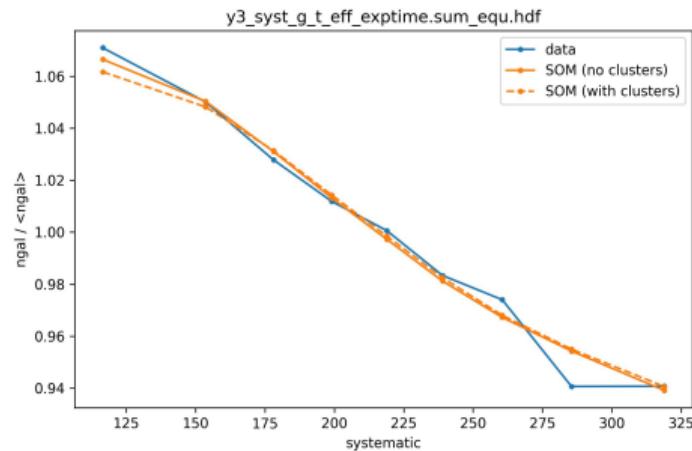
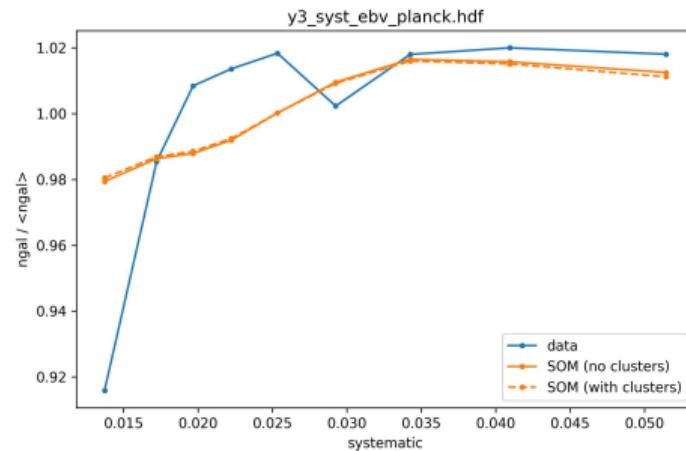
Scientific validation

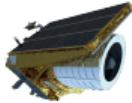




# Data-to-systematic correlations

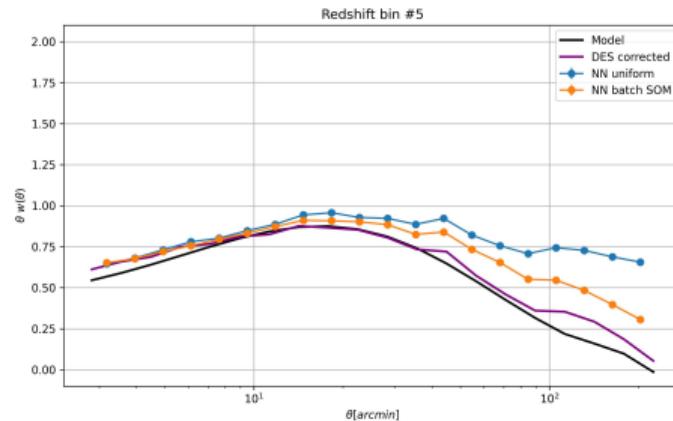
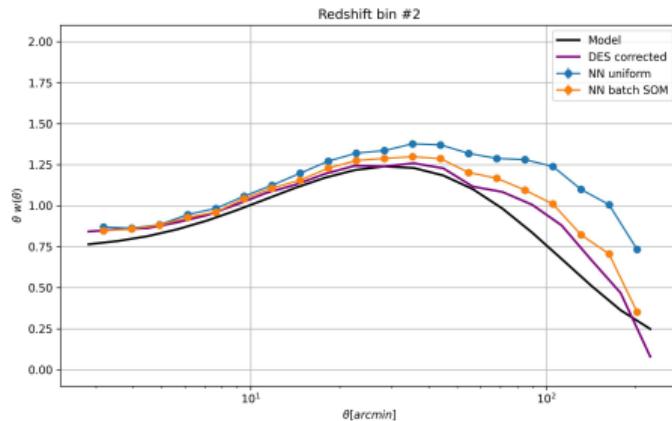
The algorithm is able to fit linear and non-linear effects





# Two-points angular correlation functions

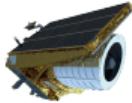
- Performed SOM decontamination with 35 DES systematics (4 photometric bands)
  - Decontamination is not perfect: DES selection cuts ?



Thanks Nicolas Tessore for the help with GLASS\* and TreeCorr†!

\*GLASS: Generator for Large Scale Structure, Tessore N. et al. [[arXiv:2007.14499](https://arxiv.org/abs/2007.14499)]

†TreeCorr, Jarvis M., Bernstein G. and Jain B . [[ascl](#)]



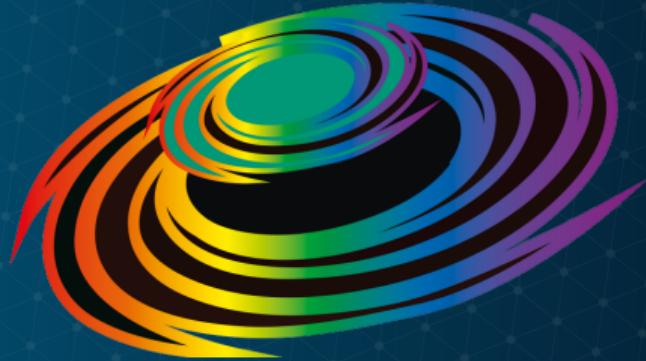
# Conclusions

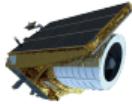
- Self-organizing maps as a promising alternative for systematic decontamination
  - On-line and batch implementations
  - The real challenges :
    - A huge amount of data, about  $15\ 000\ \text{deg}^2$ , billions of galaxies
    - Handling missing data
- Scientific validation ongoing on DES data
  - Two-points angular correlations
    - Need to tune the SOM parameters?
    - Apply the DES selections?
    - Project systematics into a principal component basis?
  - Other algorithms to implement
    - Iterative systematic decontamination
    - ElasticNet regularization (ongoing)
    - Principal component analysis (ongoing)
- First tests on Euclid PV phase data

Our decontamination tool is publicly available at [ipsc-euclid.github.io/decontamination](https://ipsc-euclid.github.io/decontamination)

Dedicated Slack channel: #sos-ou-le3-id-vmnz

Thanks for your attention!





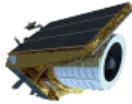
# SOM equations

On-line:

$$\vec{w}_k(t+1) = \vec{w}_k(t) + \alpha(e) h_{ck}(e) [\vec{x}(t) - \vec{w}_k(t)]$$

Batch:

$$\vec{w}_k(t+n) = \frac{\sum_{t'=t}^{t+n} h_{ck}(t') \vec{x}(t')}{\sum_{t'=t}^{t+n} h_{ck}(t')}$$



# Two-points angular correlation functions

