

Invisible Higgs and dark matter

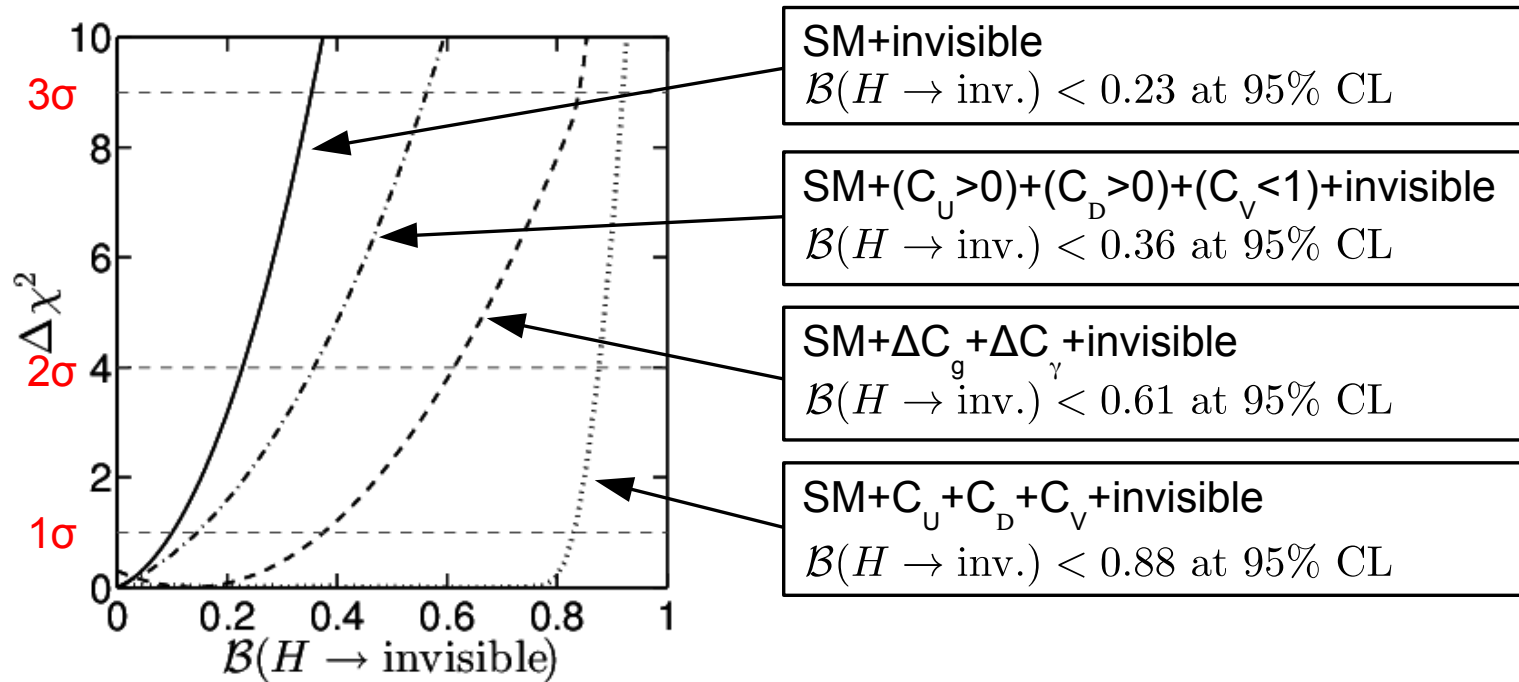
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LAPTH, Annecy-le-Vieux

LPSC, March 2013

- Invisible Higgs decay constrain DM models with $M_{\text{DM}} < m_h/2$
- Strong link with direct detection
- Model independent: Scalar/fermion/vector
- Models : MSSM/NMSSM

Constraints on $\text{Br}(H_{\text{inv}})$

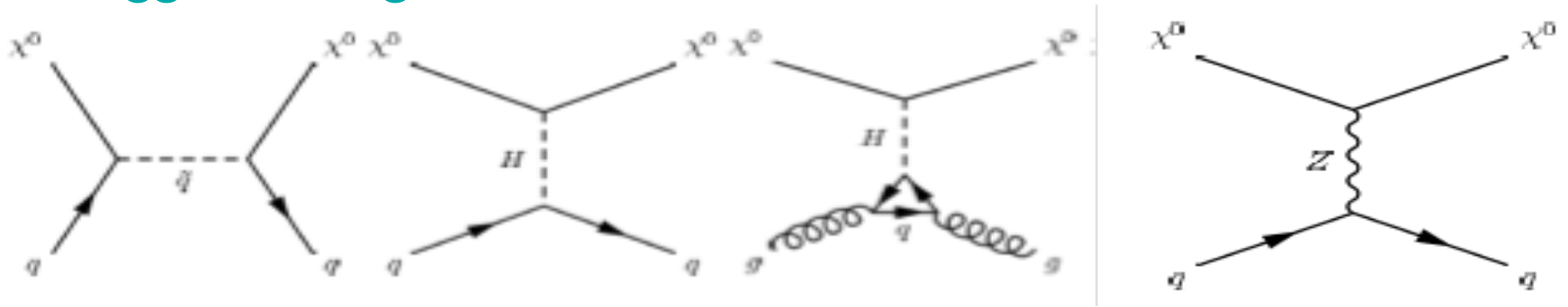
- From general fits (GB, Dumont, Elwanger, Gunion, Kraml, 1302.5694)



- From direct searches (see Kirtimaan's)

Direct detection

- Higgs exchange often dominates



For Dirac fermions Z exchange contributes to SI and SD

- Both Invisible branching of Higgs and spin-independent cross-section depend on $h\chi\chi$ coupling squared

Invisible Higgs and direct detection

- For light DM particle, relation between invisible width and direct detection

$$\sigma_{\text{SI}} = \eta \mu_r^2 m_p^2 \frac{g^2}{M_W^2} \Gamma_{\text{inv}} \left[C_U(f_u^N + f_c^N + f_t^N) + C_D(f_d^N + f_s^N + f_b^N) + \frac{\Delta C_g}{\widehat{C}_g} f_g^N \right]^2$$

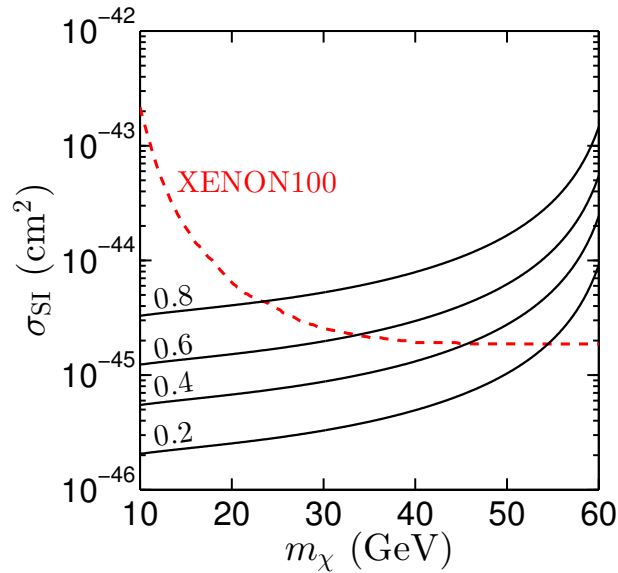
- Majorana fermion

$$\eta = 4 / (m_H^5 \beta^3)$$

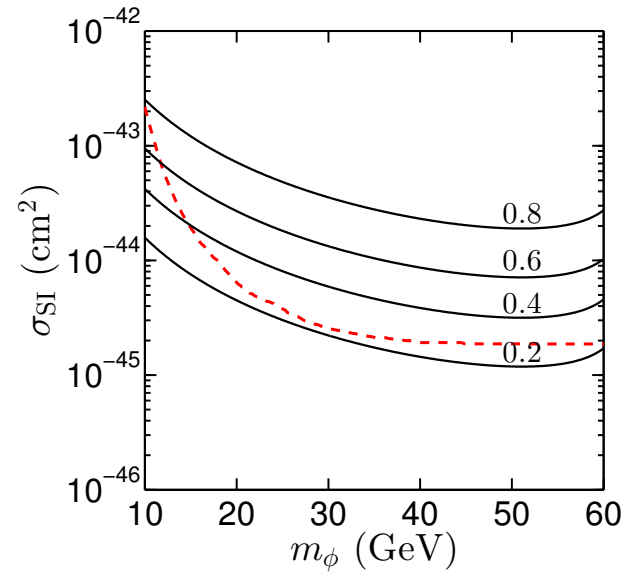
- Real scalar

$$\eta = 2 / (m_H^3 m_\phi^2 \beta)$$

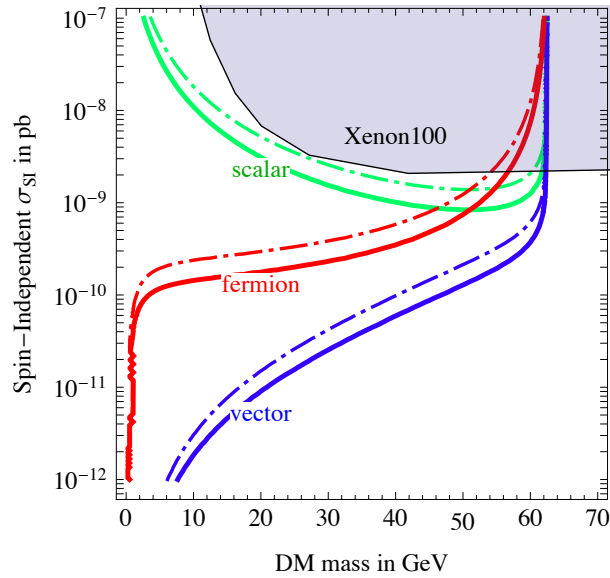
Fermion DM



Scalar DM



- Fermions : for light DM - Higgs invisible much stronger constraint than DD
- Scalar - DD strong constraint on Br_{inv} except for very small masses



$Br_{inv}=0.19-0.28$
 Giardino et al, arXiv 1303.3570

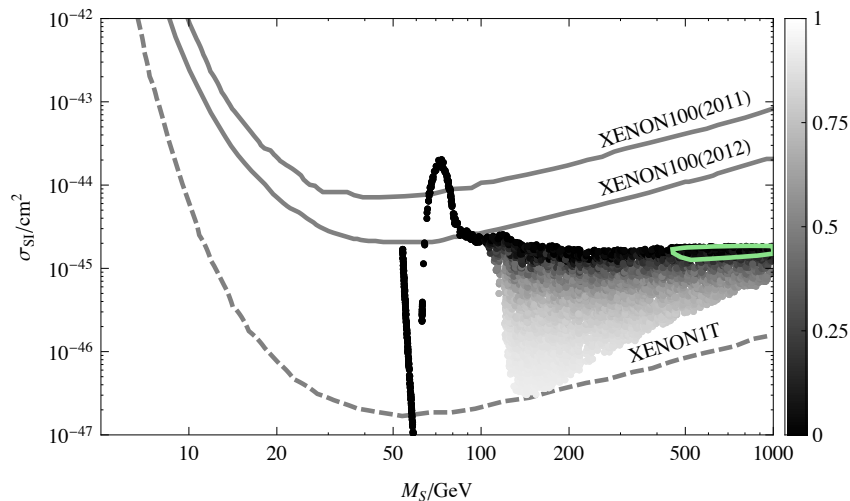
See also Djouadi et al, 1205.3169

$$\sigma_{\text{SI}} = \eta \mu_r^2 m_p^2 \frac{g^2}{M_W^2} \Gamma_{\text{inv}} \left[C_U(f_u^N + f_c^N + f_t^N) + C_D(f_d^N + f_s^N + f_b^N) + \frac{\Delta C_g}{\widehat{C}_g} f_g^N \right]^2$$

- Non-standard Higgs couplings
 - potential cancellation ($C_U < 0$, $C_D > 0$) - at first sight relax direct detection bound
 - however need to generate fermion mass - other Higgs should couple to fermion - give contribution to direct detection
- Direct detection can involve other particles
 - coloured particles (e.g. squarks in SUSY) contribution suppressed by mass scale
 - other Higgses

Specific models

- Take into account various constraints, in particular relic density
 - light DM need efficient annihilation mechanism
 - through Z, Higgs, new particles
 - Example: Scalar dark matter



SM+scalar singlet, Z_3
 $Br_{inv} < 0.4$
GB et al, 1211.1014

See also Djouadi et al ,
1112.3299

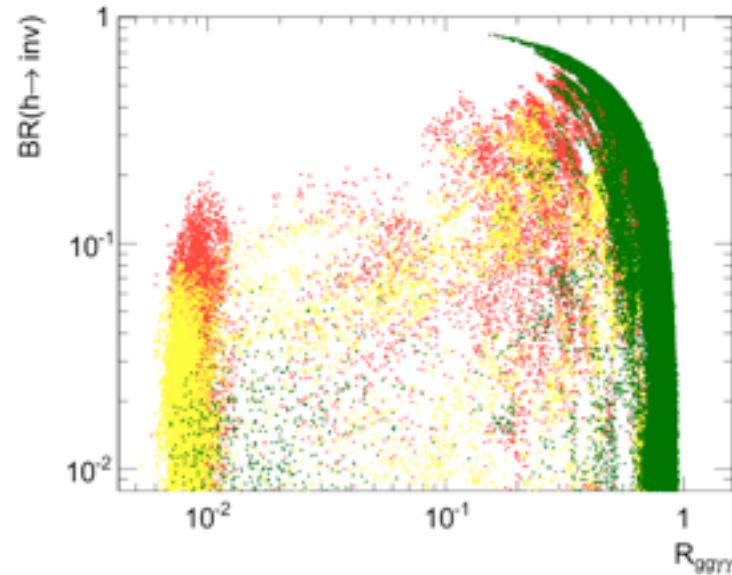
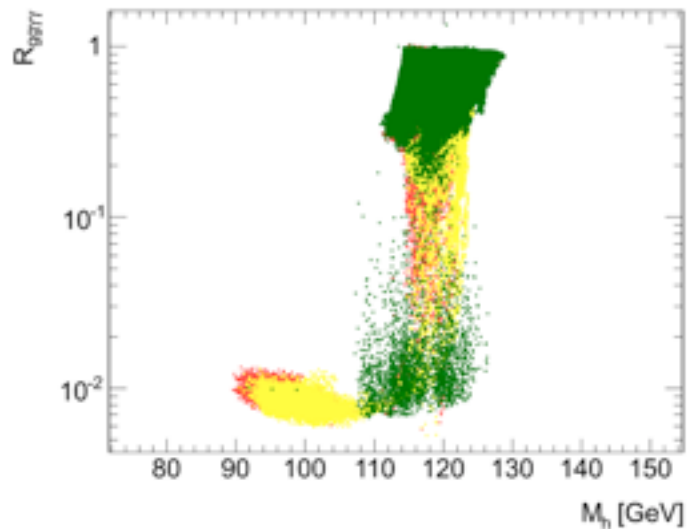
What about MSSM?

- Higgs coupling to neutralinos

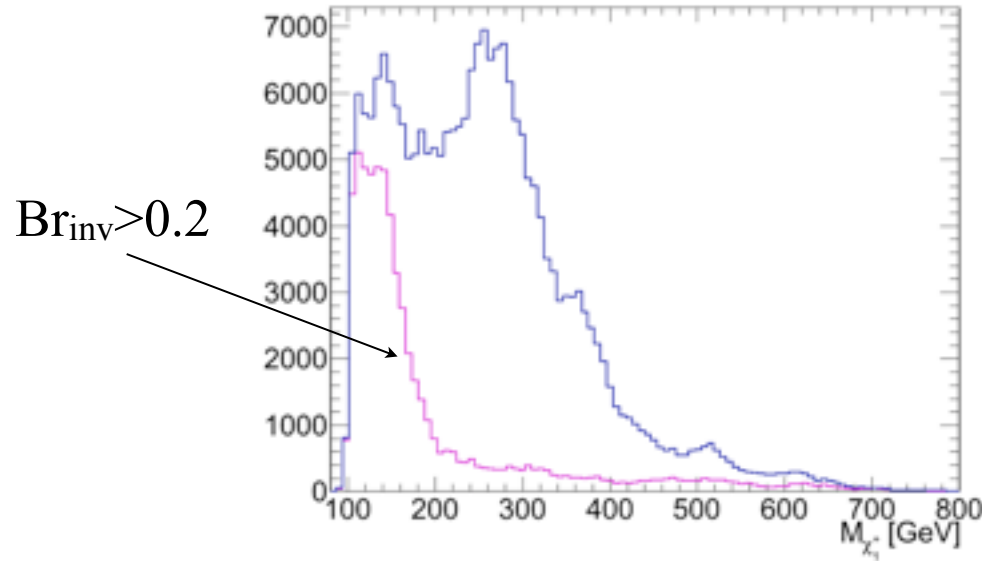
$$C_{h\tilde{\chi}_1^0\tilde{\chi}_1^0} = (O_{12}^N - \tan\theta_W O_{11}^N)(\sin\alpha O_{13}^N + \cos\alpha O_{14}^N)$$

- Significant coupling for mixed bino/Higgsino LSP - μ small
- Light LSP must have sufficient annihilation for $\Omega h^2 \sim 0.1$
 - through Higgs exchange --- mixed bino-higgsino
 - Z exchange -- higgsino component ($O_{13}^2 - O_{14}^2$)
 - into fermions through stau exchange - bino

Results before LHC-Higgs discovery



- Scan in MSSM-11 (Albornoz Vasquez et al, 1112.2200)
- Taking into account $m_h \sim 125$ + Higgs couplings, + Xenon limits + relic density, what are predictions for Br_{inv} ?

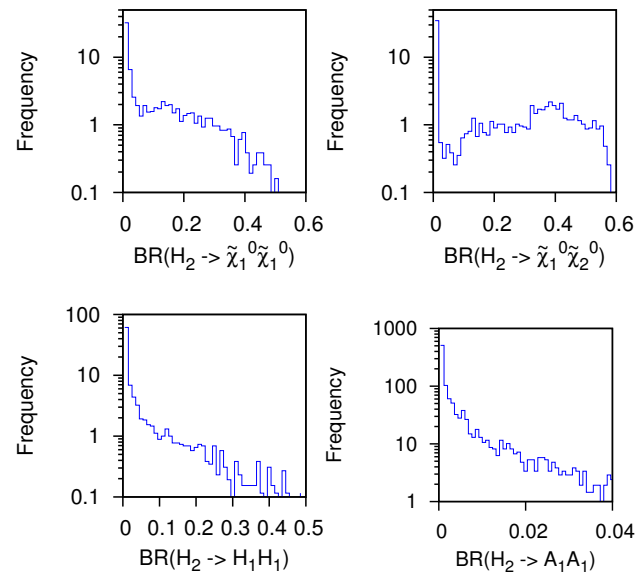


- Large Br_{inv} - associate with light chargino (small μ)
 - see also Dreiner, Kim, Lebedev, 1206.3096

NMSSM

- Additional light Higgses can contribute to annihilation of light neutralino, direct detection
- Both ‘invisible and undetected’ decays
- Destroy simple relation Br_{inv} --direct detection
- Important to search new Higgs decay channels AND H_{inv}

NMSSM with 14
parameters
with light neutralino
and DM constraints
arXiv:1203.3446



$122 < m_h < 128$

Summary

- Constraints on invisible Higgs - limits on light DM
- Correlation invisible Higgs direct detection
- Remains to be seen : what is still allowed in specific models such as MSSM, NMSSM ...