Analysis of storage times of UCN in gravitational quantum states

Les Houches, Febuary 15 2009

1 Lifetimes of quantum states

2 Losses of quantum states Verticality of the walls Holes in the corners Mirror's waviness Vibrations Earth rotation

Conclusion

Resonant transitions

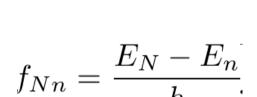


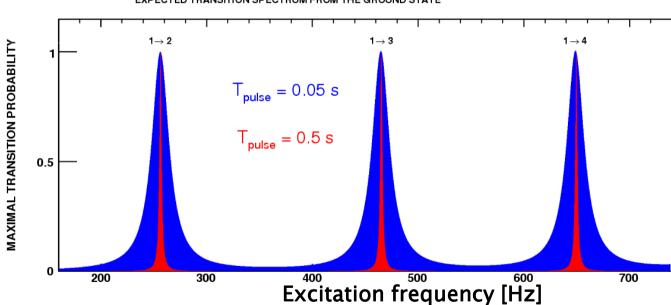
One applies harmonic excitation V(t) at frequency ω during time t

$$|4\rangle$$
 $|3\rangle$

$$P_{N\to n}(t) = \frac{\sin^2\left(\sqrt{(\omega - \omega_{Nn})^2 + \Omega_{Nn}^2} \frac{t}{2}\right)}{1 + \left(\frac{\omega - \omega_{Nn}}{\Omega_{Nn}}\right)^2} \quad \Omega_{Nn} = \frac{2}{\hbar} \langle n|V(z)|N\rangle$$







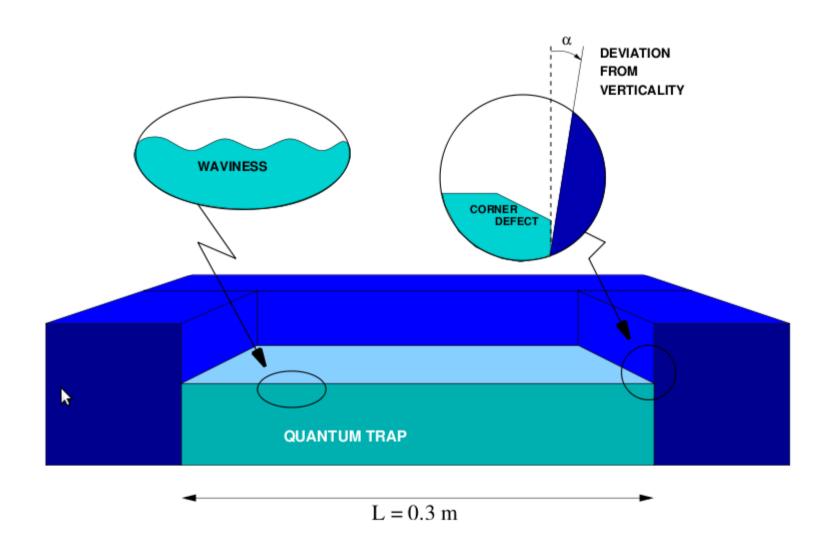
Lifetimes of quantum states are important

Width of the resonances

$$\Delta E \cdot T_{\text{pulse}} = h$$

- Pulse time needed to resolve neighbouring states: 10 ms
- Flow through mode $T \approx \frac{0.3 \text{ m}}{5 \text{ m/s}} \approx 50 \text{ ms}$

We want the lifetimes of quantum states to be larger than 50 ms and eventually larger than the beta decay lifetime

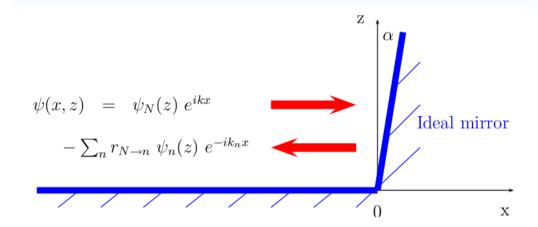


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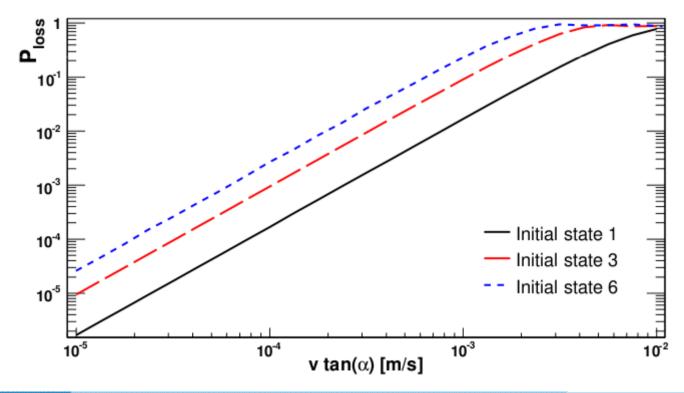
Verticality of the walls



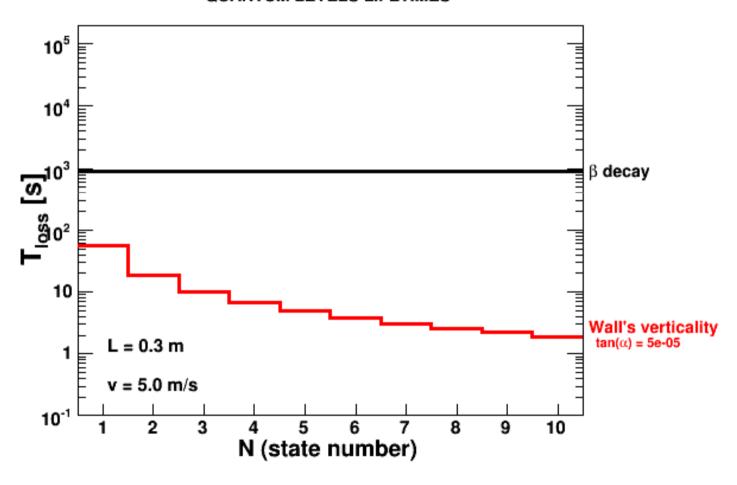
Imposing $\psi(\tan(\alpha)z, z) = 0$ we get:

$$P_N = 1 - |r_{N \to N}|^2$$

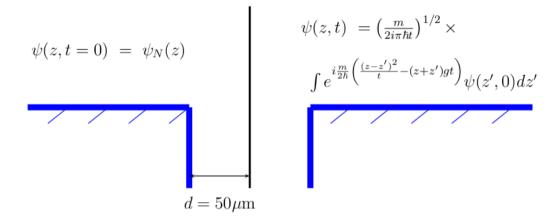
= $1 - |\langle N|e^{2ik\tan(\alpha)\hat{z}}|N \rangle|^2$



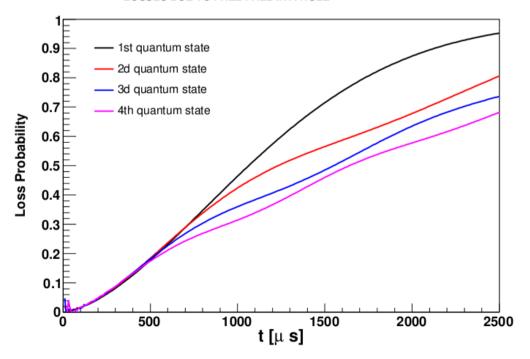
Verticality of the walls



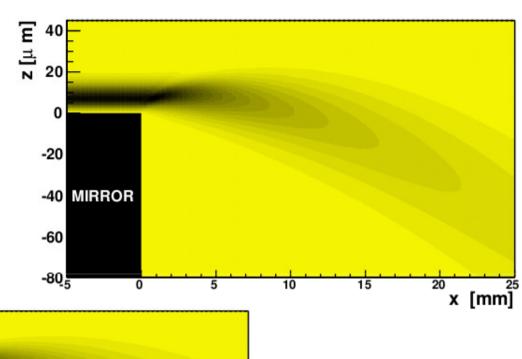
Holes in the corner

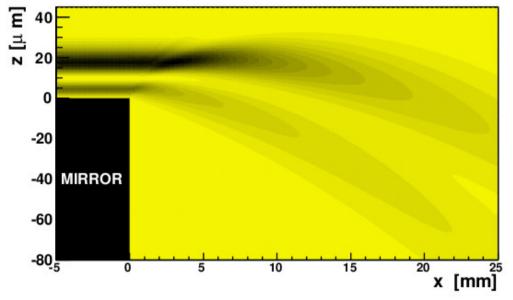


LOSSES DUE TO FREE FALL IN A HOLE



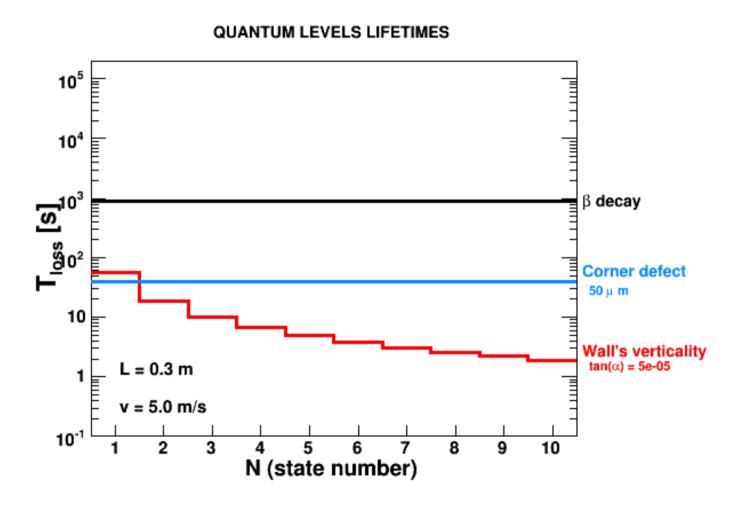
A nice picture





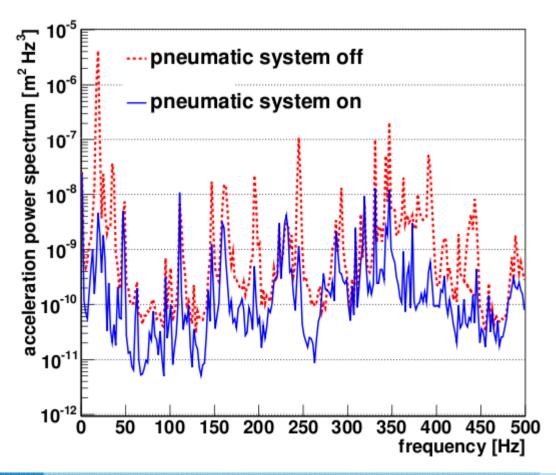
The free fall of quantum states 1 and 2

Holes in the corner

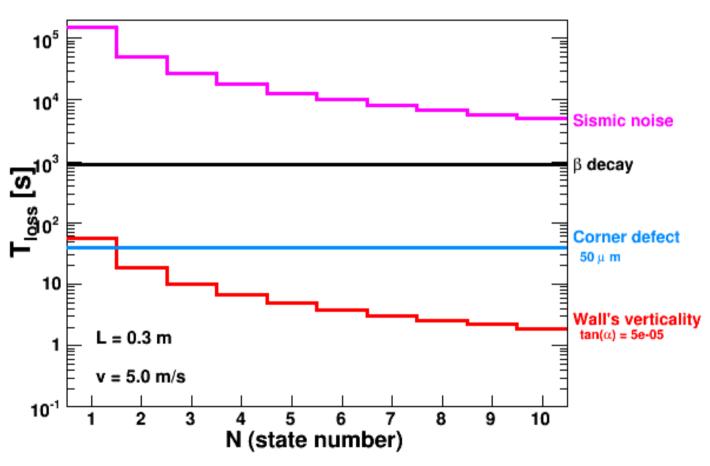


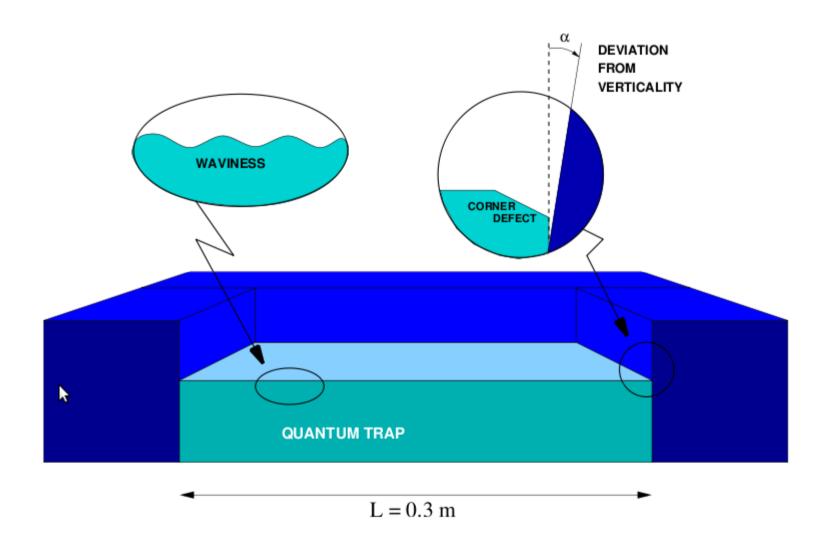
Losses due to vibrations

$$S_a(f) = \frac{1}{T} \left| \int_0^T a(t)e^{2i\pi f} dt \right|^2 \frac{1}{T_{N\to n}} = \left(\frac{\hbar}{mg}\right)^2 \frac{(2\pi f_{Nn})^4}{S_a(f_{Nn})}$$

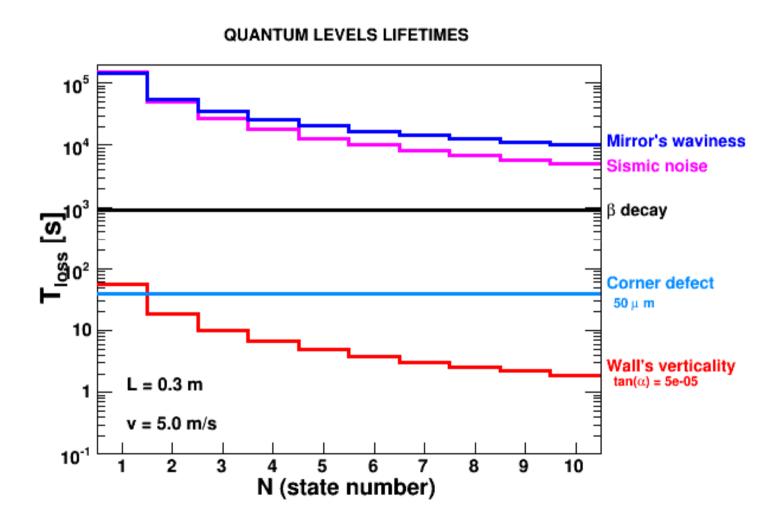


Losses due to vibrations





Losses due to mirrors waviness



Limitations du to Earth rotation

The Coriolis force $\vec{F}_c = 2m\vec{v} \times \vec{\Omega}$

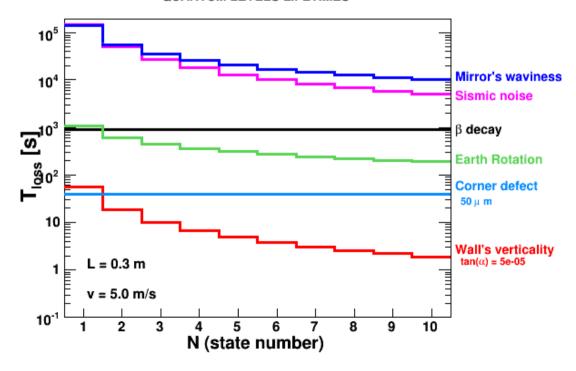
$$\vec{F}_c = 2m\vec{v} \times \vec{\Omega}$$

Modifies effectively the gravitational acceleration

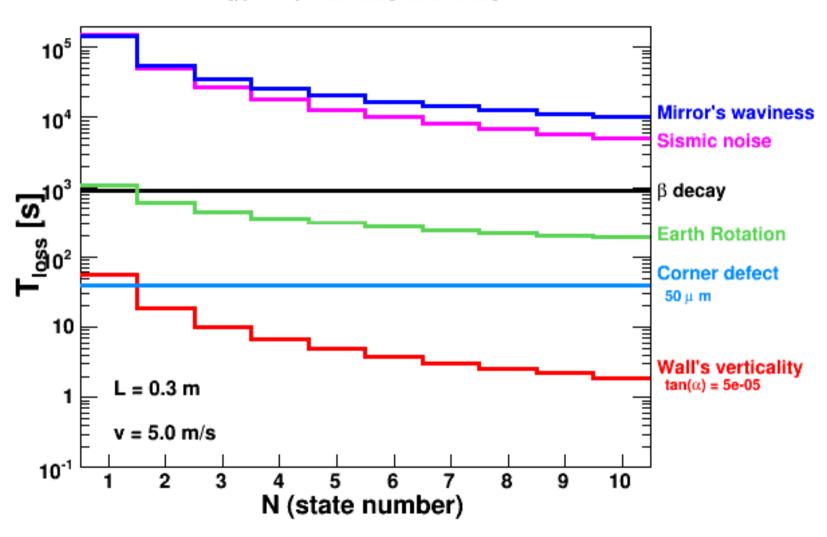
$$\vec{F}_c \cdot \vec{e}_z = 2mv\Omega\cos\lambda$$
 $g_{\text{eff}} = g - 2v\Omega\cos\lambda$

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Leading to Coriolis widening of the resonances



The hole picture



One last academic remark

Spontaneous decay of quantum states through graviton emission?

A semiclassical estimate Pignol Protassov Nesvizhevsky Class Quant Grav 24 (2007)

$$T_{decay} \simeq 10^7 \, s$$

Conclusions

Requirements in flow through mode T > 1 s

- 1) Mirrors verticality has to be better than 10⁻⁴ rad
- 2) Holes have to be smaller than 0.5 mm
- 3) Vibrations have been tested to be ok
- 4) Mirror's waviness should be ok

In latter stages

Beta decay and Coriolis force are the fundamental limitations Wall's verticality has to be better than 10^{-5} rad Holes have to be smaller than 50 µm Vibrations have been tested to be ok Mirror's waviness should be ok