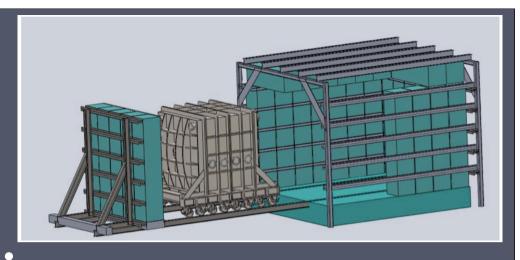
The DRIFT Directional Dark Matter Experiments....



Neil Spooner, University of Sheffield DRIFT collaboration, Boulby (Occidental, UNM, USC, Sheffield, Edinburgh, STFC)

- Directionality
- DRIFT II backgrounds
- DRIFT II unblind and blind analysis
- DRIFT III and scale-up

for the collaboration - contributions from many —



CYGNUS Workshops





Boulby

CYGNUS20

CYGNUS 2007

First Workshop on Directional Detection of Dark Matter

International Journal of Modern Physics A © World Scientific Publishing Company



bratory, UK rs meeting

THE CASE FOR A DIRECTIONAL DARK MATTER DETECTOR AND THE STATUS OF CURRENT EXPERIMENTAL EFFORTS

S. AHLEN, 1 N. AFSHORDI, 23,31 J. B. R. BATTAT, +,15 J. BILLARD, 11 N. BOZORGNIA, 3 S. BURGOS, 21 T. CALDWELL, 15,22 J. M. CARMONA, 12,13 S. CEBRIAN, 12,13 P. COLAS, 4 T. DAFNI, 12,13 E. DAW, 26 D. DUJMIC, 18 A. DUSHKIN, 2 W. FEDUS, 18 E. FERRER, 4 D. FINKBEINER, 6 P. H. FISHER, 18 J. FORBES, 21 T. FUSAYASU, ¹⁶ J. GALAN, ^{12,13} T. GAMBLE, ²⁶ C. GHAG, ⁵ I. GIOMATARIS, ⁴ M. GOLD, ¹⁸ H. GOMEZ, ^{12,13} M. E. GOMEZ, ⁷ P. GONDOLO, ²⁰ A. GREEN, ²⁰ C. GRIGNON, 11 O. GUILLAUDIN, 11 C. HAGEMANN, 18 K. HATTORI, 10 S HENDERSON 15 N HIGASHI 10 C IDA 10 F I IGUAZ 12,13 A INGLIS 1 I. G. IRASTORZA, 12,13 S. IWAKI, 10 A. KABOTH, 18 S. KABUKI, 10 J. KADYK, 14 N. KALLIVAYALIL, 15 H. KUBO, 10 S. KUROSAWA, 10 V. A. KUDRYAVTSEV, 26 T. LAMY, 11 R. LANZA, 15 T. B. LAWSON, 26 A. LEE, 15 E. R. LEE, 18 T. LIN. 6

F. MAY White Paper 112 authors

A. TAKEDA, ²⁸ T. TANIMORI, ¹⁰ K. TANIUE, ¹⁰ A. TOMAS, ^{12,13} H. TOMITA, ¹ K. TSUCHIVA 10 J. TURK 18 E. TZIAFERI 26 K. UENO 10 S. VAHSEN 14

International Journal of **Modern Physics A** Vol. 25, No. 1 (2010) 1-51

¹¹Laboratoire de Physique Subatomique et de Cosmologie

tter Detection



of Technology 11-13, 2009

¹⁰ Kyoto University Kitashirakawa-oiwakecho, Sakyo-ku, Kyoto, 606-8502, Japan

June 8-10, Aussois, France

13 Directional R&D Challenges

Techniques

Implementation

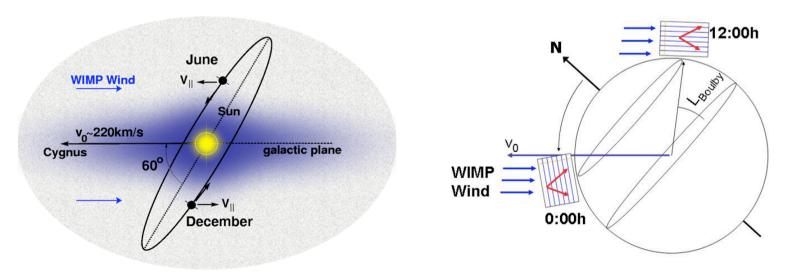
Theory

- 1. Development/demonstration of directional sensitivity for low energy
- 2. Development/demonstration of head/tail discrimination for low energy nuclear recoils
- 3. Development/demonstration of background discrimination and
- 4. Demonstration of robustness and stability for long-term operation
- 5. Selection/optimisation of gases or gas mixtures for SD and SI sensitivity
- 6. Determination of gas parameters, gains, sensivities, W and form factors
- 7. Development of end-to-end simulations
- 8. Development/optimisation of readout techniques and instrumentation
- 9. Optimisation of gas pressure (or pressures) for directional and nondirectional operation.
- 10. Development/demonstration of cost reduction techniques for scale-up
- 11. Assessment of infrastructure requirements size, depth, vetos?
- 12. Study of halo / cosmology theory and likely science reach.
- 13. Study of wider applications: KK axions? DAMA?

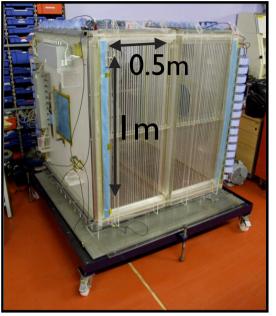
Dark Matter Signals

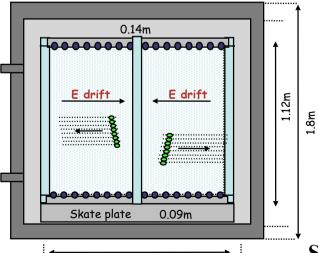
and directionality

- Motion of the Earth through a static WIMP 'halo' -> Earth is subject to a 'wind' of WIMPs
- of average speed ~220kms⁻¹ coming roughly from the direction of the constellation Cygnus.
- The Earths rotation relative to the WIMP wind -> Direction changes by $\sim 90^{\circ}$ every 12 hours

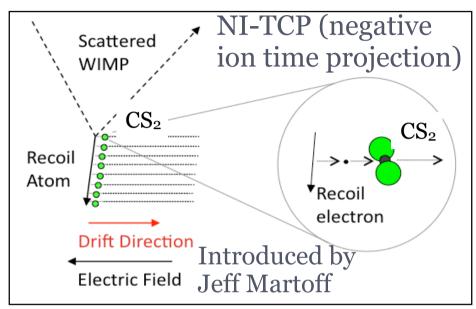


D<u>RIFT II Con</u>cept





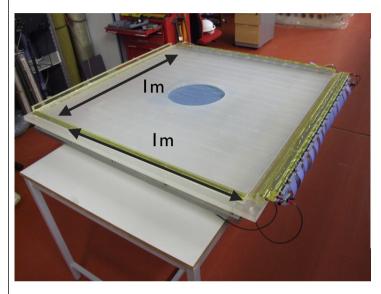
1.00m



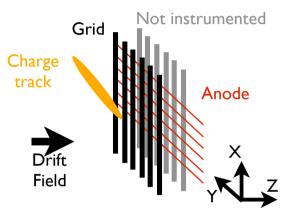
- 1 m³ active volume back to back MWPCs
- Gas fill 40 Torr CS₂ => 167 g of target gas
- 2 mm pitch anode wires left and right
- Grid wires read out for Δy measurement
- Veto regions around outside
- Central cathode made from 20 μm diameter wires at 2 mm pitch
- Drift field 624 V/cm
- Modular design for modest scale-up

S. Burgos et al., Nucl. Instr. Meth. A 584, 114 (2008)

MWPC Readout



- Anode plane of 512 20µm wires with 2mm pitch
- 2 cathode planes of 512
 100μm wires perpendicular to anode plane, 2mm pitch one of which is read out

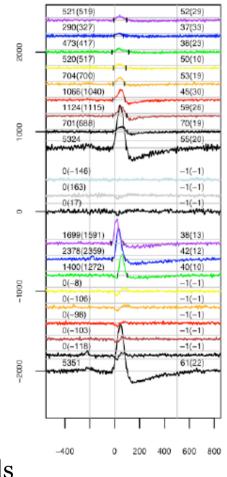


 ΔX : Number of anode wires crossed

 ΔY : Progression across

grid wires

 Δ Z: Drift time between start and end of track



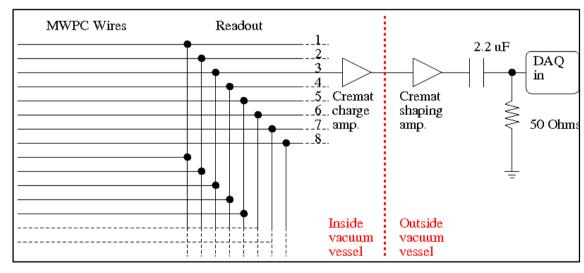
Time (µS)

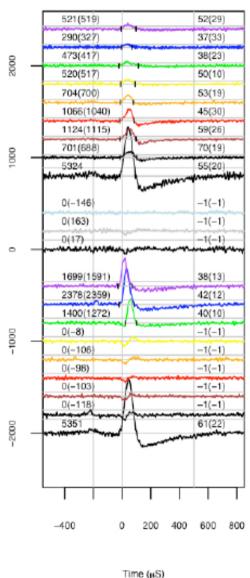
Multiplexed to 18 channels of digitised waveform output for 1m² readout plane

Simple, cheap & scalable

MWPC Readout

- Multiplexed to 8 lines of output per plane
- 1m² 2D readout 18 ADC channels (8×anode, 8×grid, 2×veto)
- Cheap and scalable
- No absolute x-y position only dx, dy



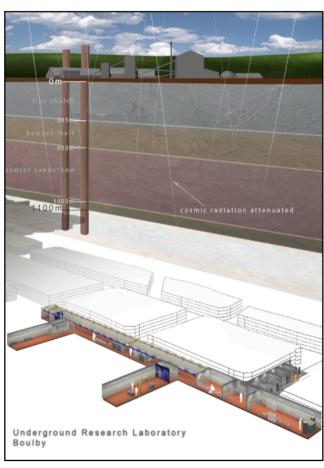


Boulby Mine (UK)

• Current site (1.1 km deep) in salt rock

• Deeper excavation underway in dolomite rock

• Suitable for a large TPC!











DRIFT II Shielding Simple and cheap poly pellet neutron shielding

- Lab at depth of 1100m (2800 m.w.e)
- Cosmic ray flux = 4.1×10^{-8} cm⁻² s⁻¹ M. Robinson et. al, NIM Ā 511 (2003)]
- Polypropylene pellets of >67cm depth on all sides
- Equivalent to 40g/ cm² solid hydrocarbon passive shielding
- Lead shielding not required due to detector's inherent insensitivity to electron recoil events









DRIFT II Summary

- Operational in the Boulby Mine since 2001
- DRIFT-I, DRIFT-IIa, DRIFT-IIb, DRIFT-IIc, DRIFT-IId
 - Low threshold potential (< 3 keV, S-recoil)
 - Directional signatures (and 3D reconstruction)
 - Head-tail (sense) is feasible, and verified by theory
 - Radon backgrounds (RPR) understood reduced
 - Fiducialisation via +ve ions looks to work
 - Thin cathode works
 - Neutron backgrounds understood
 - Stable and safe operation with CS₂ and CF₄
 - Competitive SD WIMP-P limits with directionality

B. Morgan, A.M. Green and N.J.C. Spooner, Phys Rev D71 (2005) 103507

P. K. Lightfoot, N. J. C. Spooner et al., Astropart. Phys. 27 (2007) 490

S. Burgos et al., Astropart. Phys. 28 (2007) 409

N.J.C. Spooner, J. Phys. Soc. Japan, 76 (2007) 11101

E. Tziaferi et al., Astropart. Phys. 27 (2007) 326

K. Pushkin et al., (2008) arXiv:0811.4194

S. Burgos et al., Nucl. Instrum. and Meth. in Phys. Res. A 584 (2008) 114

S. Burgos et al., JINST 4 (2009) P04014

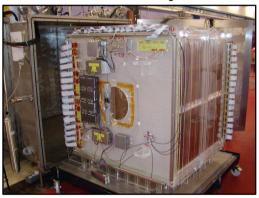
S. Burgos et al., Nucl. Instrum. and Meth. in Phys. Res. A600 (2009) 417

S. Burgos et al., Astroparticle Physics 31 (2009) 261

N.J.C. Spooner et al. Astroparticle Physics 34 (2010) 284

E. Daw et al, sub Astroparticle Physics (2011) - arXiv:1012.5967

BIG PROGRESS in the last 2 years







Ready for new experiment DRIFT III Module - 24m³

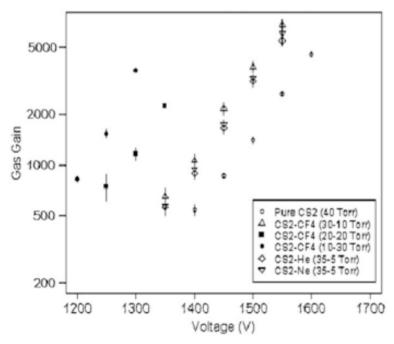
DRIFT IId - Gas Mixtures

• Recent emphasis on backgrounds and spin-dependent limits

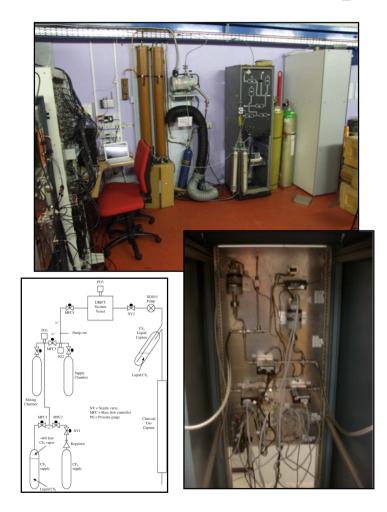
• Measured ionization, gain, drift velocity and diffusion in various CS₂ gas

mixtures with CF₄

DRIFT-IId new set-up with CS₂/CF₄
 (=DRIFT-IIb + gas mix system)

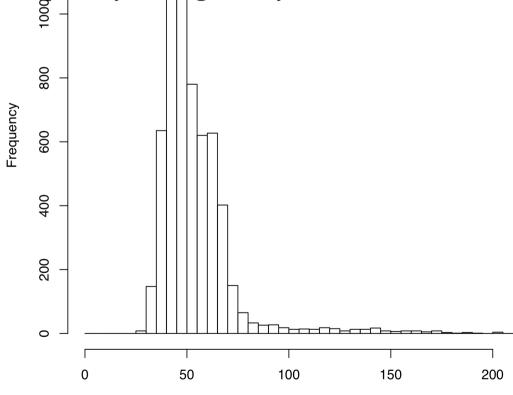


CS₂/CF₄ gain vs mixture



Dark Matter Runs with CS₂/CF₄

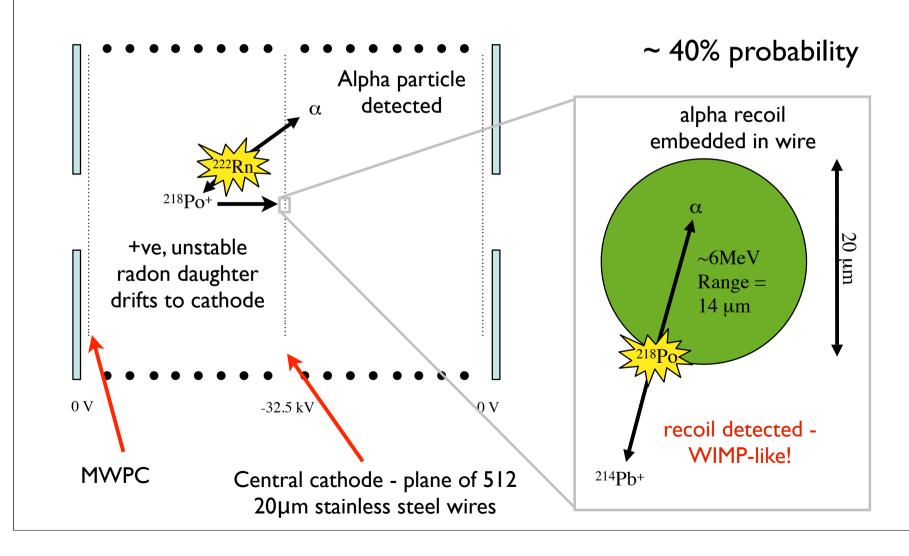
- 47.4 days of live time states collected in Winter 2009/2010
- Target was 30 Torr CS_2^+ + 10 Torr CF_4^- , 139 g of target mass
- F has a large-spin-dependent WIMP-proton cross-section
- 130 events per day through analysis cuts



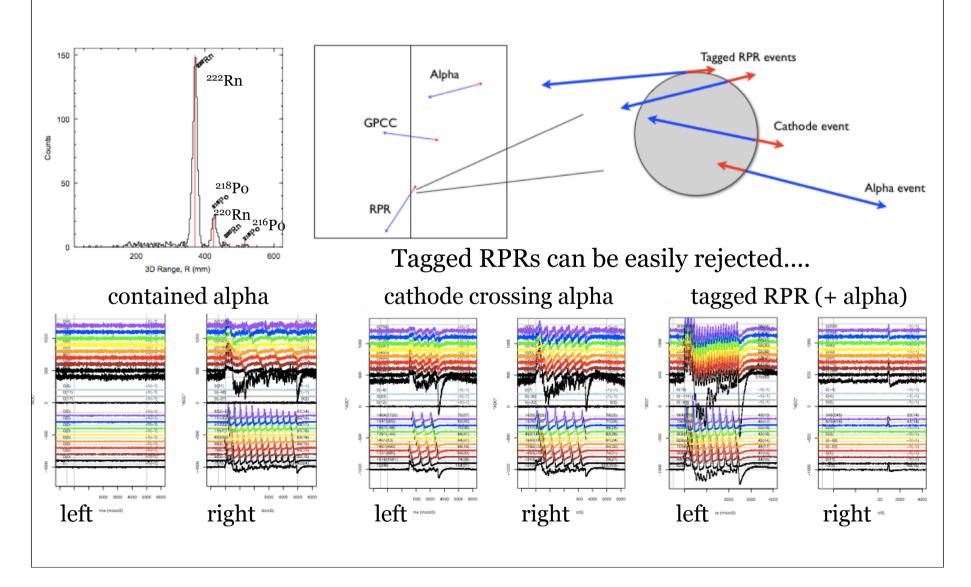
F equivalent recoil energy (keV)

DRIFT II's dominating background

• The main background is from radon progenies (RPRs)



Separating out/reducing backgrounds



RPR Reduction (PART 1)

• Reduce radon producing contaminants inside detector:

Sample	Fill gas	Emanation	Humidity	Raw result	Adjusted result
(Emanating into vacuum)		time (days)	(%)	(Bq/m ³)	(Rn atoms.s ⁻¹)
RG58 coax cables (72m)	Dry N2	12.5	24	9.4 +/- 0.7	0.36 +/- 0.03
Electronics boxes	Dry N2	12	37	1.5 +/- 0.3	0.05 +/- 0.02
Ribbon cables	Dry N2	6.5	23	10.1 +/- 0.7	0.50 +/- 0.04
Electronics & PCBs	Dry N2	10	37	0.3 +/- 0.2	<0.02 *
Single core & thin coax cables	Dry N2	7	19	1.3 +/- 0.3	0.04 +/- 0.02
Field cage parts	Dry N2	7	33.3	0.6 +/- 0.2	<0.03 *
				Total	0.95 +/- 0.5



S. Sadler, S. Paling et al. (Sheffield)

• RPRs still produced from Pb isotopes plated out on cathode. Clean cathode with nitric acid

D. Snowden-Ifft, Oxy, J. Turk, UNM (PhD thesis 2008)

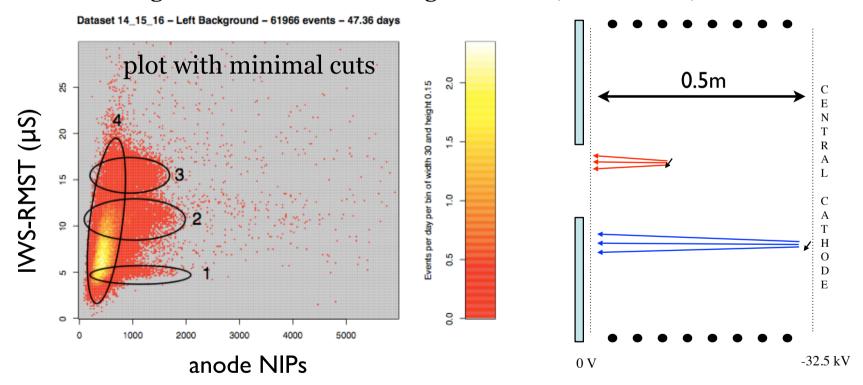
Together, these reduced the RPRs by 96% relative to D-IIa rate

Use pulse z-direction shape shape



Separating out/reducing backgrounds

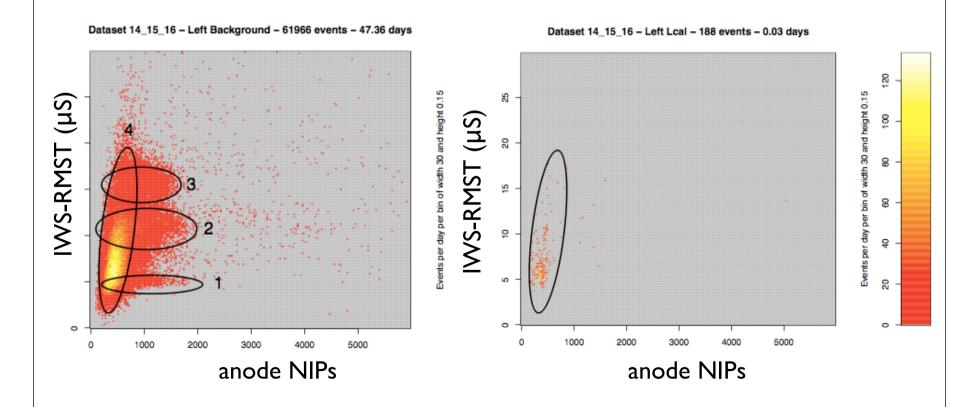
- Neutrons occur throughout the detector volume
- RPRs come from the central cathode or MWPC and suffer different diffusion
 on average cathode RPRs have higher width (IWS-RMST)



Plots with reduced cuts help explain various backgrounds at low recoil energy: (1) MWPC sparks, (2) MWPC RPRs (3) central cathode RPRs (4) betas

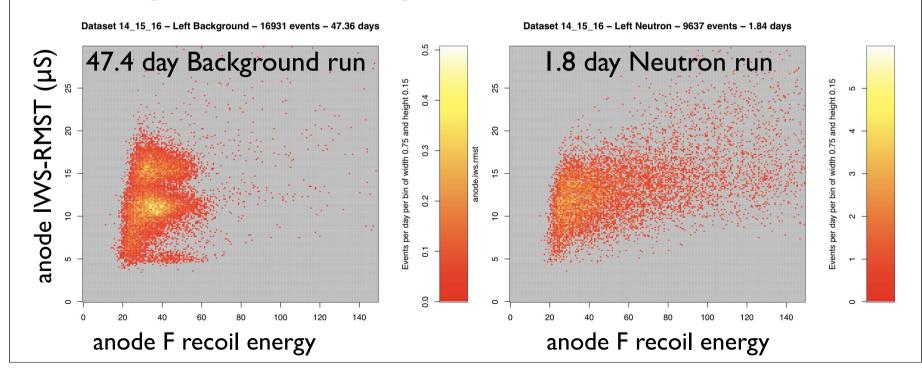
Separating out/reducing backgrounds

"WIMP" analysis of 55Fe reveal some leakage of electron events at very low threshold (~500 NIPs)



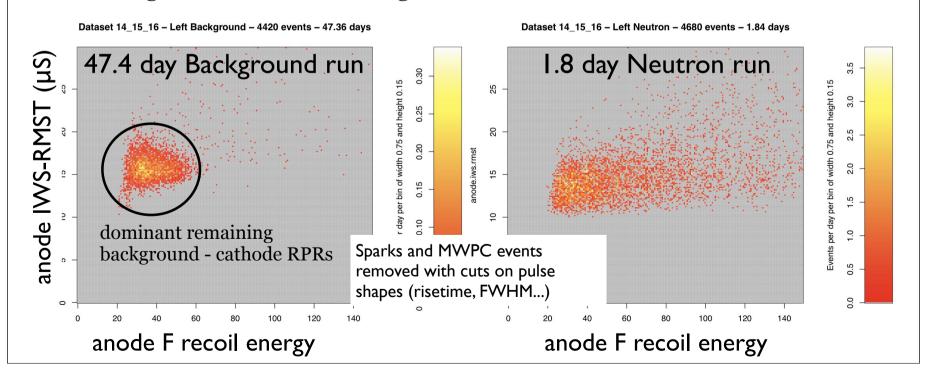
Backgrounds and neutron calibration

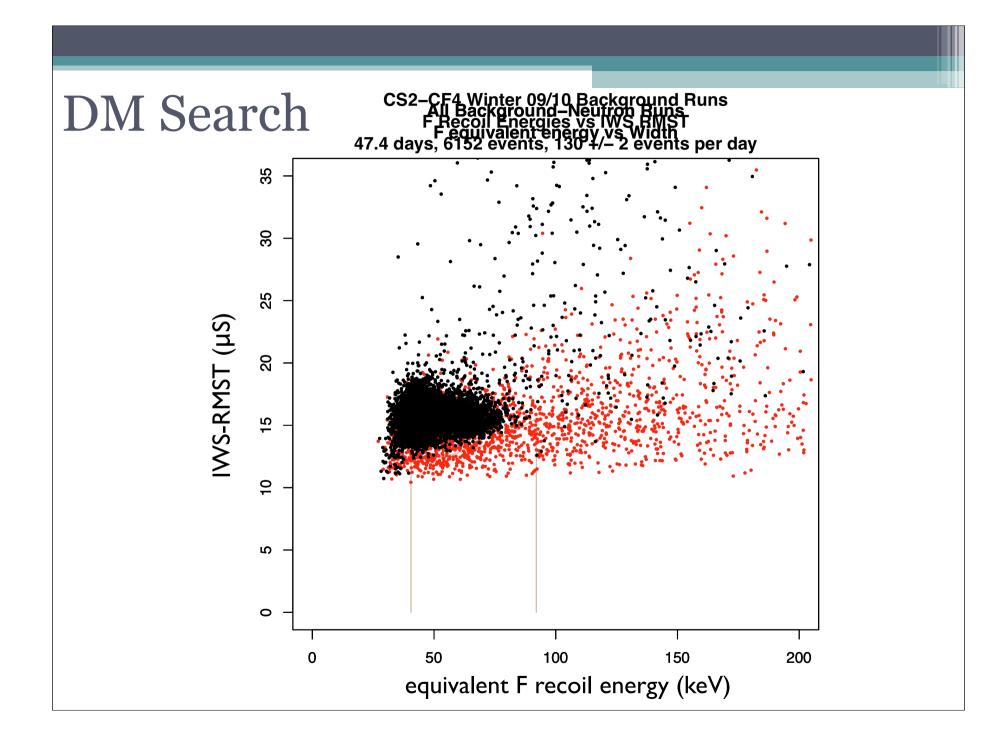
- X-axis equivalent F recoil energy (keV)
- Y-axis IWS-RMST (Induced Waveform Subtracted RMS Time) measure of width of the track in the drift field (Z) dimension
- Three main background populations:
 - Low RMST sparks consistent with shaping time of amplifiers
 - Mid RMST events in the MWPC (RPRs?)
 - High RMST RPRs coming from the central cathode



Backgrounds and neutron calibration

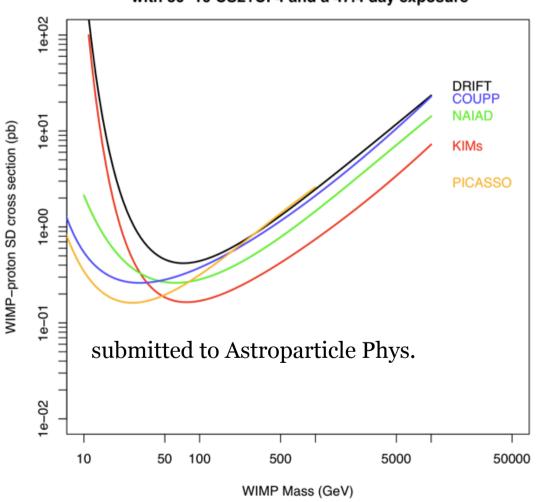
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 - Mid RMST events in the MWPC (RPRs?)
 - High RMST RPRs coming from the central cathode





SD Limit from 47.2 days

SD WIMP-proton Limits with 30-10 CS2+CF4 and a 47.4 day exposure



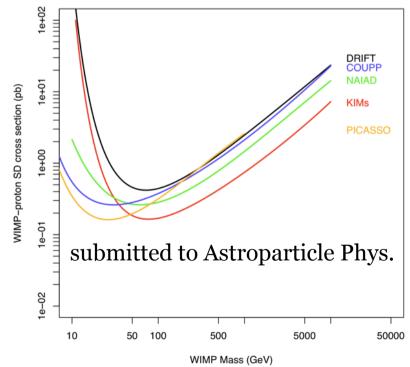
DRIFT: 1.5 kg-days COUPP: 28 kg-days

NAIAD: 12,500 kg-days KIMS: 3,400 kg-days

SD Limit from 47.2 days

- 30 Torr 10 Torr CS2-CF4, 47.2 days background data $= 1.5 \text{kg-days} (^{19}\text{F})$
- MC simulation calibrated by neutron data
- No compromise on directional sensitivity
- Signal region chosen for zero events (unblind analysis)
- Further 53 days data on disk for a full blind analysis

SD WIMP-proton Limits with 30-10 CS2+CF4 and a 47.4 day exposure



Min. SD limits. directional detectors:

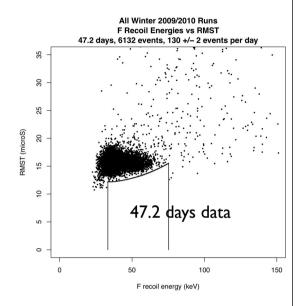
DRIFT: 1.8 pb

NEWAGE: 5400 pb DM-TPC: 2400 pb

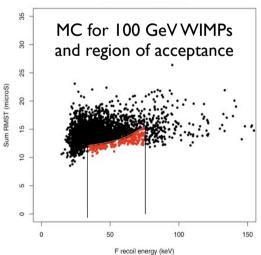
DRIFT: 1.5 kg-days

COUPP: 28 kg-days

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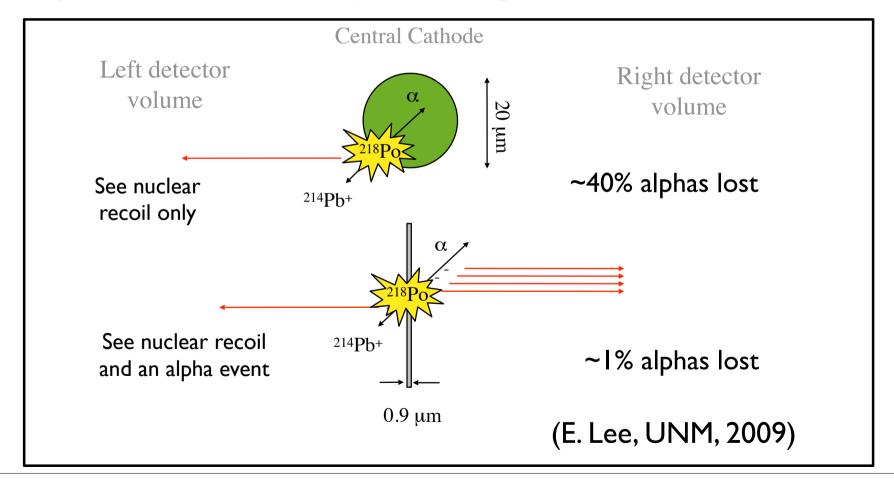
drift2d-20100313-01-0003-cgmc Nips vs RMST 228/2742/10000 events



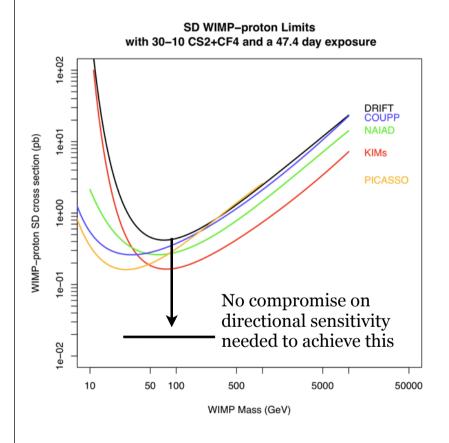
RPR Reduction (PART 2)

• Make central cathode transparent to alphas

A cathode highly transparent to α 's from RPRs will provide a tag to veto events Change from 20 μ m wire to 0.9 μ m thin film predicts x40 reduction in RPR



Thin cathode sensitivity Prediction



- ×40 reduction in RPR background expected
- 0.02pb limit projected assuming RPRs have same distribution
- ~2000 days live time required to achieve this
- DRIFT II is then volume limited not background limited

DRIFT IId 0.9µm cathode

Use of multi-panel 0.9µm thick DRIFT cathode

cathode tested at full voltage (32.5kV)

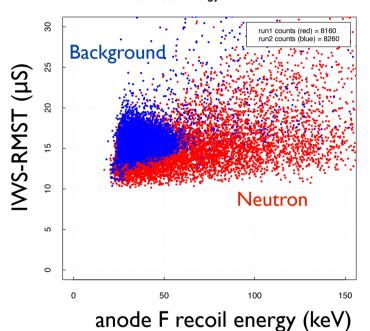


Running stably since installation ~65 days of live-time data collected and counting...

20 μm wire cathode

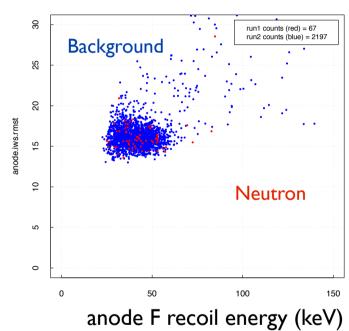
Background events 174 events/day

. anode.F.recoil.energy vs anode.iws.rmst



Tagged RPRs 47 events/day

. anode.F.recoil.energy vs anode.iws.rmst



47.2 days data as used to set limit

0.3 tagged RPRs per background event

0.9 μm film cathode (new data)

Background events 14.7 events/day

Tagged RPRs 63 events/day

. anode.F.recoil.energy vs anode.iws.rmst

. anode.F.recoil.energy vs anode.iws.rmst

30 **Background**

WS-RMST (µS)

anode F recoil energy (keV)

run1 counts (red) = 91 Background 8 Neutron

anode F recoil energy (keV)

11.6 days - 53 days on disk ready for full blind analysis 4.3 tagged RPRs per background event = x 14 improvement

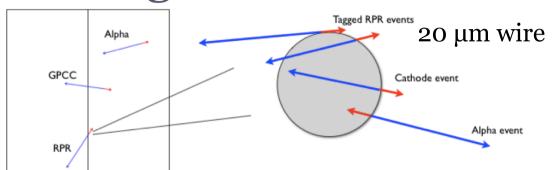
Other smaller backgrounds are now revealed that account for the lower difference seen compared to expectation of ~ x40

These are likely Low Energy Alphas (LEAs) and betas

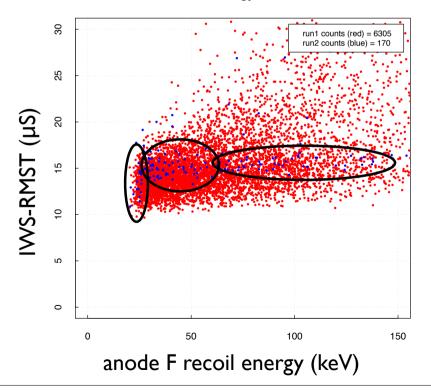
Neutron

Also shifts in the IWS-RMST distributions

Thin film backgrounds



. anode.F.recoil.energy vs anode.iws.rmst



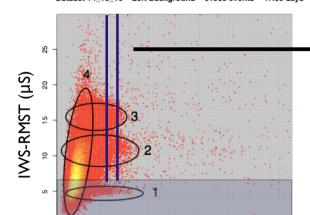
0.9 μm thin film

Background populations and rejection:

- (i) RPR (standard)
- (ii) double RPR RPR plus low energy alpha
- (iii) low energy alpha(straggling)
- (iv) betas

Backgrounds RMST and rates vs. time

Lots of information on backgrounds by studying time changes over 1500 days



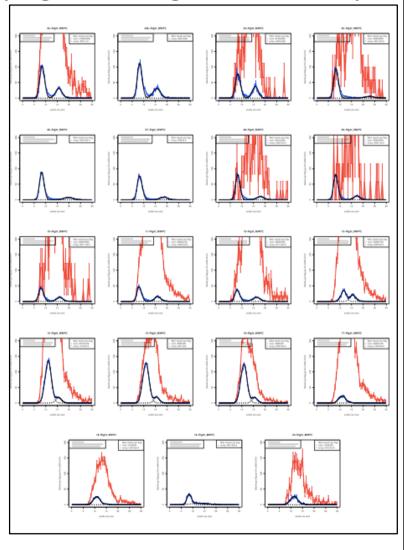
look at energy slice IWS-RMST distributions vs. time

Population changes in MWPC and cathode RPR events during evolution of the detector through radon reduction upgrades:

Position and rates changed by:

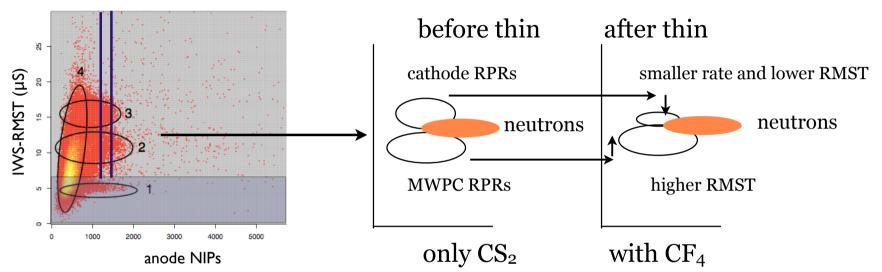
anode NIPs

- -cleaning of MWPCs
- -cleaning of central cathode
- -addition of CF₄
- -gas flow
- -change to thin cathode



Backgrounds RMST and rates vs. time

Lots of information on backgrounds by studying time changes over 1500 days



Conclusions:

The thin film successfully reduces the cathode RPRs. However, RMST factor reduces. This is most likely a geometric effect that untagged thin film RPRs are ejected more parallel to the film and so have smaller RMST



Meanwhile on change to CF₄ the MWPC RPRs move up in RMST. This is most likely due to a longer free electron capture time cf pure CS₂

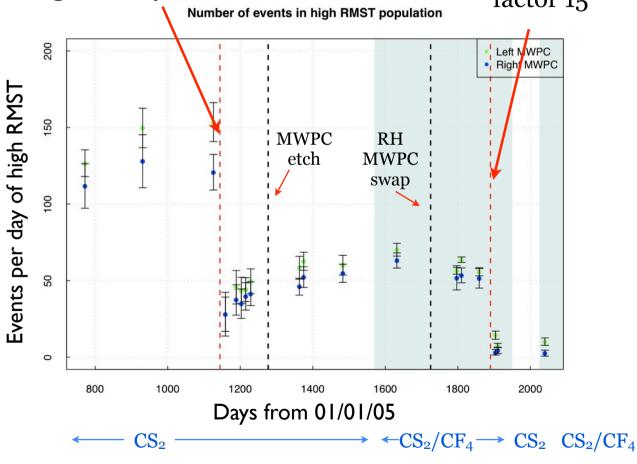
The result is that the signal (neutron) region gets squeezed.

Backgrounds rates vs. time

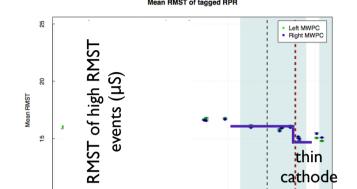
For instance cathode RPR region (higher IWS-RMST) vs. time

nitric acid etch of cathode reduced background by 5

thin film cathode reduces by further factor 15



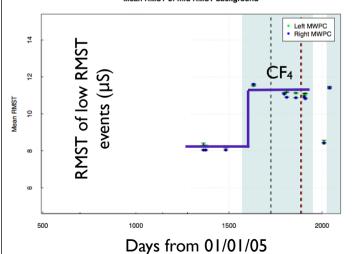
Backgrounds RMST and rates vs. time



high RMST population

Days from 01/01/05

Mean RMST of mid RMST background

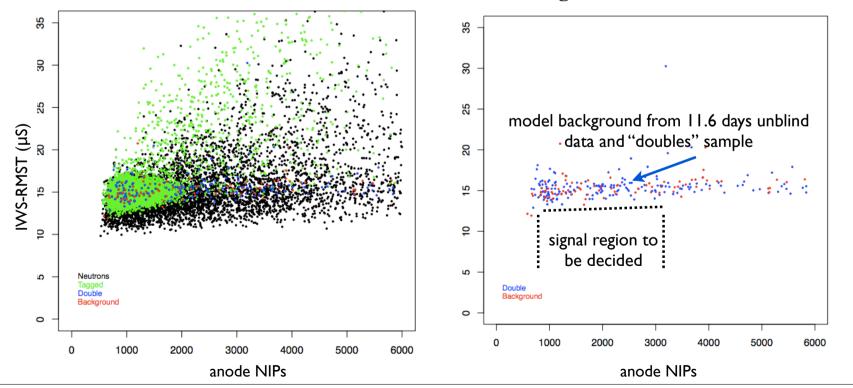


low RMST population

Towards new blind analysis

By relaxing cuts to see tagged RPRs and "doubles" we can produce "model" background events in the data that can be used to set the data box region for a full blind analysis

11.6 days unblinded and 53 days on disk ready for full blind analysis
all runs unblind background and double recoil runs



Next with DRIFT II, towards DRIFT III

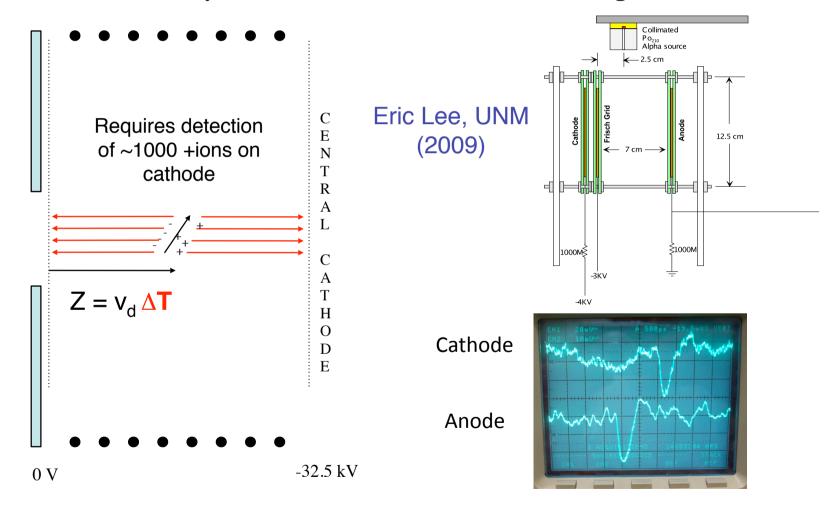
- (1) Main thrust is RPR elimination:
 - (a) more reduction of intrinsic radon/RPR contamination
 - (b) improved PSD/position analysis and cuts
 - (c) further improvement to <u>alpha-transparent cathode</u>
 - (d) full z-fiducialisation via +ve ion

- (2) Upgrade/streamlined electronics and gas system
- (3) DRIFT III scale-up design 24 m³

in 4 m³ segments

Z-fiducialisation (E. Lee, D. Loomba et al., UNM)

Z-fiducialisation by +ve ion detection demonstrated to give ΔT measurement



Z-fiducialisation scale-up

New mini-DRIFT test vessel at UNM to optimise z-fiducialisation design includes:

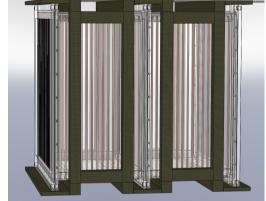
- Vibration suppression via acoustic shield and elastic suspension
- Lexan open ended box with Cu tape on inside for field cage with minimal HV coronal discharge and anti-vibration support
- I-beam design to maximise mechanical strength. Minimal dielectric (acrylic, polycarbonate) inner frame needed to maximize the open aperture area
- Then fit to DRIFT IIe



Inner surfaces of frame can mount circuit boards

Kevlar-epoxy laminate Structural frame.



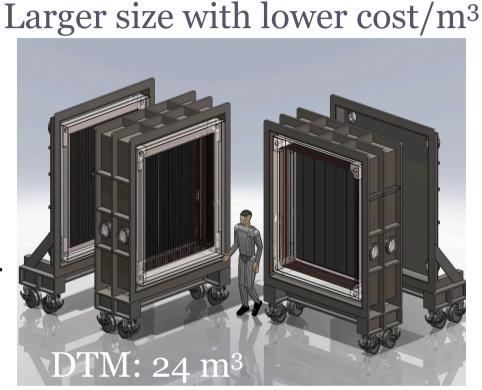




Bottom piece (MWPC frame) resting on bottom of vacuum vessel may be partly aluminum for internal electronics heat transfer.

DRIFT III Module (DTM) Concept

- Modular design to allow approach to ton-scale
- 4 kg target 24 m³
- One DTM well suited to Boulby
- Large number fits 30m x 150m DUSEL "Standard Lab Module" or new tunnel excavation at Boulby for 1 ton
 - advantages of no cryogenics
 - multiplexed electronics
 - neutron shielding



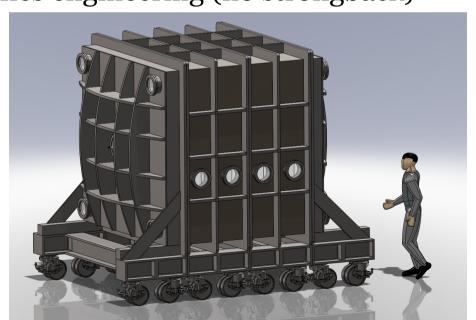
cost split evenly:

- vacuum vessel
- electronics
- gas system

NB: current cost of complete DRIFT II module (with shielding) ~ \$80K Extrapolation gives ~\$250K for DTM (with shielding)

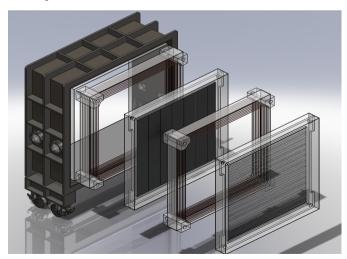
DTM Baseline Design Subject to change

- 4kg fiducial mass, CS₂ -ve ion plus CF₄ (different target mixes)
- Thin low RPR central cathode (1 μm), partial segmentation
- Nitric acid process cleaning and radon emanation tests
- +ve ion detection for Z-fiducialisation
- 2 x 2m single plane anode with alternate grid wires, 1mm pitch reduced tension simplifies engineering (no strongback)
- Head-Tail sensitivity
- 2D readout but with 3D side veto using resistive wires
- neutron shielding
- No gamma shielding needed

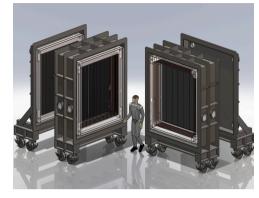


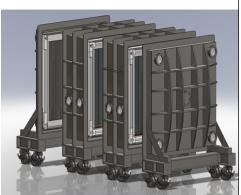
DTM Vessel

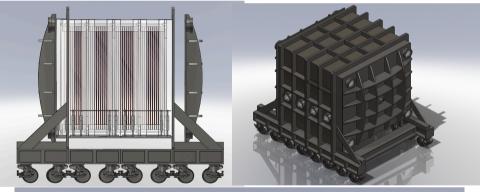
- One DTM vessel composed of 6 segments footprint ~6 m by 3 m.
- Each detector segment observes 4 m³ of gas or, at 40 Torr of CS₂, 0.67 kg of target mass
- 250 of these 4 kg modules gives 1 ton and would fit into a standard DUSEL module or 500m tunnel at Boulby

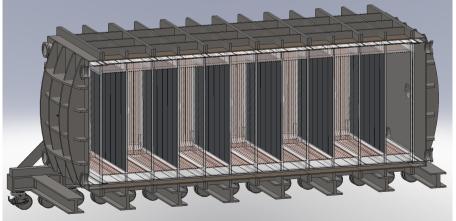


Preference for CH-based material









DRIFT III Module Readout

Sense plane

• Transparent readout plane to sense two sides (eliminates the mechanical

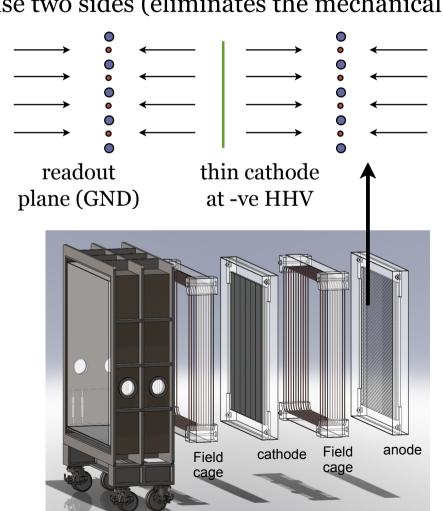
support "strong back")

 20 μm diameter stainless steel wires on a 2 mm pitch

• X-wires, Y-resistive wires

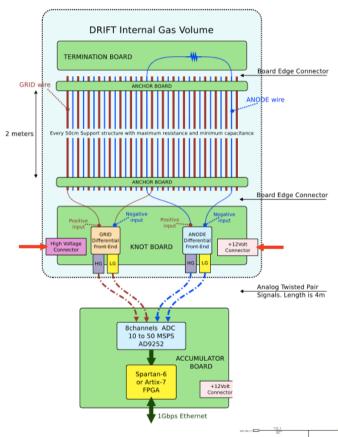
Cathode

- 70 kV with well-engineered field cage and high-voltage system; diffusion (reduced by 40% c.f. DRIFT II)
- +ve ion detection segmented to reduce the input capacitance for Z fiducialisation
- Orientation perpendicular to anode wires to give more y dimension information



DRIFT III Module Electronics

(J. Harton et al., CSU)



Anchor Board

Glued with epoxy to support frame.

Wires are soldered directly to the board.

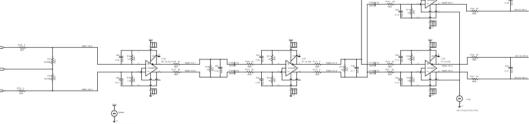
Physically separates anodes from grids.

Termination Board

Plugs into Anchor Board through board edge connector.

Completes one side of the anode loop for resistive readout.

Every 9th anode is connected. Only one loop is shown for picture clarity.



Anode Front End

More than 1GhZ Differential opamp: ADA4930. High Supply transient rejection rate, fully balanced input. Low and High gain channels (ratio is 200).

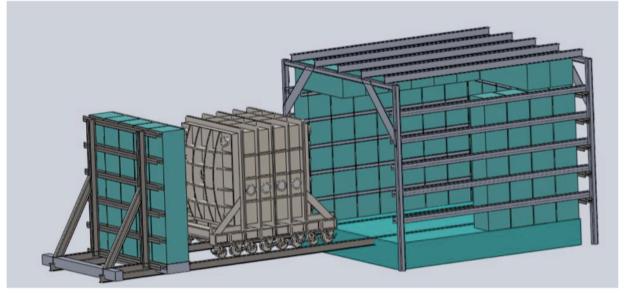
DTM Shielding

Simple and cheap poly pellet or water neutron shielding





DRIFT II

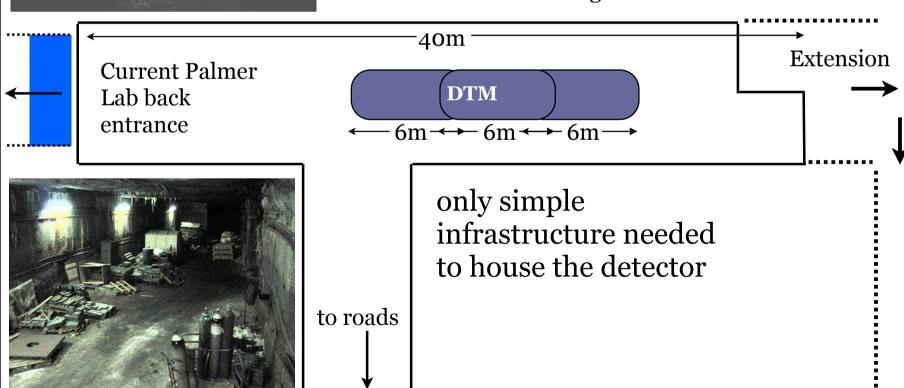


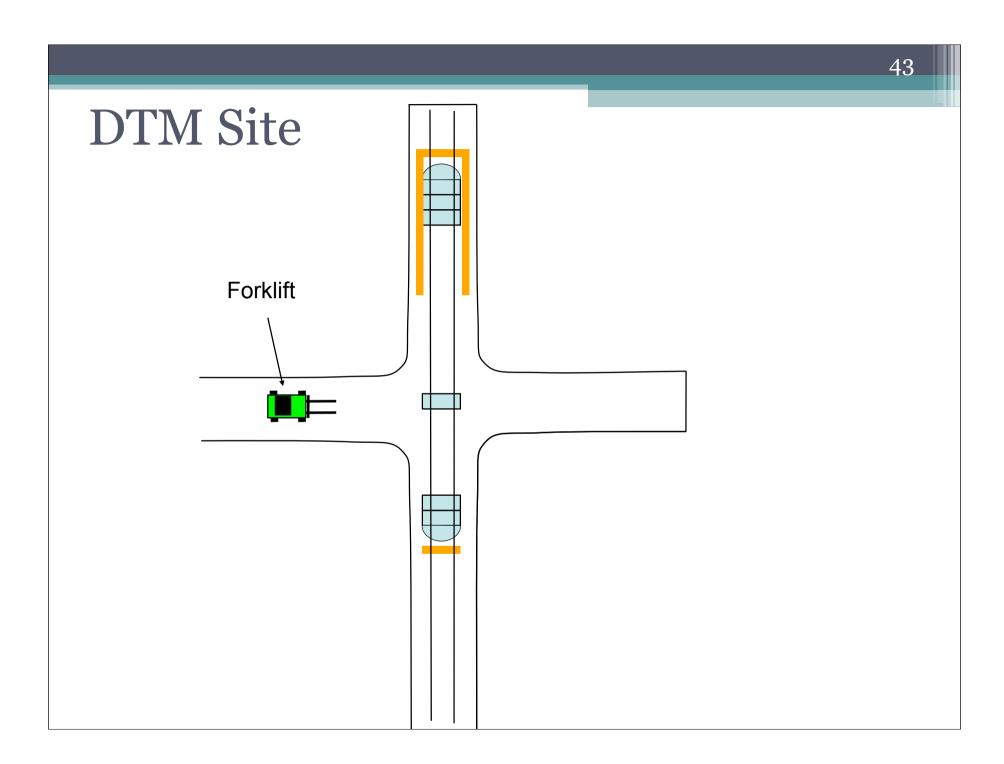
DTM



DRIFT III Module Site

- Site for first DRIFT III module identified at Boulby
- CPL willing to excavate for DTM at "no charge" scale drawing

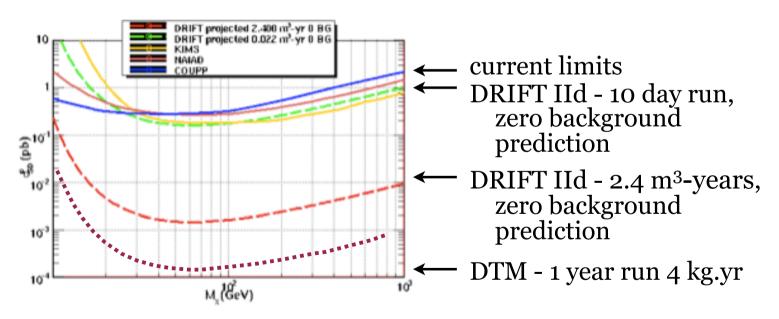




DRIFT III Module (DTM) Sensitivity with directional capability SD Prediction

- Projected limits for a DTM (DRIFT III Module) (24 m³)
- With no compromise on directional sensitivity

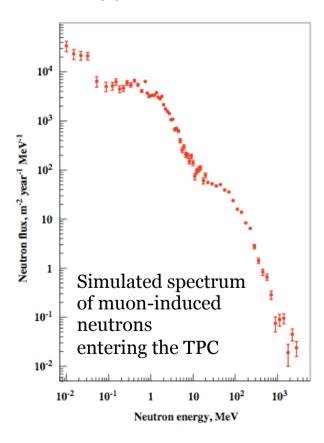
Expected WIMP-proton spin dependent sensitivity



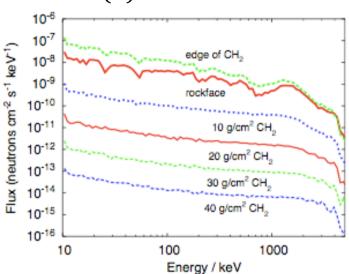
Neutron Summary DRIFT II studies

e.g. see M.J. Carson et al NIM A 546 (2005) 509-522

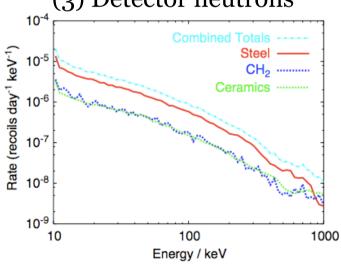
(1) Muon neutrons



(2) Rock neutrons



(3) Detector neutrons



Neutron Summary

Assumptions

Includes 40 g cm⁻² CH₂ shielding against rock neutrons (estimates)

Result (prelim estimates for DTM) see M.J. Carson et al NIM A 546 (2005) 509–522

Estimated neutron backgrounds per year at 10-50 keV recoil energies

rson et al NIM A 546 (2005) 509–522	kg	Rock	Muons	Detector	Total
DRIFT II	0.167	0.01	0.12	0.06	0.19
DTM (as multiple DRIFT IIs)	4.00	0.24	2.88	1.56	4.68
DTM using steel, no muon veto	4.00	0.20	2.00	1.50	3.70
DTM acrylic, no muon veto	4.00	0.20	<1.00	<1.00	<0.4

Conclusion for single DTM (prelim):

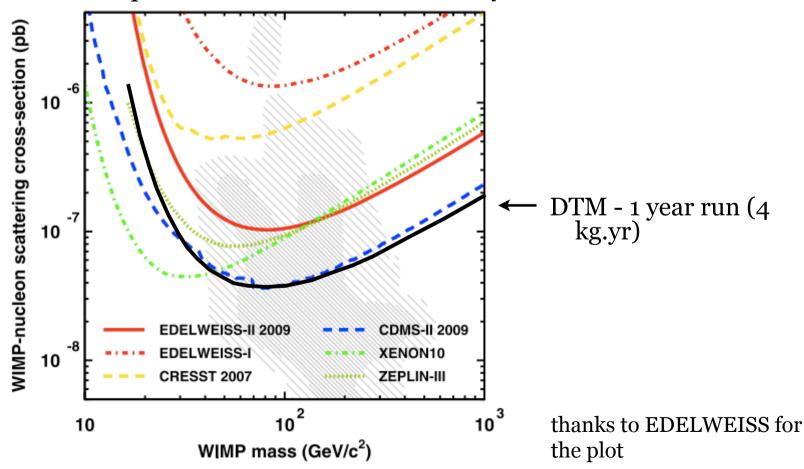
- Requires 40 gcm⁻² CH neutron shielding (like DRIFT II)
- Steel construction just about alright optimization, selection, internal CH?
- No need for muon active veto at Boulby for single module

DTM Sensitivity

SI Prediction

with directional capability

- Projected limits for one DTM (4kg)
- With no compromise on directional sensitivity



Conclusion - a 24m³ directional detector

- New DRIFT IId WIMP-proton limits
- Blind analysis of further data with thin cathode soon
- Development from DRIFT II to DTM (x24) is now not a major technical leap; the main challenges are:
 - Vessel design (reduction of muon and detector neutrons)
 - Full implementation of z-fiducialisation (RPR reduction)
 - Gas recycling and handling underground
- CCDs, Micro-pix, Micromegas alternative readouts also possible

