

# **Mirrors for GRANIT, production, characterization**

## The LMA in the project



**LMA** : Laboratoire des Matériaux Avancés (Lyon)

❑ Features : Optical coatings

### Main projects

- ✓ Mirrors for gravitational waves antennas : VIRGO – LIGO
- ✓ Astrophysics telescope : LSST
- ✓ GRANIT

### Our role in GRANIT (ANR project)

- ✓ Expert and advice for the construction of the GRANIT cleanroom
- ✓ Expert and metrology of the different mirrors in GRANIT
- ✓ Coating of the trap, transportation and extraction optical pieces

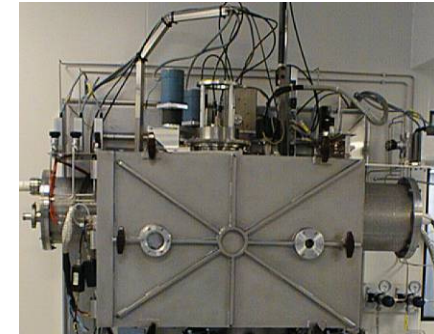
# LMA IBS (Ion Beam Sputtering) deposition Facilities

IBS technique is the deposition process to obtain optical stacks with the lower optical losses on large surface

- **Small IBS coater : DIBS**

- Able to coat homogeneously up to 3" substrates
  - ✓ Continual upgrades since 1990 : now equipped with an RF ion source (filament ion source before) like in the large coater
- Very flexible machine ⇒ ideal for prototyping

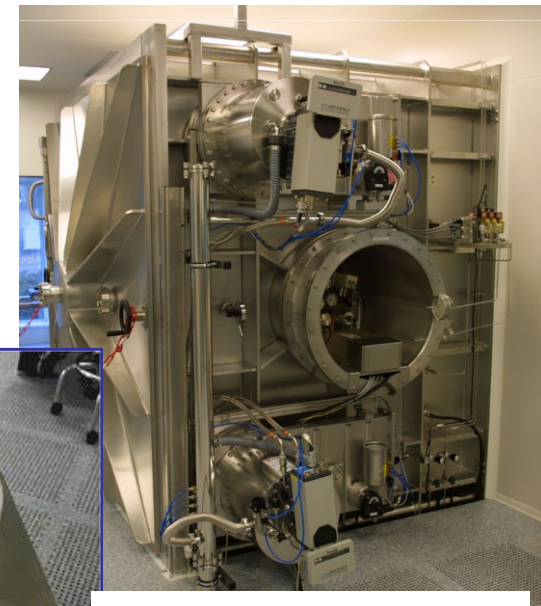
## Small IBS coater



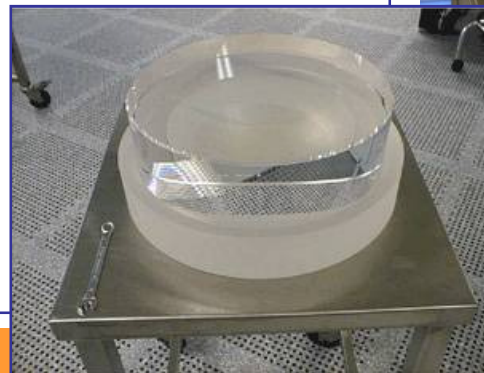
- **Large IBS coater : GC**

- ✓ 2,2 m X 2,2 m X 2,0 m inner deposition chamber
- ✓ Designed to coat substrates up to 1 meter diameter
- ✓ Used for VIRGO large mirrors since 2001
  - ❖ Periodic quarter wave doublet stacks ( $\text{Ta}_2\text{O}_5$  and  $\text{SiO}_2$ )
  - ❖ Between 130 and 180 nm layer thickness

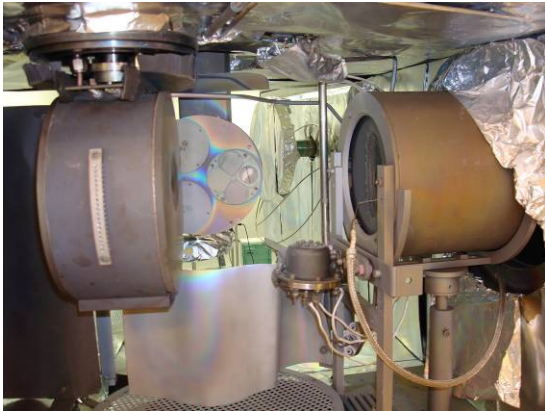
**350 mm diameter VIRGO mirrors**



## Large IBS Coater



## (D) IBS : (Dual) Ion Beam Sputtering

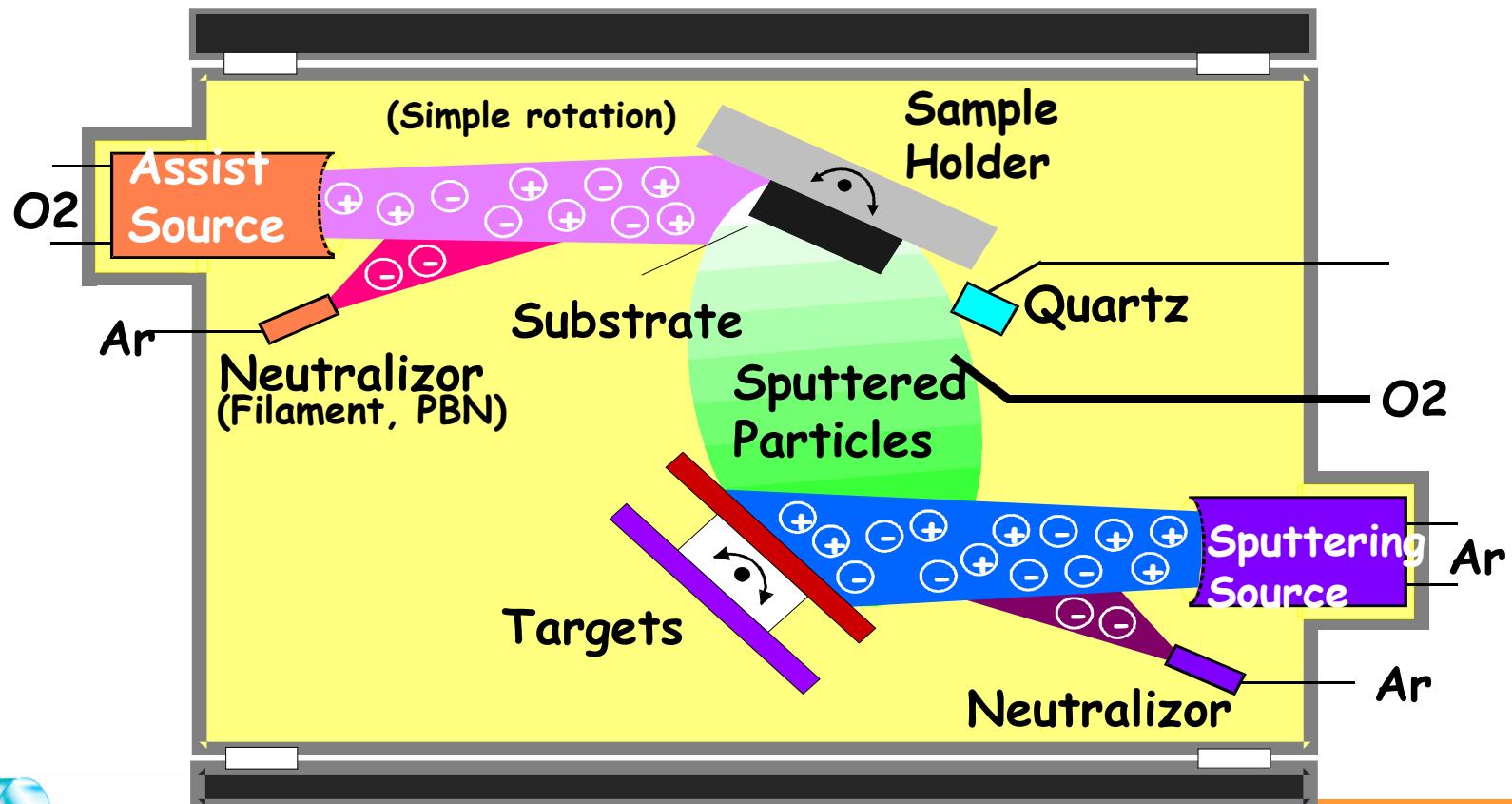


**Layers properties** : - ultra dense materials

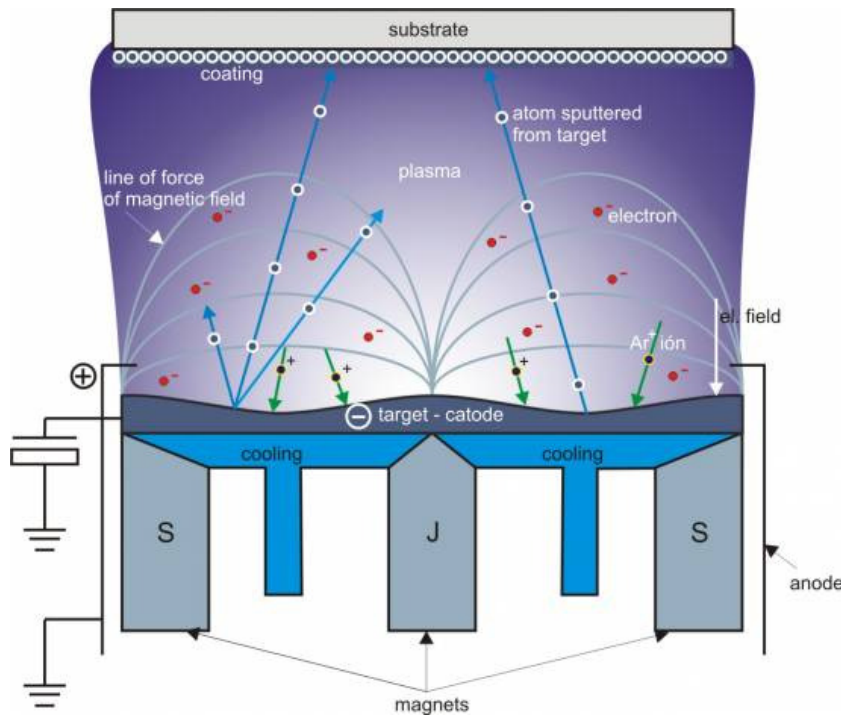
- ultra pure materials

**But** : - very slow deposition rate ( $0.5 \text{ \AA} / \text{sec.}$ )

- Stressed coatings



# Magnetron Cathodic sputtering



**Leybold Z550**  
**(LMA)**

**Layers properties** :

- dense materials
- pure materials
- low cost
- fast deposition rate  
(100 Å/sec.)

**But** : some difficulties with dielectric materials

ex : **Carbon**

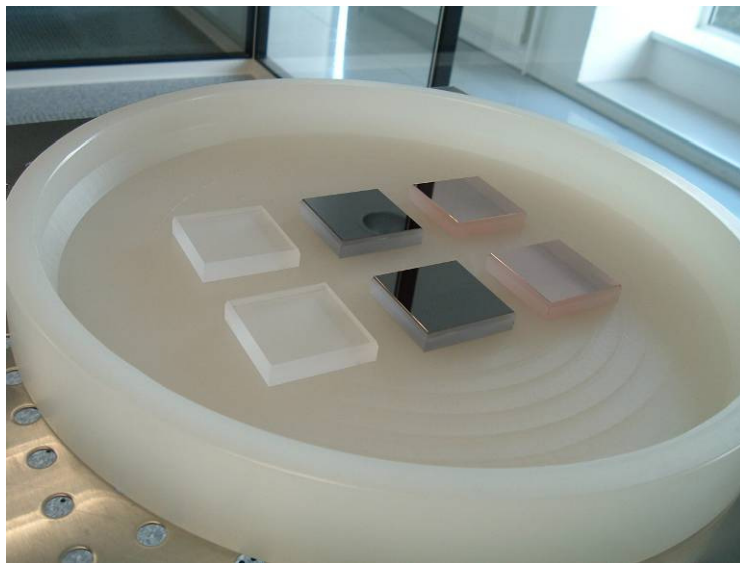
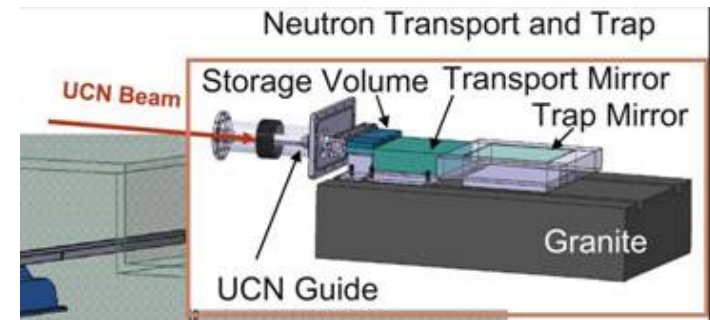


## Choice of the coating and material for the walls of the trap

### ➤ Choice and test of the coating layer

Sample : Silica 50 X 50 X 10 mm<sup>3</sup> from General Optics

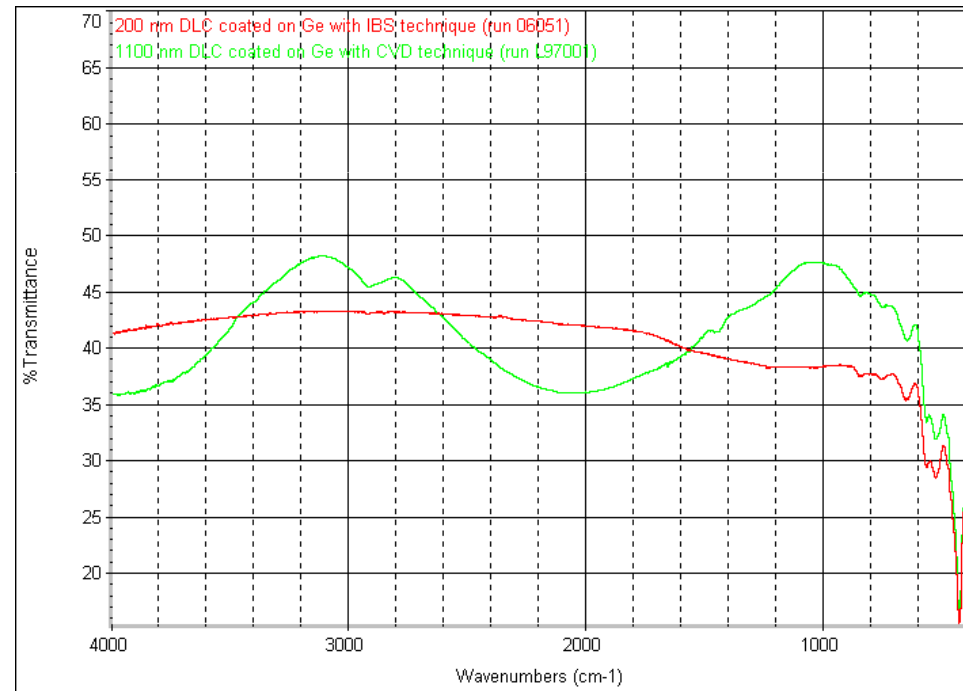
- ✓ 2 without coating
- ✓ 2 with 200 nm copper (RF magnetron sputtering)
- ✓ 2 with 200 nm Carbon IBS (IBS) not as hard as DLC CVD but without Hydrogen



**Left : uncoated silica**

**Middle : carbon coated silica**

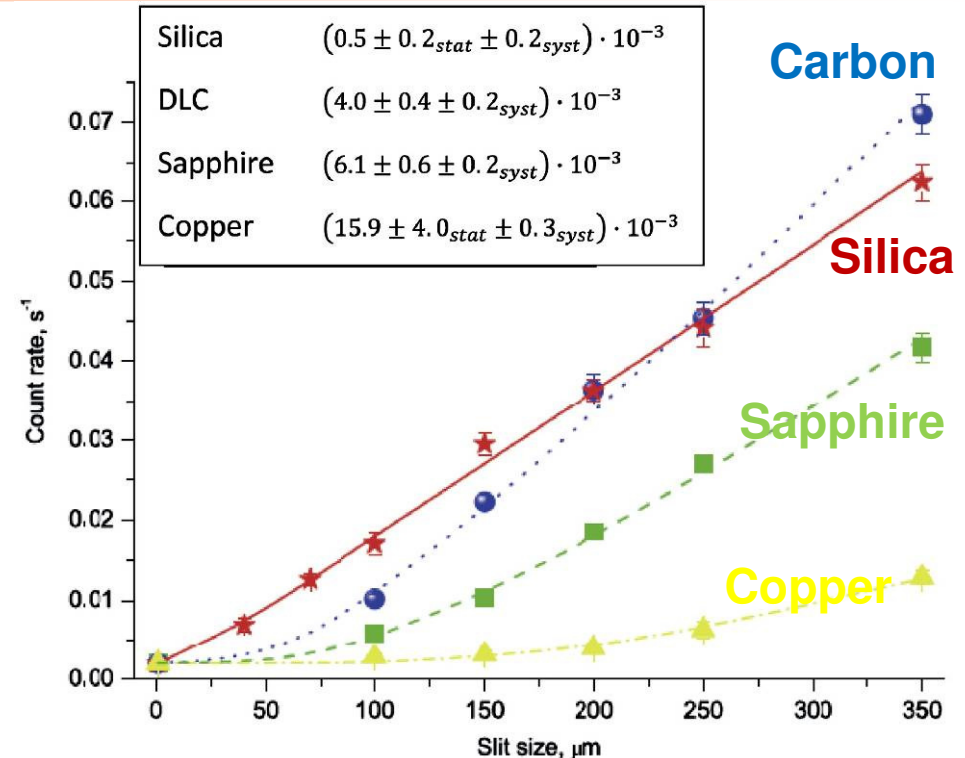
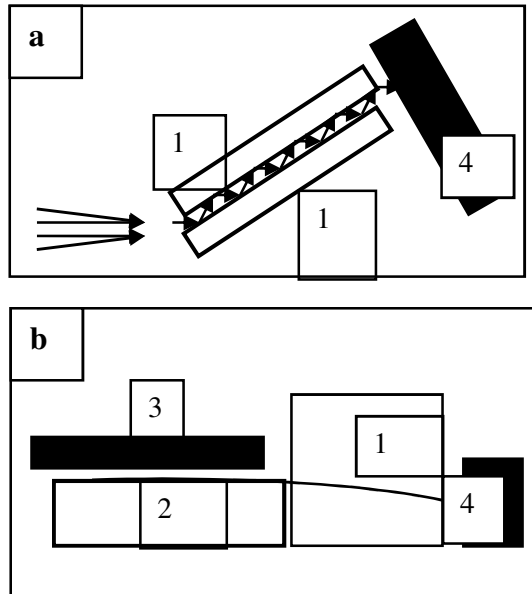
**Right : copper coated silica**



*IR transmission spectra of carbon layers on Ge substrate*

## Choice of the coating and materials

### Gravitational spectrophotometer



Total flux of transmitted neutrons as a function of slit size

### Conclusions :

- ✓ Silica for mirrors **OK**,
- ✓ Silica coated IBS Carbon **OK**, choice for use on the vertical walls of the GRANIT trap
- ✓ Silica coated copper (large coefficient of UCN total loss) and unstability of the coating

## Test on Stainless steel sheets

PVD test on 150 X 150 mm<sup>2</sup> stainless steel foils

- o 200 nm thick Carbon layer by Cathodic Sputtering deposition



*One week after coating*



*Few months after coating*

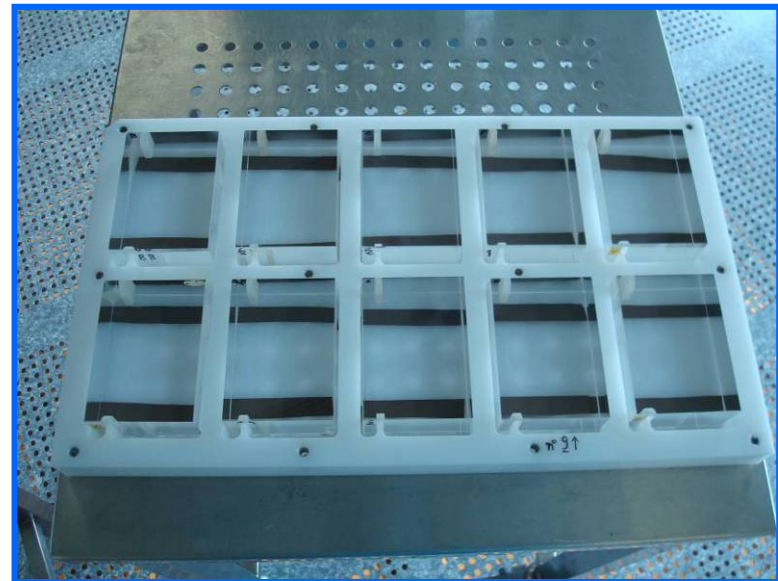
**Poor adhesion and complete delamination of the Carbon layer**

## Mirrors for the wall of the trap

- ✓ Due to the IBS coating chamber dimensions, each wall is divided into 3 parts
- ✓ Shaping and polishing done by SESO (Société Européenne de Systèmes Optiques )

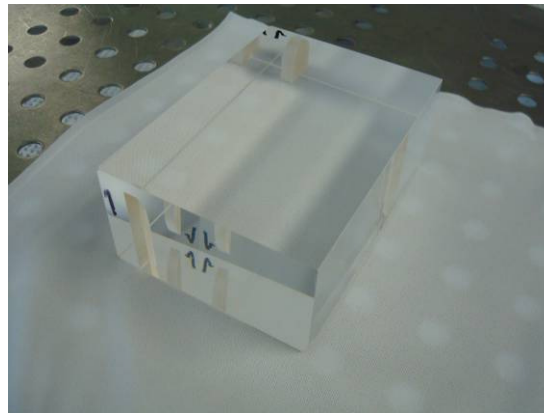


*Pôle d'Activités d'Aix-en-Provence (Les Milles)(France)*

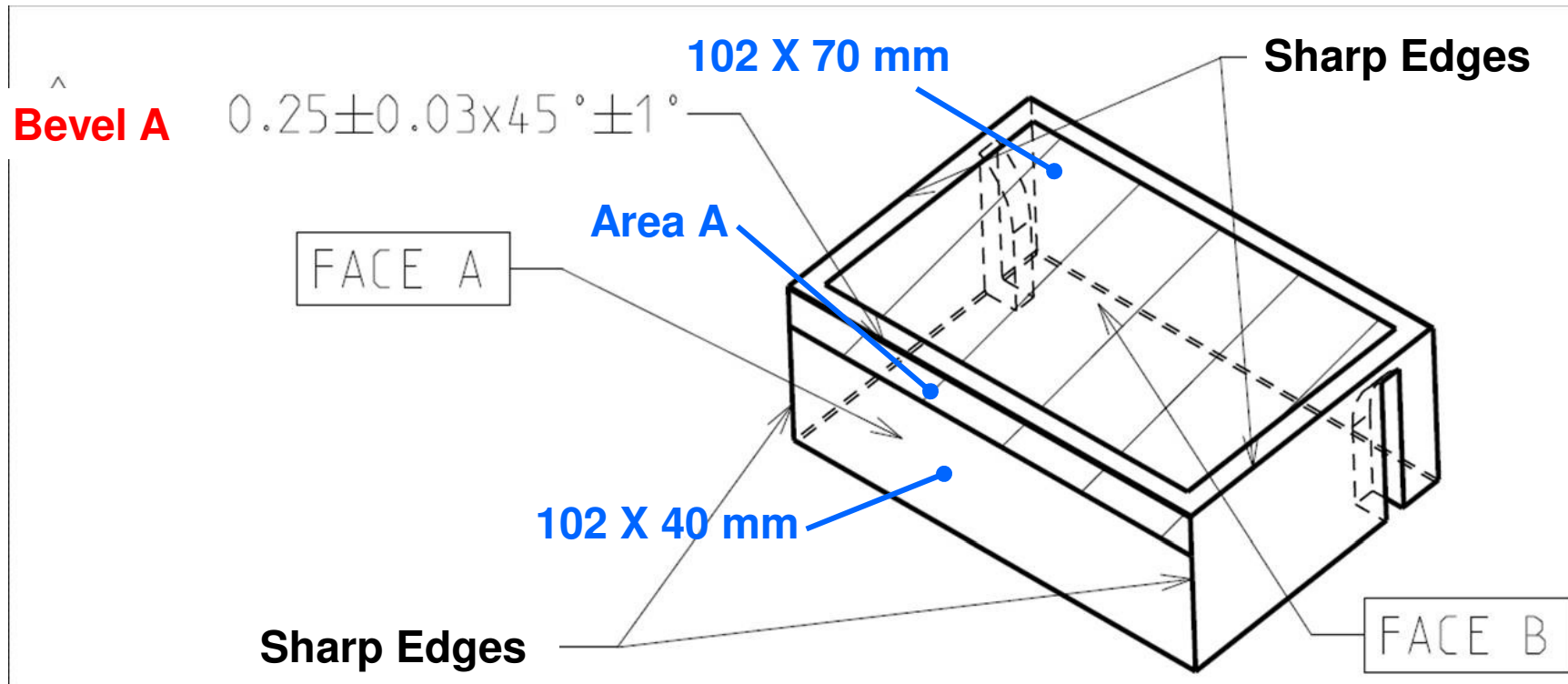


*10 over 17 mirrors for the walls*

- ✓ Optical characterization and Carbon coatings at LMA



## Mirrors for the wall of the trap : in details



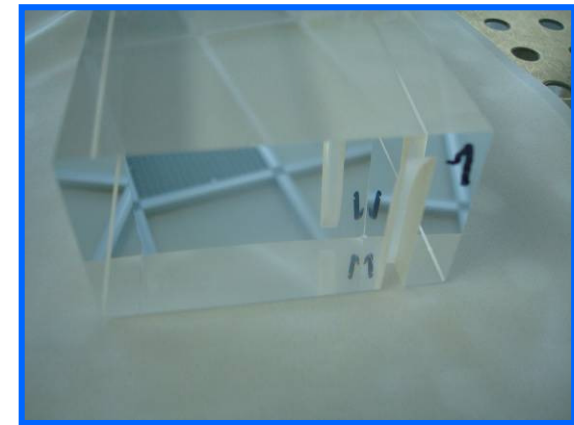
**Material : Silica**

**Face A : polishing  $\lambda/10$  PTV (Peak to Valley) or  $\lambda/4$  rms @ 633 nm**

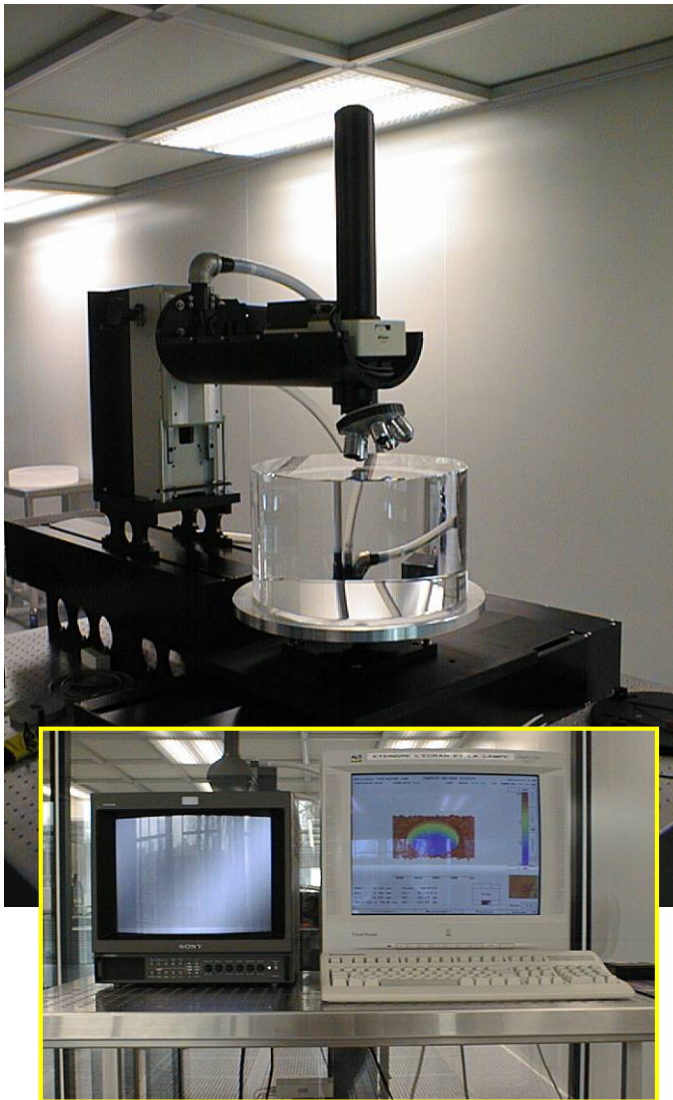
**Face B : polishing  $\lambda/4$  PTV on the area 92 X 60 mm**

**Area A : Roughness 2 Å rms**

**Bevel A : 0.25 mm +/- 0.03**



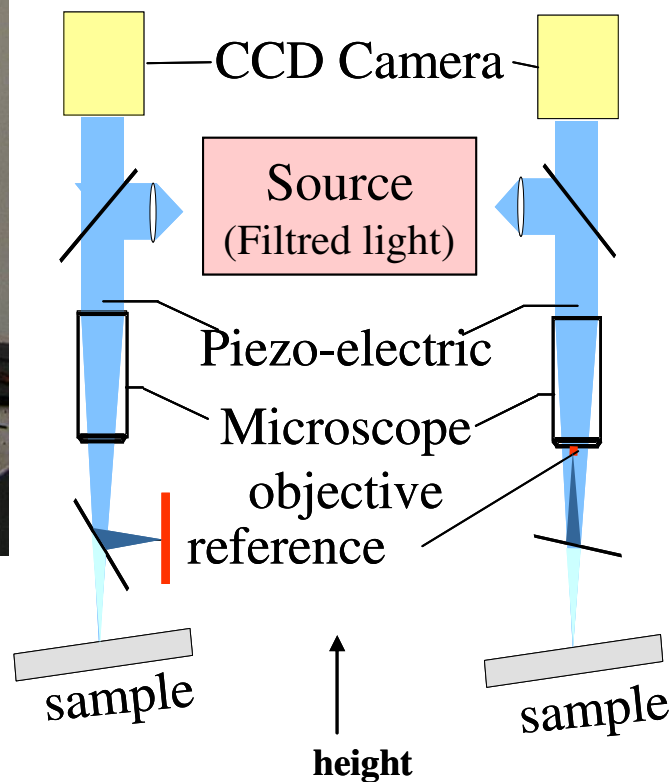
## Roughness measurement facility



### Interferometric Microscope

*Michelson*

*Mirau*

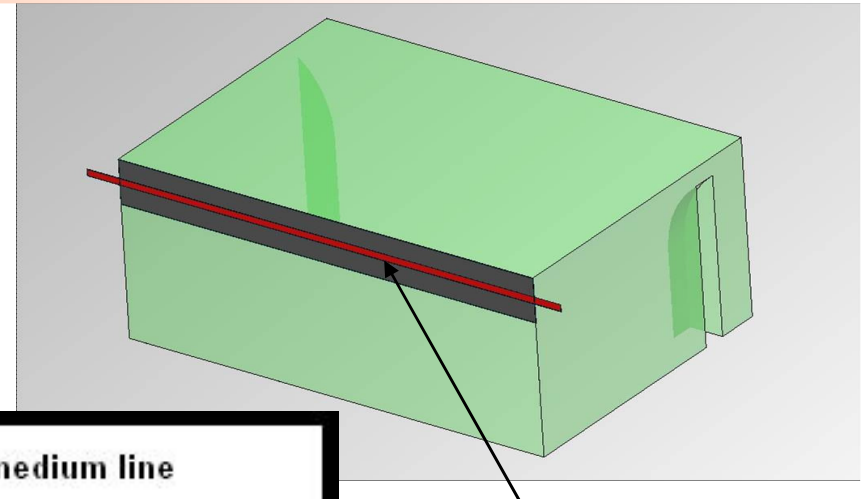


- ◇ Roughness  
(.1Å rms)
- ◇ Substrate point  
defects detection  
(.3 μm)

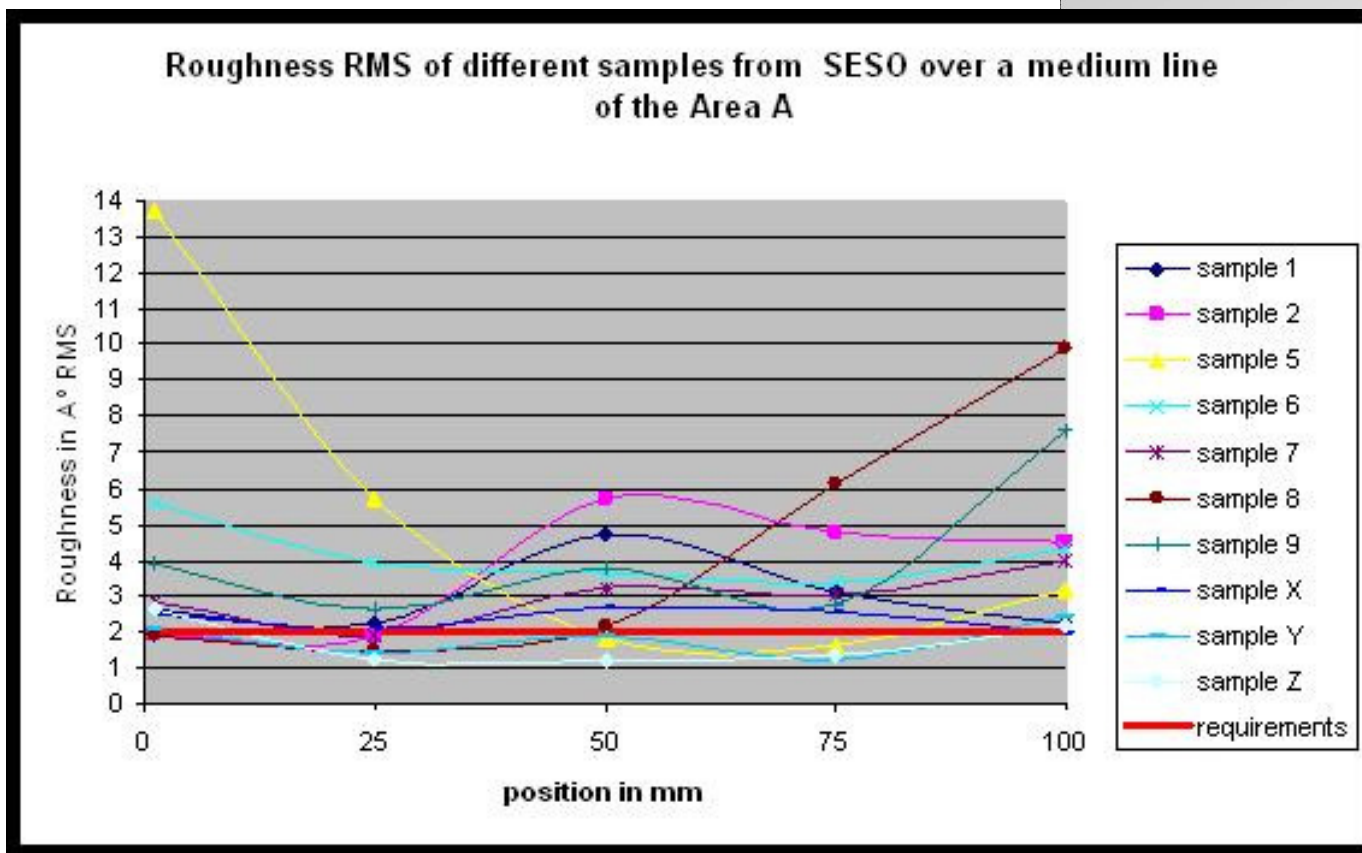
## Roughness measurements : Area A (samples 1 to Z)

Measurements done on 17 samples

Requirements : 2 Å rms



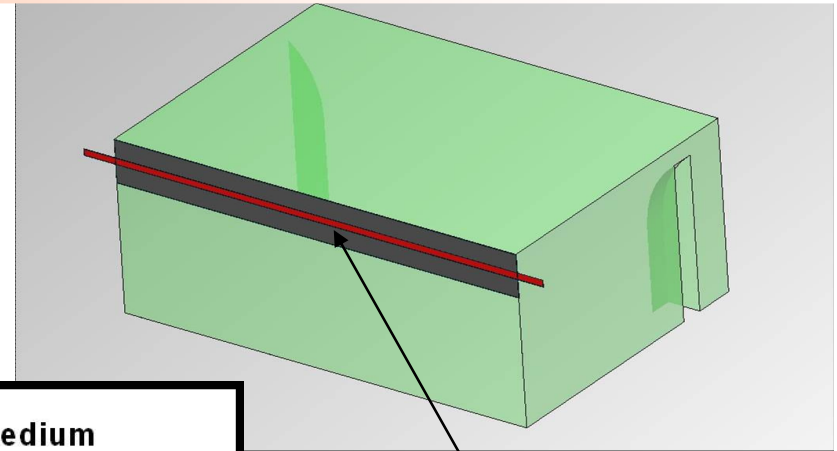
Medium line



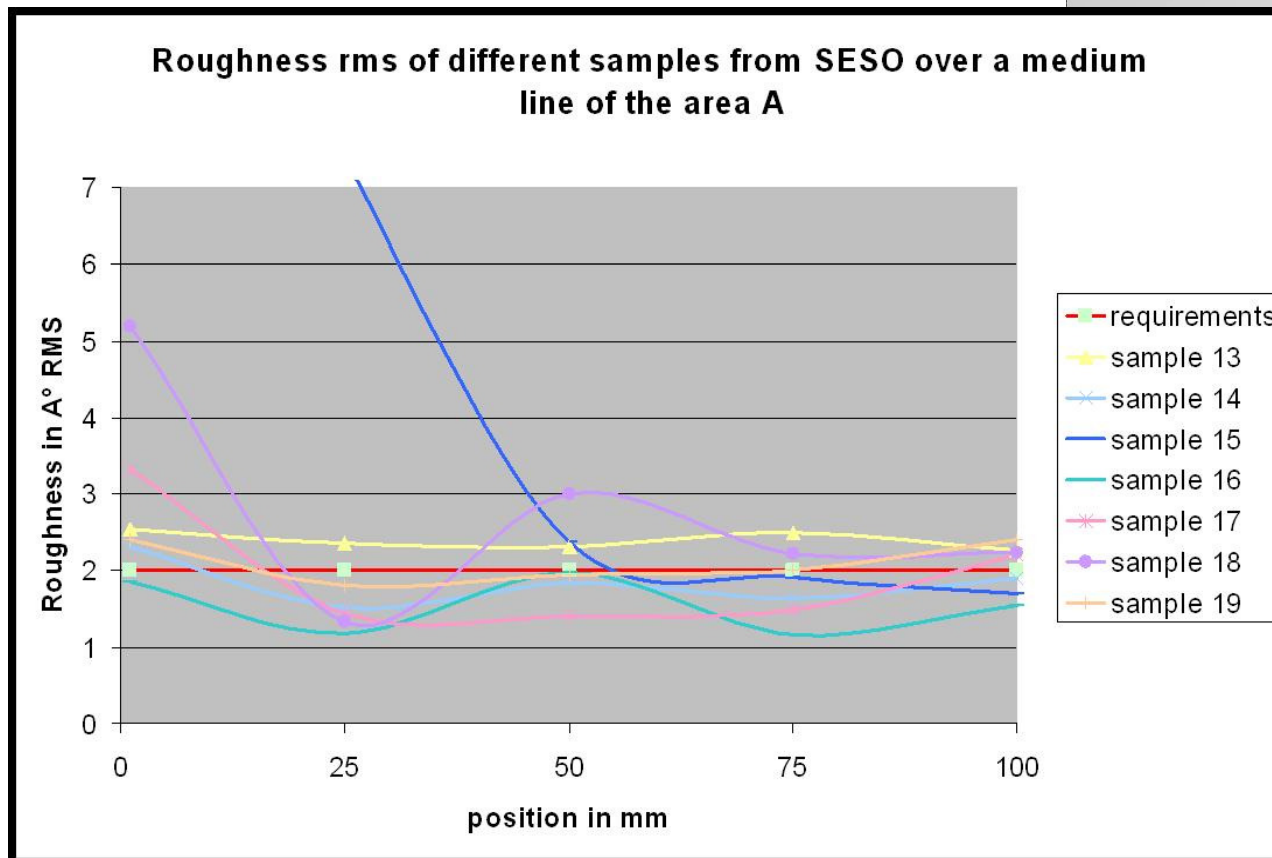
## Roughness measurements : Area A (samples 13 to 16)

Measurements done on 17 samples

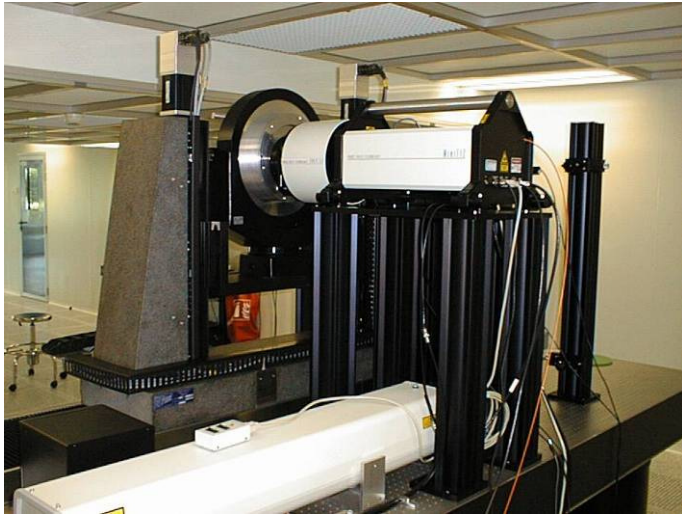
Requirements : 2 Å rms



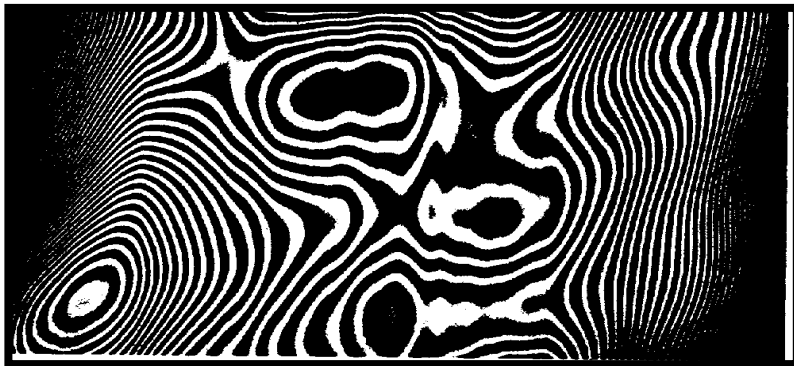
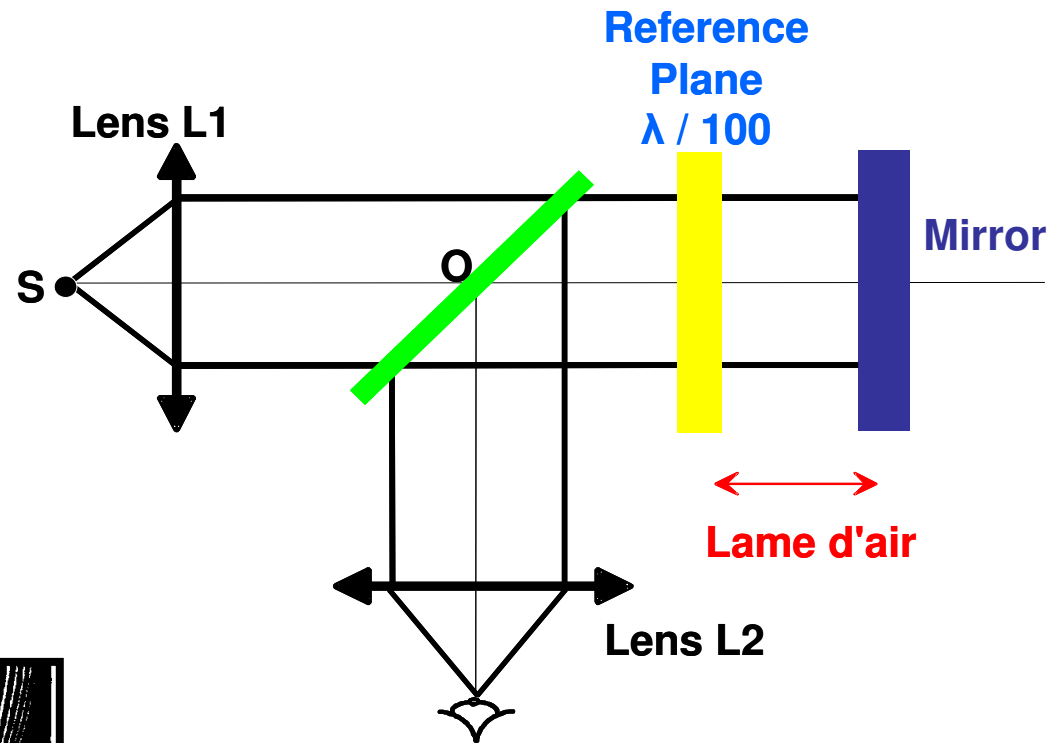
Medium line



## Planeity measurements



### Fizeau Interferometer



*Interferences between :*

- *reflected wave by the reference plane*
- *reflected wave by the mirror*

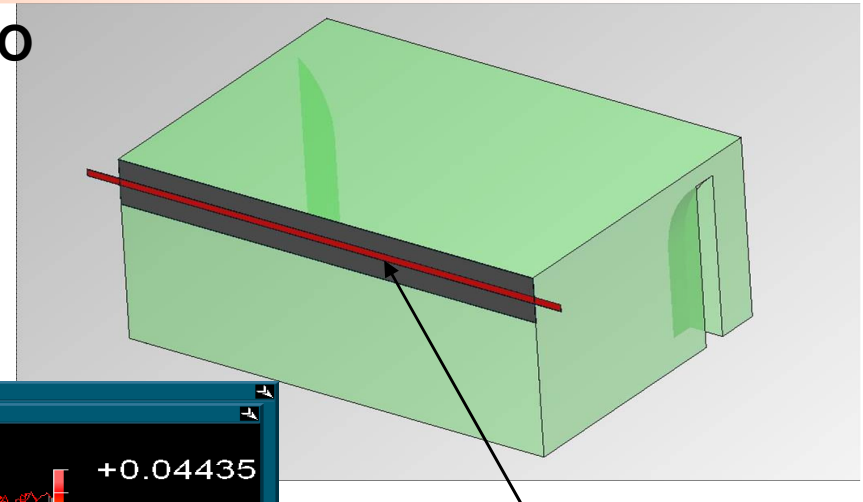
## Planeity measurements : Area A

Measurements done on 17 samples from SESO

Requirements planeity Area A Face A :

PTV :  $\lambda/10$  @ 633 nm (63 nm)

Rms :  $\lambda/40$  @ 633 nm (16 nm)

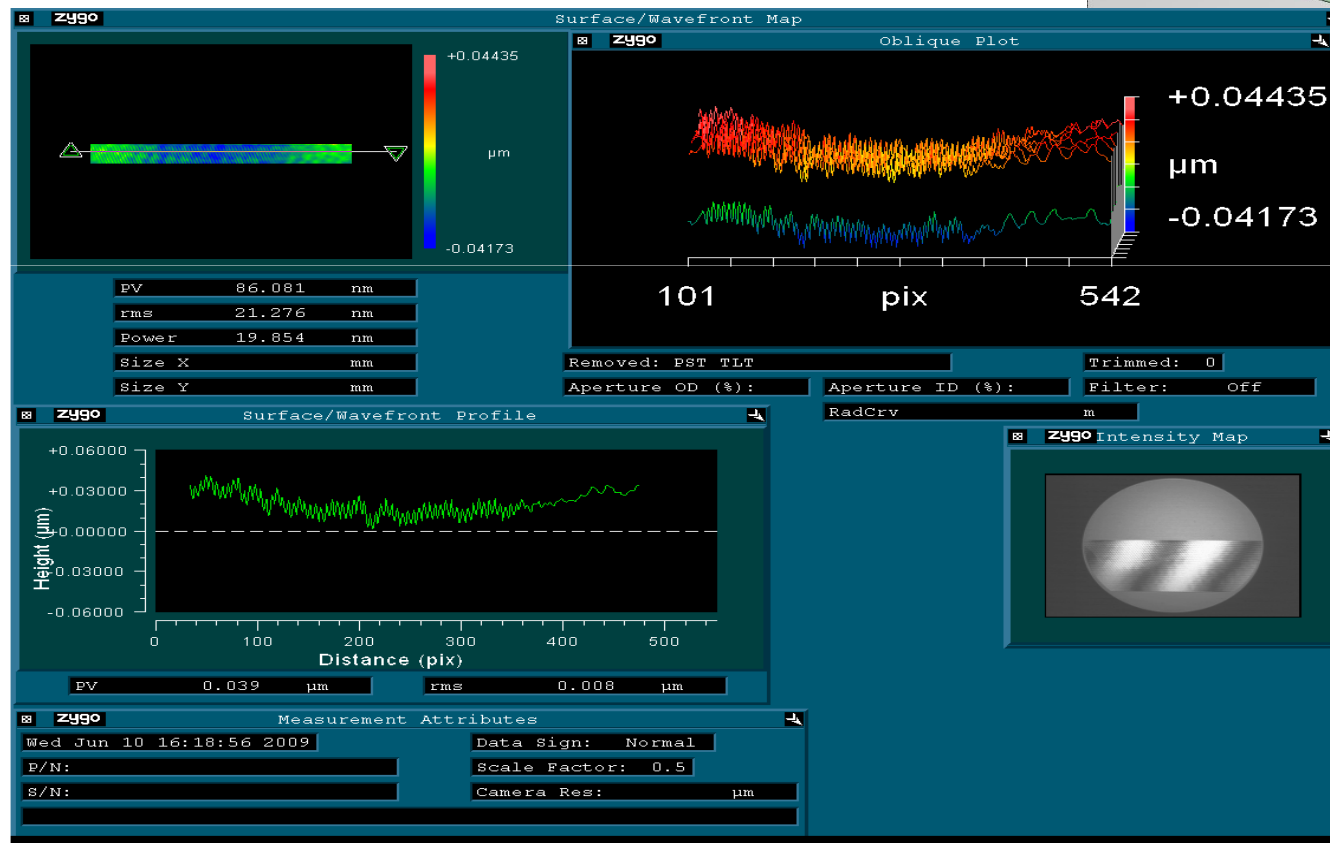


Medium line

Averaged values :

-PTV : 85 nm (high)

-Rms : 16 nm (OK)

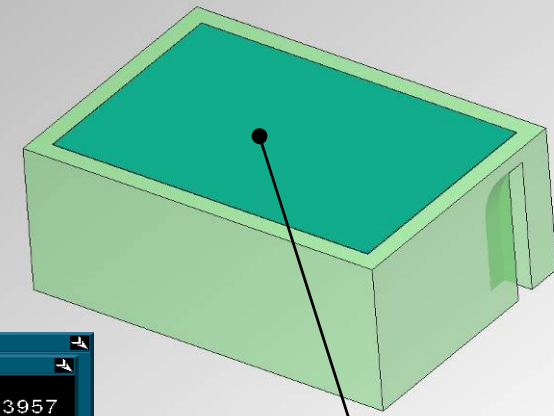


## Planeity measurements : Area B

Measurements done on 17 samples from SESO

Requirements planeity Face B :

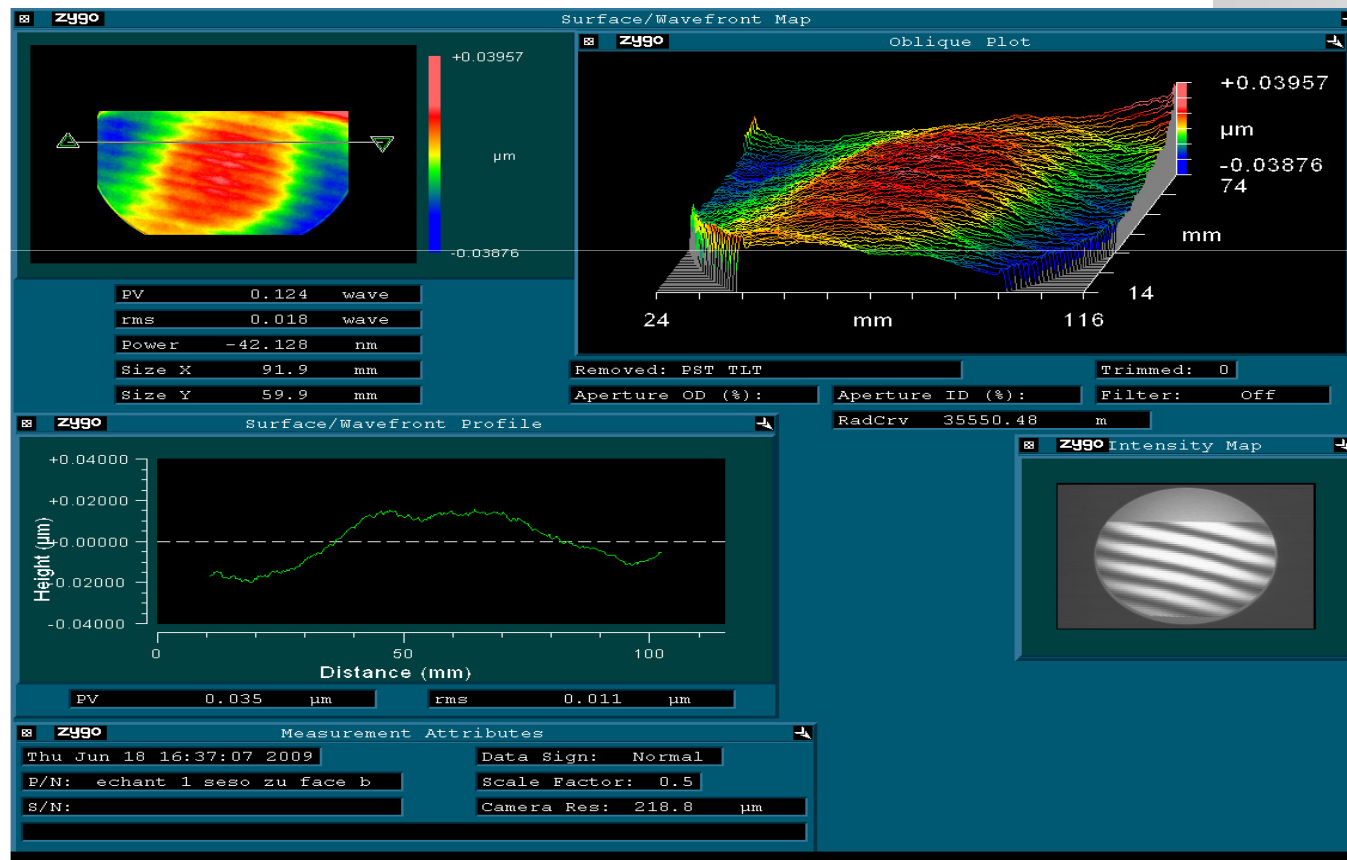
PTV :  $\lambda/4$  @ 633 nm (160 nm)



Face B

Averaged values :

- PTV : 120 nm (OK)



## Conclusions and Next to do

### Preliminary results

- ✓ Choice of coating process ( IBS) and material (200 nm Carbon)
- ✓ Shaping and polishing of 17 pieces for the walls of the trap (SESO)
- ✓ Characterization of the substrate for the walls of the trap

### Next to do

- ✓ Choice of the 12 better mirrors for the wall of the trap
- ✓ Coat the wall with 200 nm IBS carbon
- ✓ The pieces for the transportation and extraction will arrive at the end of february from SESO
- ✓ Characterization and coating of the remaining pieces