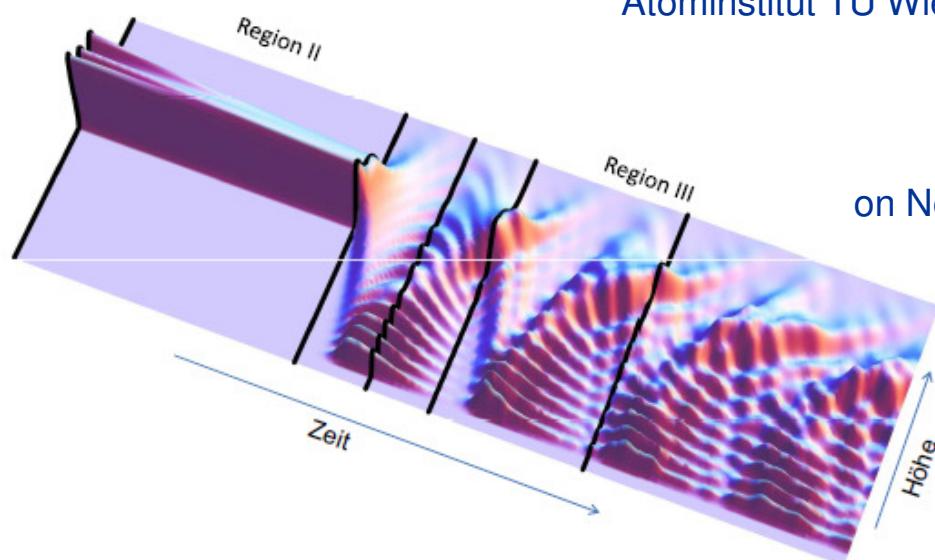


Gravity Resonance Spectroscopy within the qBounce experiment



Tobias Jenke

Atominstitut TU Wien, Vienna/Austria

NPPatLPS 2013
ESS Science Symposium
on Neutron Particle Physics at Long Pulse Spallation Sources
Grenoble
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Gravity \xleftarrow{UCN} Quantum Mechanics

Schrödinger equation:
$$\left(-\frac{\hbar^2}{2m} \frac{\partial^2}{\partial z^2} + mgz \right) \varphi_n(z) = E_n \varphi_n(z)$$

boundary conditions: $\varphi_n(0) = 0$ $(\varphi_n(l) = 0)$

scale factors:

$$z_0 = \sqrt[3]{\frac{\hbar^2}{2m_n^2 g}} \approx 5,88 \mu\text{m}$$

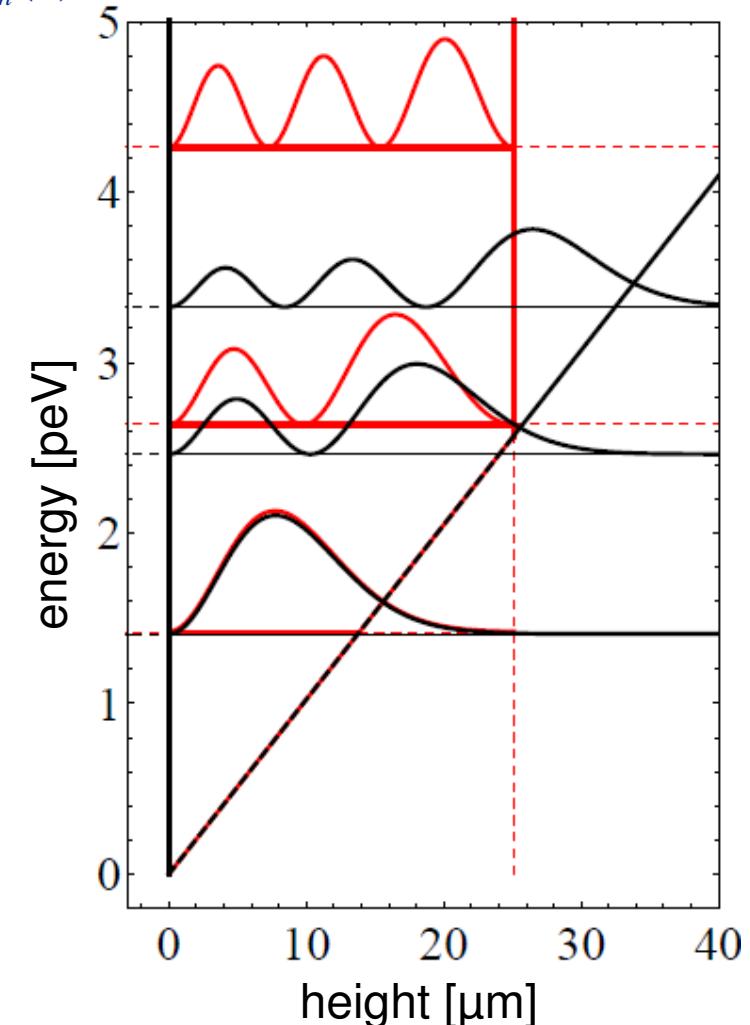
$$E_0 = m_n g z_0 \approx 0,6 \text{ peV}$$

Summary: $t_0 = \frac{\hbar}{E_0} \approx 1,1 \text{ ms}$

- **gravitationally** bound quantum states
- wave functions: Airy-functions
- discrete, **non-eqidistant** energy eigenvalues
- ground state energy: 1,4 peV

Experimental Status:

- 1978: Proposal: *Quantum Bouncer with UCN*
V.I. Lushchikov, A. Frank, JETP Letters **28**, 559ff (1978)
- 2002: *Demonstration of UCN gravity states*
V.V. Nesvizhevsky et. al., Nature **415**, 297ff (2002)



Outline

Gravity Resonance Spectroscopy within the qBounce experiment

- Motivation: Scientific Case
- Gravity Resonance Spectroscopy
 - The key technologies
 - State of the Art (2010-2011)
- GRS at ESS?
 - Statistics
 - Systematics
- Outlook: Requirements for the ESS

UCNs and short-ranged gravity

Gravity at short distances (μm)

$$V(r) = -G \frac{mM}{r} \left(1 - \alpha e^{-\frac{r}{\lambda}} \right)$$

- Large Extradimensions (ADD)
 $10^6 \leq \alpha \leq 10^9$ ($\lambda = 1\mu\text{m}$)
- Cosmological Constant (Burgess et. al.)
 $\alpha \leq 10^6$ ($\lambda = 5\mu\text{m}$)
- Hypothetical short-ranged spin-dependant interactions („axion window“)
 $0,2\mu\text{m} \leq \lambda \leq 2\text{cm}$ ($10\mu\text{eV}..1\text{eV}$)
- Hypothetical scalar fields
(i.e. chameleon fields)

quantum mechanics

- quantum interference
- highly sensitive methods:
resonance spectroscopy

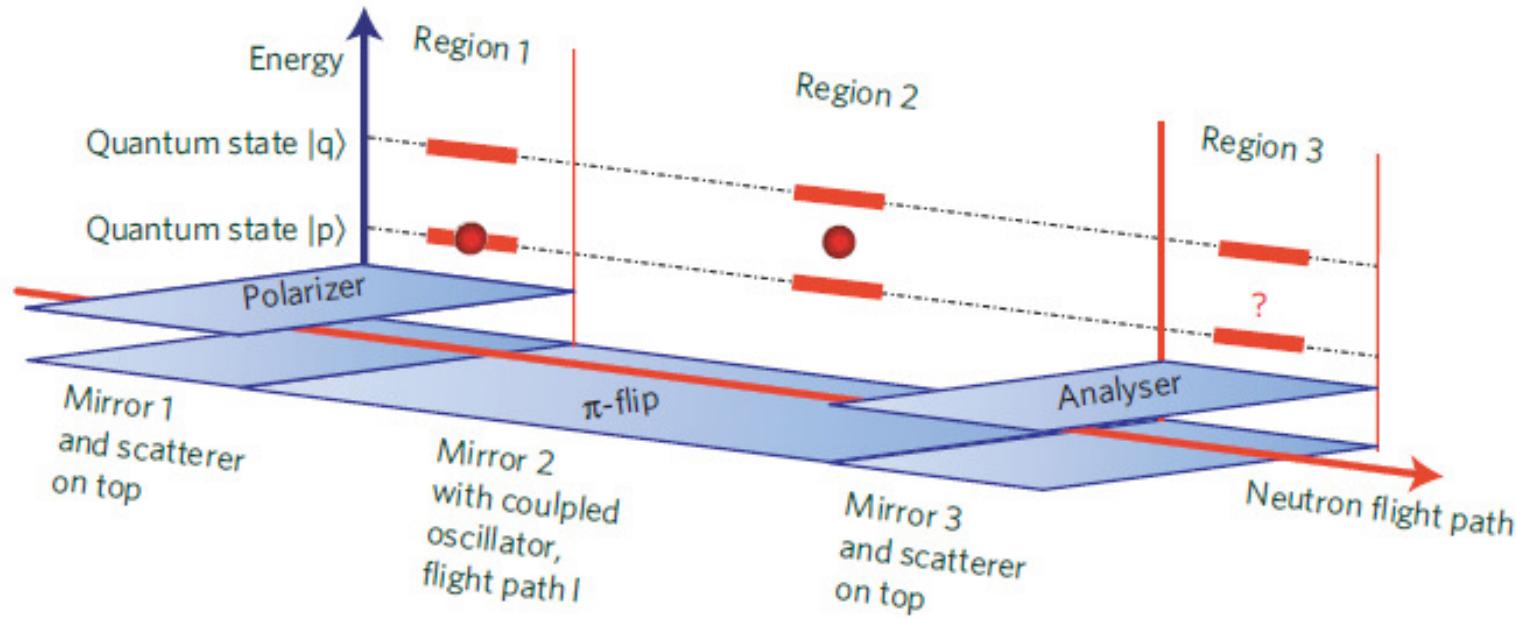
$$E = h\nu$$

UCN

Other Applications:

- Test of neutrons neutrality
- ...

Gravity Resonance Spectroscopy



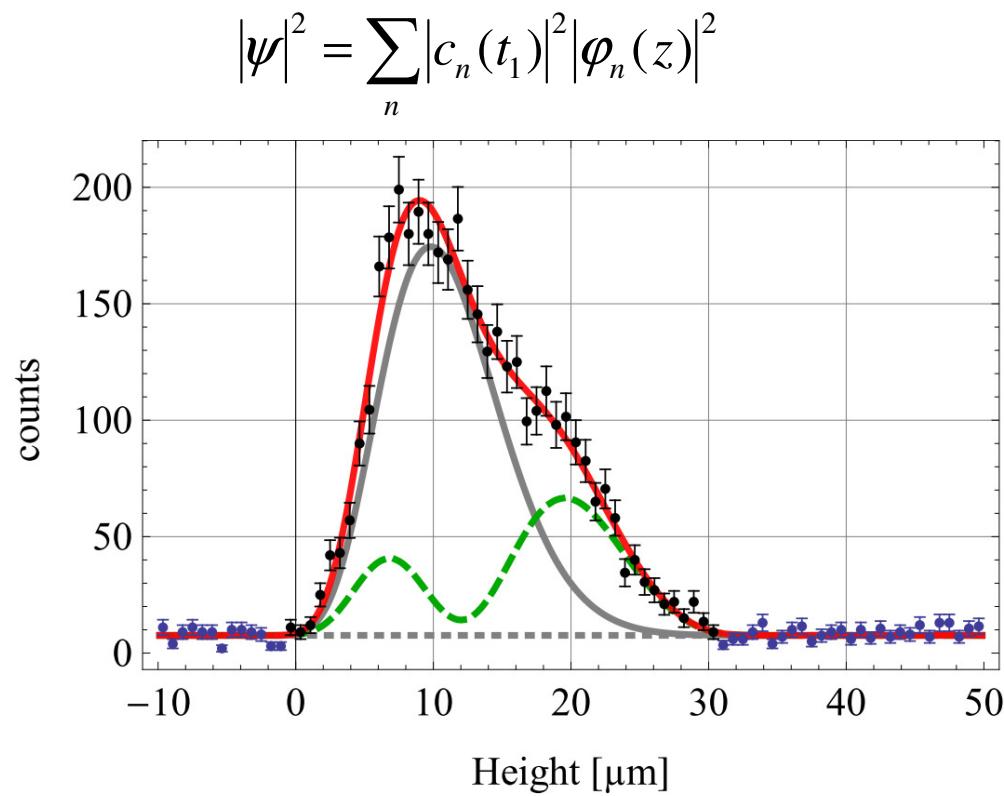
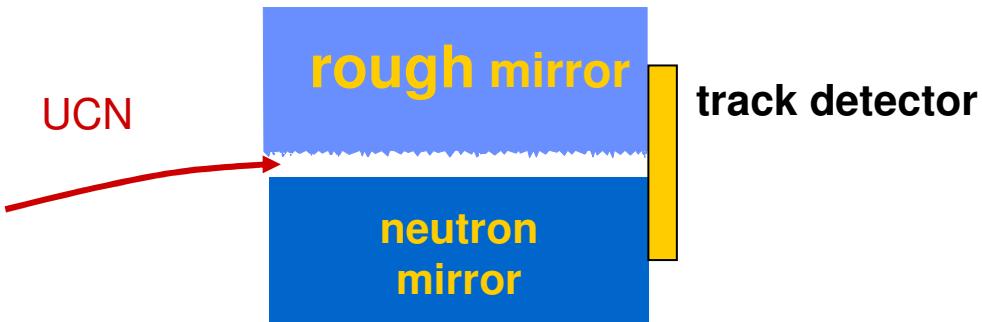
Region I & III:

- State selection via rough upper mirror

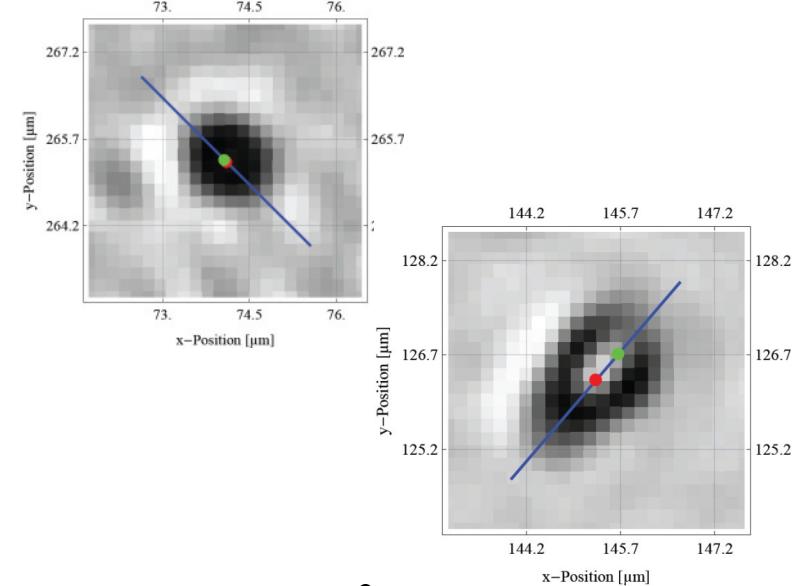
Region II:

- Mechanical Vibrations (qBounce)
- Magnetic gradient fields (GRANIT)

Region I & III: State Selection

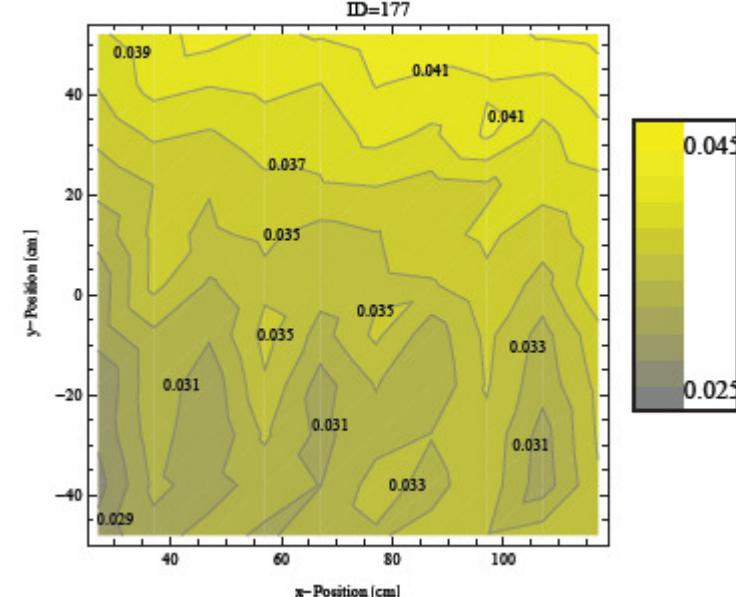
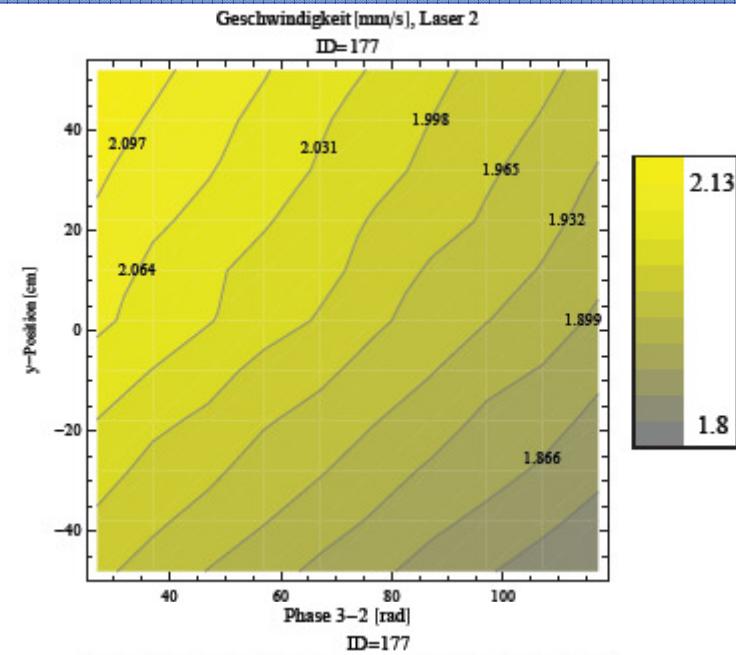
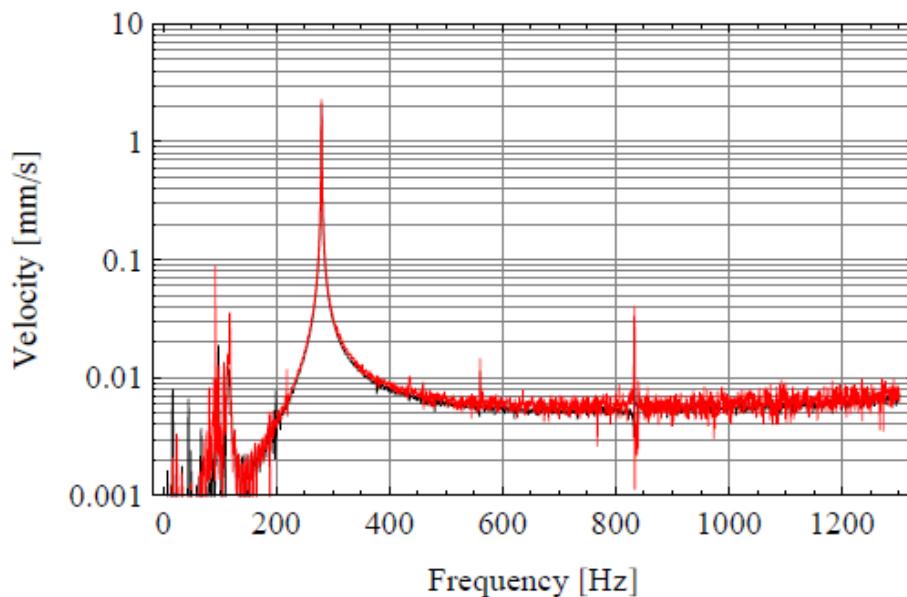
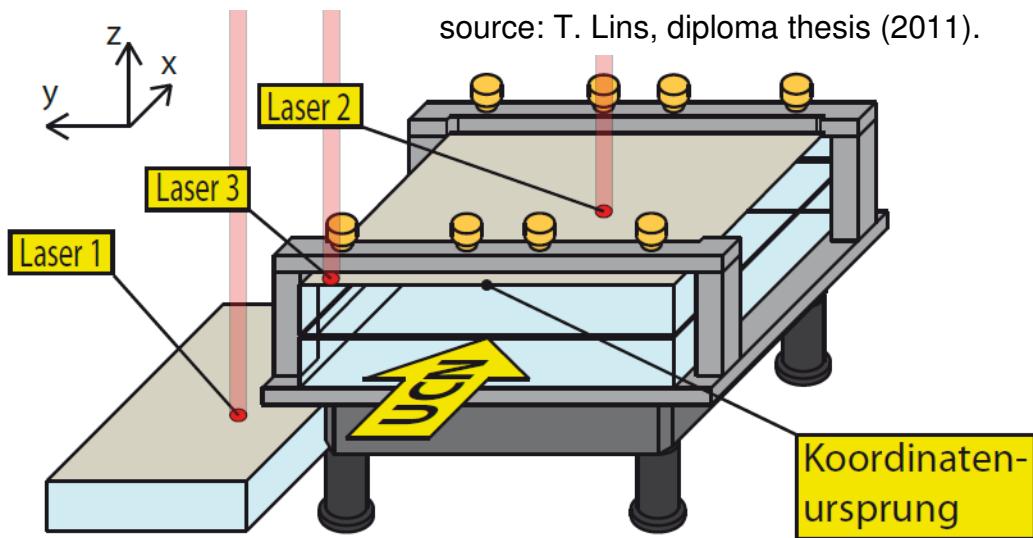


- 4.5 days of beam time
- 3600 events
(background subtracted)

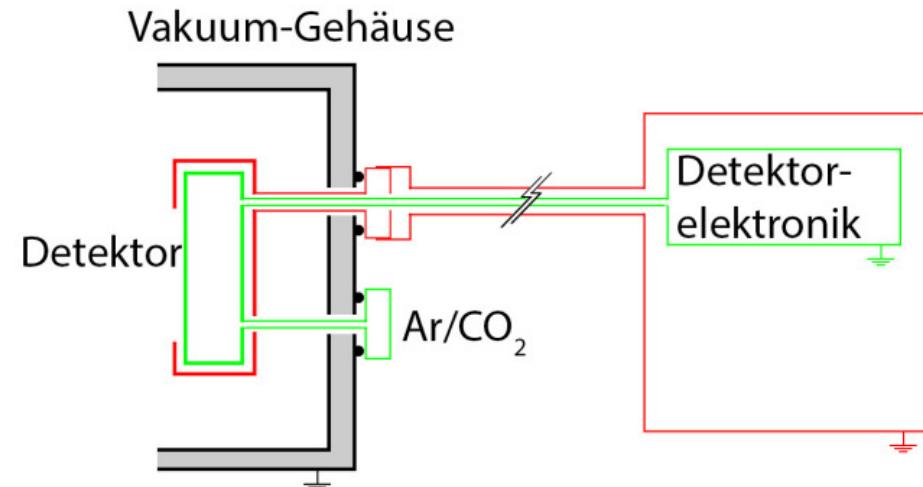
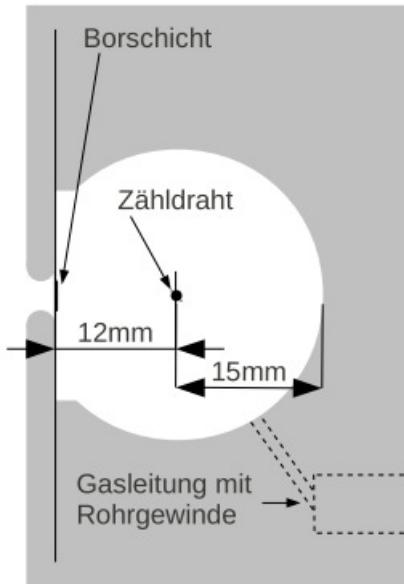


- fit: $N \cdot |\psi|^2 * PSF(\sigma) * f(t)$
- free parameters: $|c_n(t_1)|^2, l, N, z_0$
- result:
 - $|c_1(t_1)|^2 = 0,70$
 - $|c_2(t_1)|^2 = 0,30$
 - $|c_3(t_1)|^2 = 0,00$

Region II: Vibration control



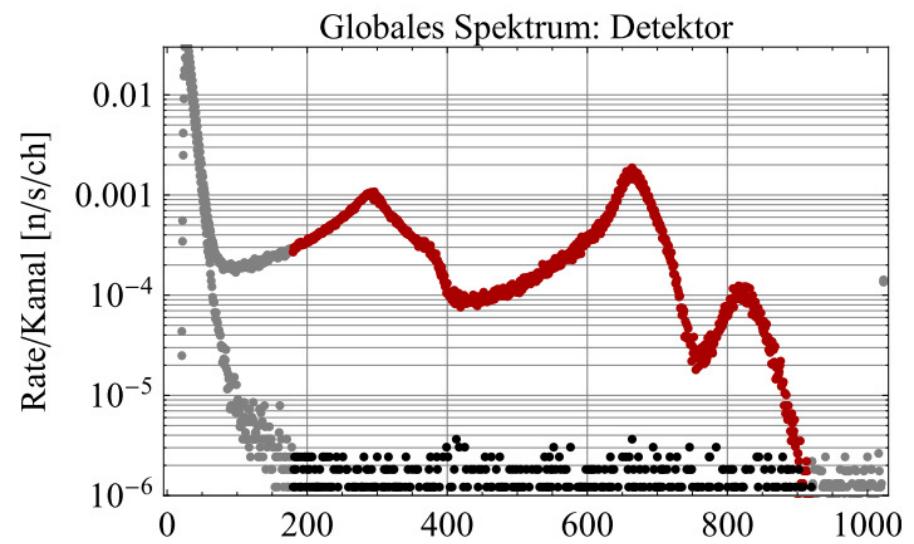
Detector



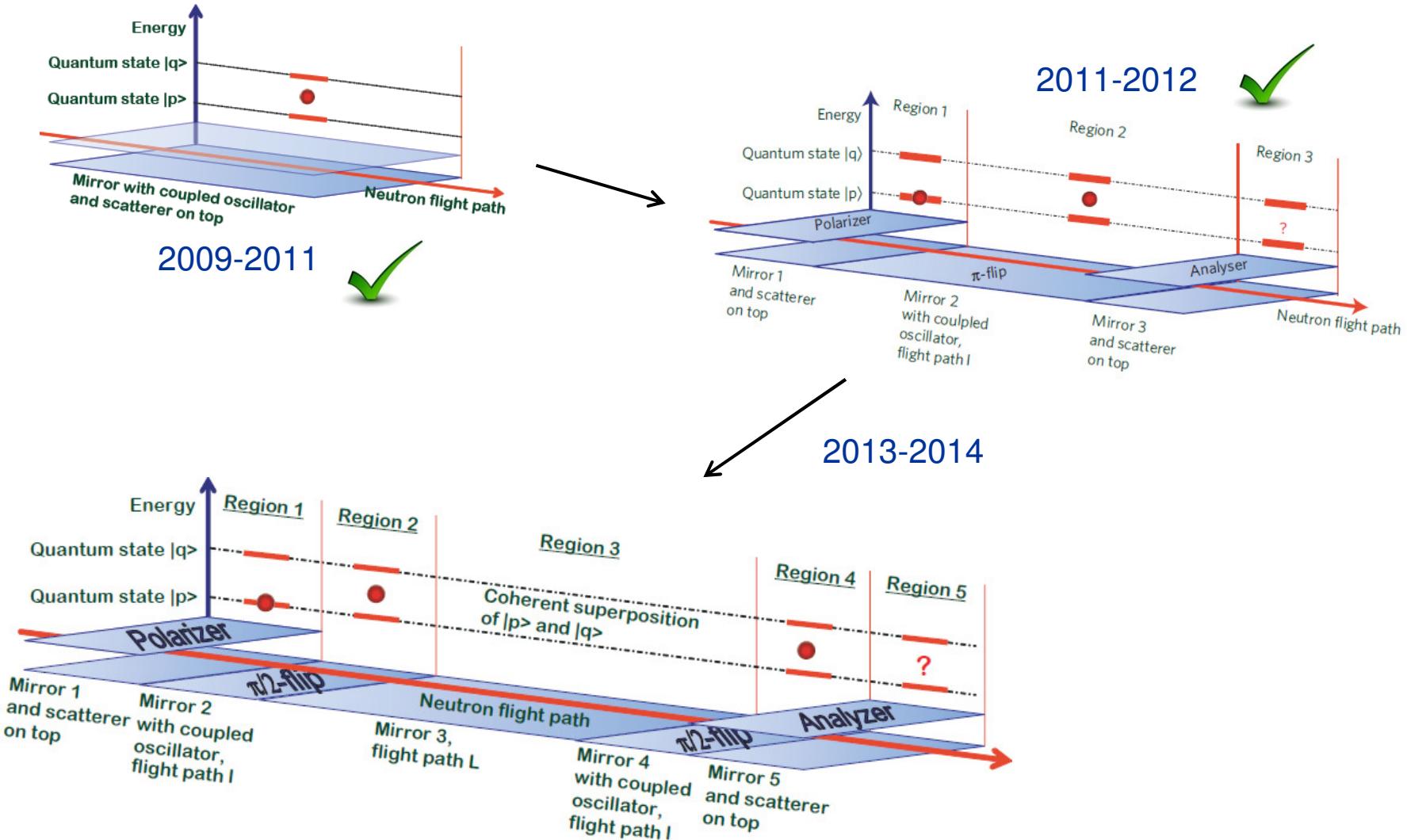
Total detector background
(@PF2 beam position):

$$R = (0,65 \pm 0,02) \cdot 10^{-3} \text{ cts}/\text{s}$$

Efficiency: $\varepsilon = 86,4\%$



Our Method: Gravity Resonance Spectroscopy

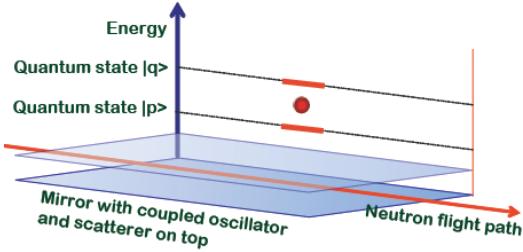


Proposal: H. Abele, T. Jenke, H. Leeb & J. Schmiedmayer, Phys.Rev. D81,065019 (2010)

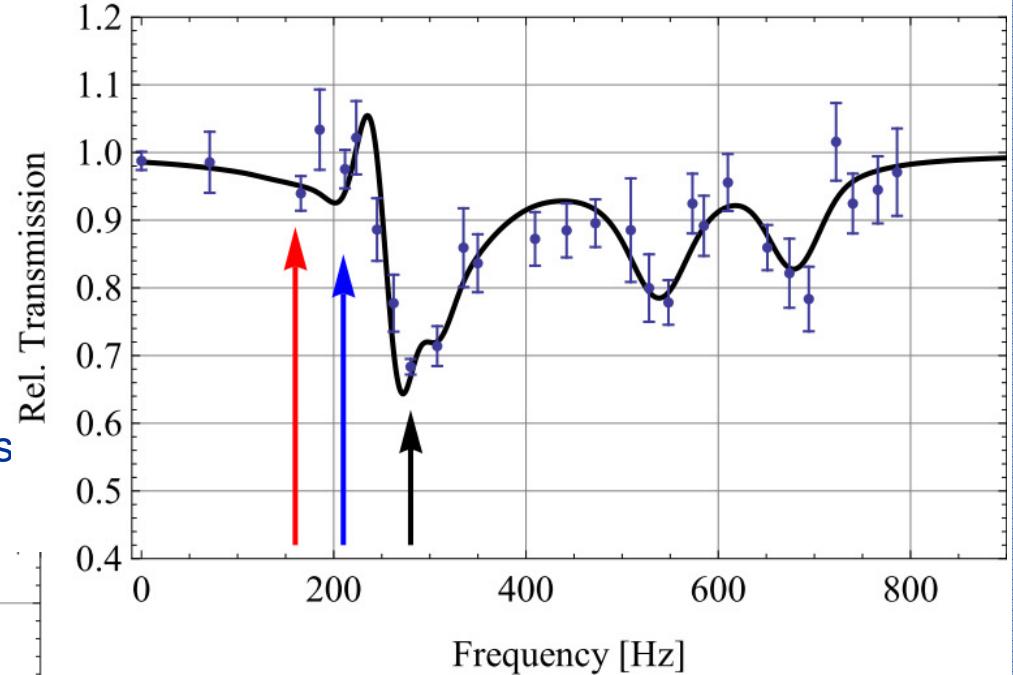
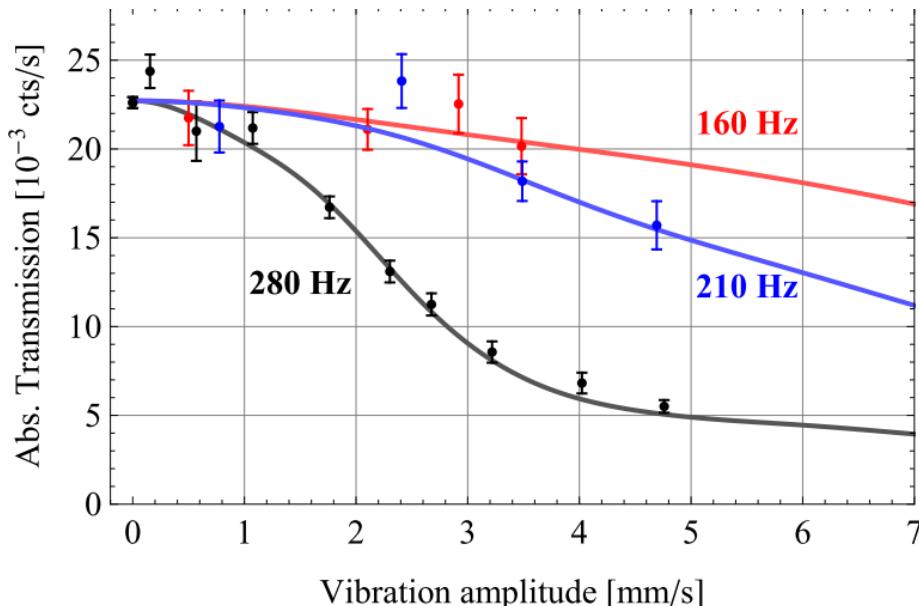
T. Jenke, NPPatLPS 2013, Grenoble, 27/03/2013

Gravity Resonance Spectroscopy

[data 2010]



- transitions:
- $|1\rangle \leftrightarrow |2\rangle, |1\rangle \leftrightarrow |3\rangle, |2\rangle \leftrightarrow |3\rangle$ and $|2\rangle \leftrightarrow |4\rangle$
- 50 days of beam time, 116 measurements



- stat. Significance: 48σ
- stat. accuracy: $\nu_{12} = 258.2 \text{ Hz} \pm 0.8\%$
 $\nu_{23} = 280.4 \text{ Hz} \pm 1.0\%$
 $\nu_{13} = 539.1 \text{ Hz} \pm 0.5\%$
 $\nu_{24} = 679.5 \text{ Hz} \pm 2.2\%$
- contrast: 60%

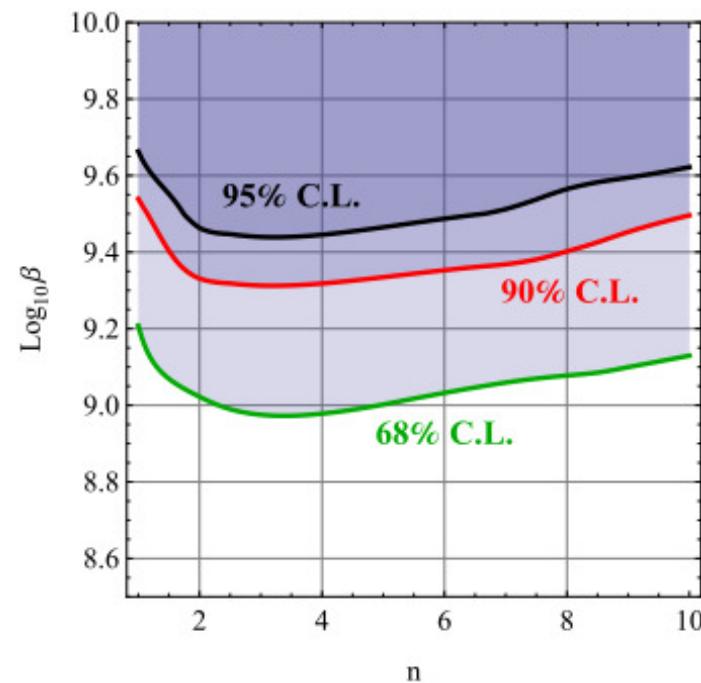
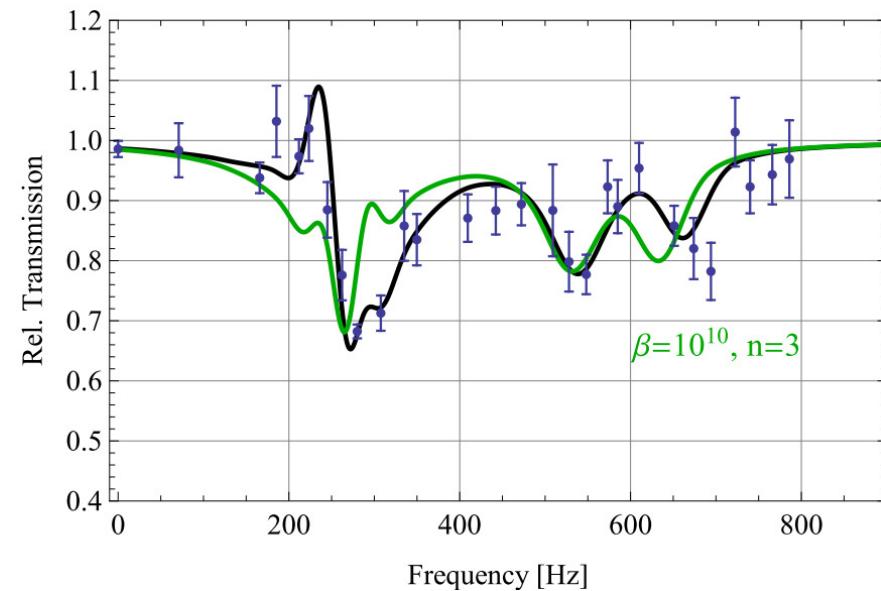
Applications I: Strongly coupled chameleons

[data 2010]



in collaboration with A.N. Ivanov (TU Vienna)

$$V_{\text{Chameleon}} = \beta \frac{m}{M_{Pl}} \Lambda \left(\frac{n+2}{\sqrt{2}} \frac{\Lambda}{d} \left(\frac{d^2}{2} - z^2 \right) \right)^{\frac{2}{n+2}}$$



A.N. Ivanov et.al., arXiv:1207.0419 (2012), accepted PRD (March 2013).
T. Jenke et.al., arXiv:1208.3875 (2012)

Applications II:

[data 2010 + 2011]



Spin-dependant short-ranged interactions

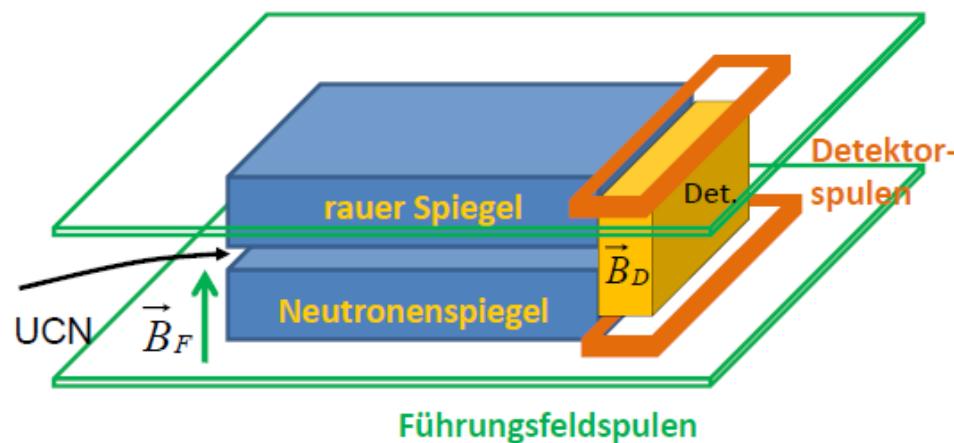
in collaboration with U. Schmidt (PI, Uni Heidelberg) and T. Lauer (FRMII, Garching)

$$V_{\text{axion}} = \frac{g_s g_p \hbar}{8\pi m_n c} \vec{\sigma} \cdot \vec{n} \left(\frac{1}{\lambda r} + \frac{1}{r^2} \right)$$

J.E. Moody, F. Wilczek, Phys. Rev. D30, 131-138 (1984)

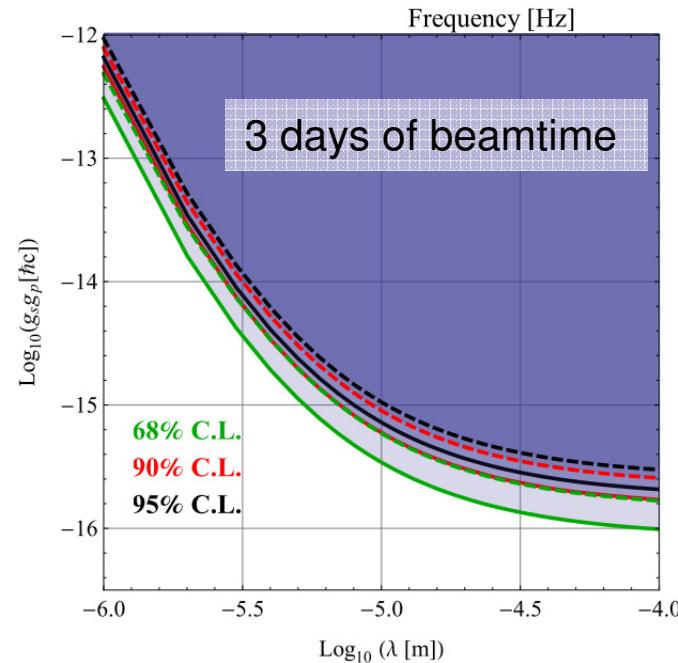
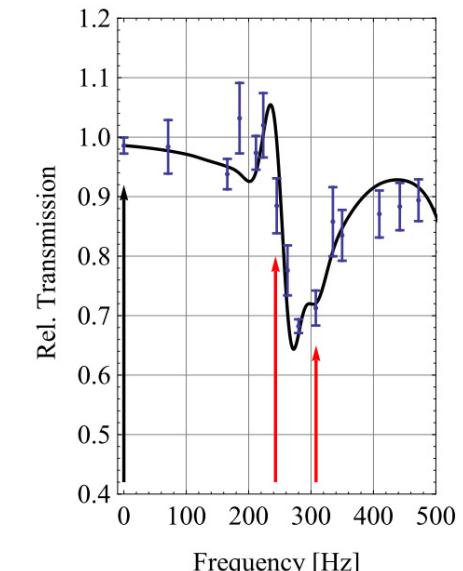
discovery potential [Setup 2010]:

$$g_s g_p / \hbar c \geq \frac{3 \cdot 10^{-16}}{\sqrt{\text{days}}} \quad (\lambda = 10 \mu\text{m}, 68\% \text{ C.L.})$$

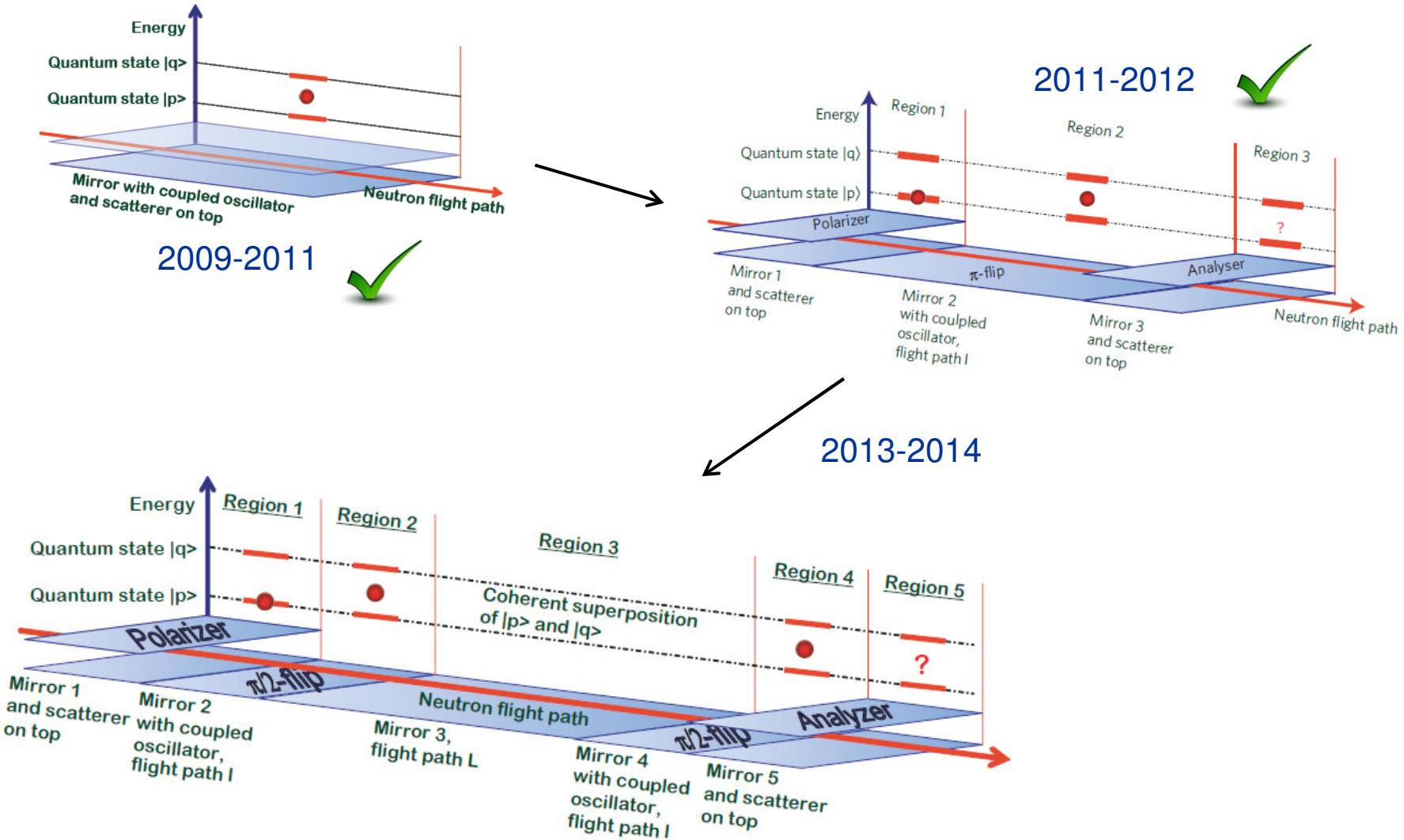


T. Jenke et.al., arXiv:1208.3875 (2012)

T. Jenke, NPPatLPS 2013, Grenoble, 27/03/2013



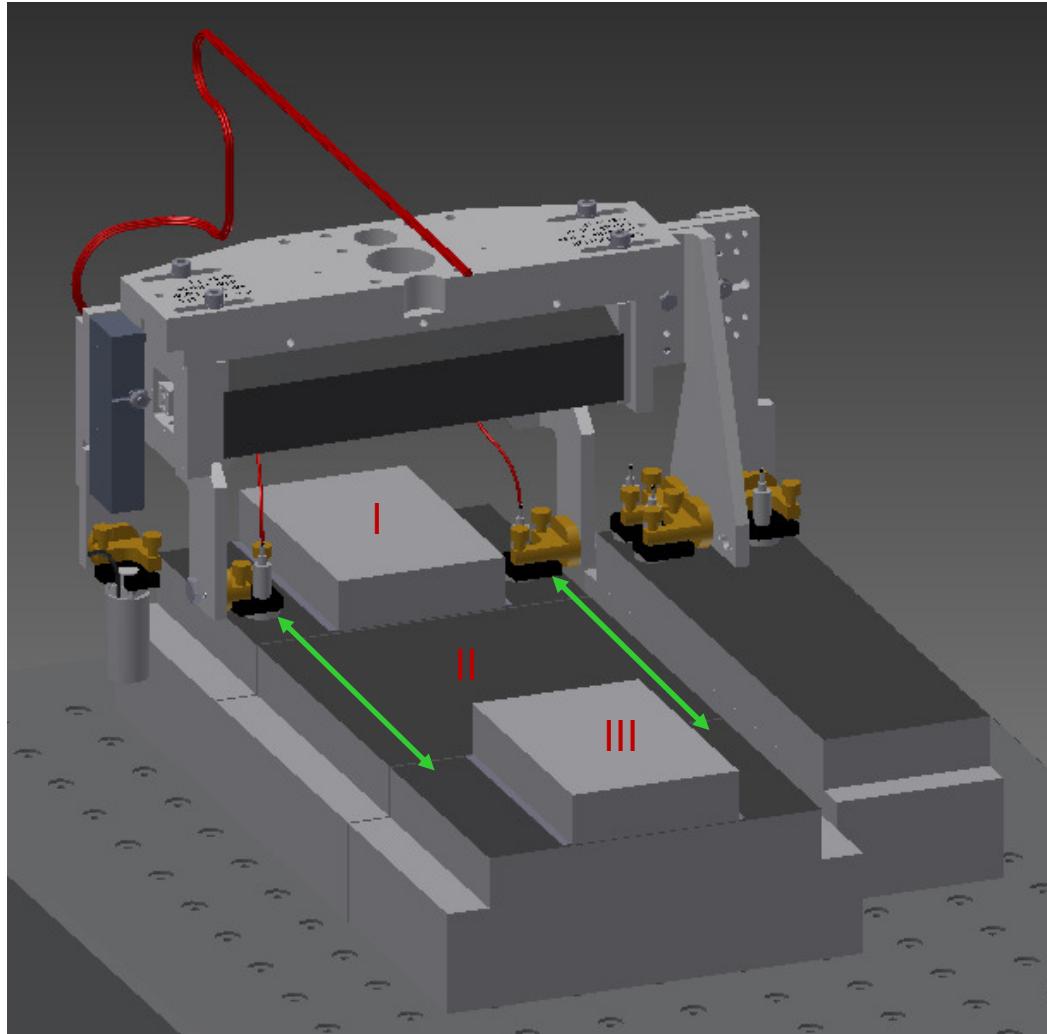
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T. Jenke, NPPatLPS 2013, Grenoble, 27/03/2013

Mind the gap... (since 2012)



Mirror Adjustment Control System

- **Highly stable mount**
- **Adjustment control system:**
 - 1 micro-positioning table (30cm traveling distance)
 - 2 capacitive sensors to measure the
- **Reference system:**
 - 1 reference mirror
 - 3 capacitive sensors

Total positioning error
(measurements 2012):

$$\delta_{sys} \leq 140nm \quad (\text{repeatability})$$

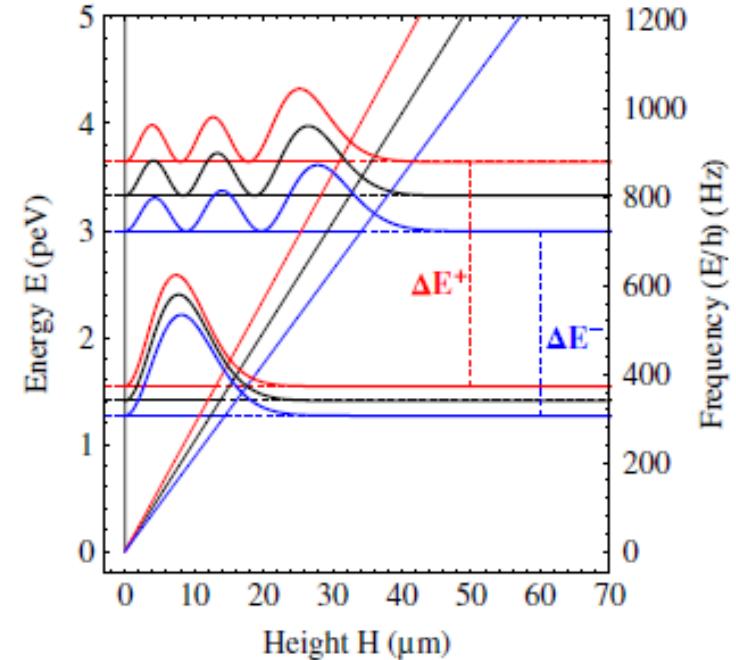
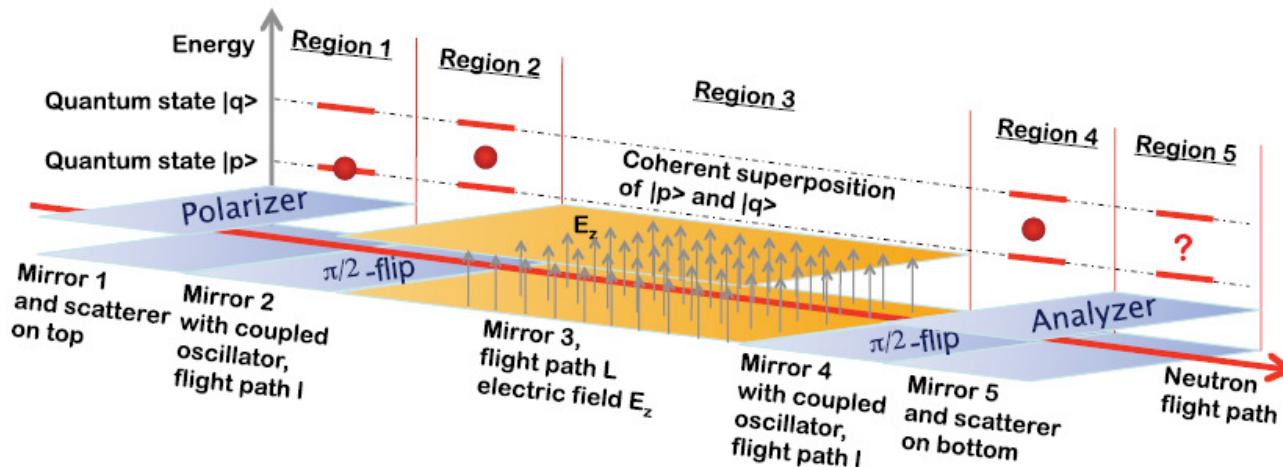
Long-term stability:

$$\delta_{stab} \leq \pm 20nm$$

Outlook: Test of neutron's neutrality

- basic ideas:
 - first Ramsey-like experiment to test neutrons's neutrality
 - null experiment
 - higher electric fields by smaller distances feasible
- frequency shift:

$$\frac{\Delta\nu}{\nu_{pq}} = \sqrt[3]{\left(1 - \frac{qE_z}{mg}\right)^2} - \sqrt[3]{\left(1 + \frac{qE_z}{mg}\right)^2} \approx \frac{4}{3} \frac{qE_z}{mg}$$



discovery potential:

$$\delta q_n(t = 1 \text{ day}) = 3 \cdot 10^{-20} q_e$$

using less than 10.000 neutrons...

Limitation I: UCN source

- flux at PF2/UCN (gold foil activation):

$$\Phi = 7,7 \cdot 10^3 \frac{n}{cm^2 s}$$

- experiment's count rate:

- 2010:

- 15cm long setup

- Vacuum chamber and beam pipe separated
 - Velocity: 5,7..9,5 m/s

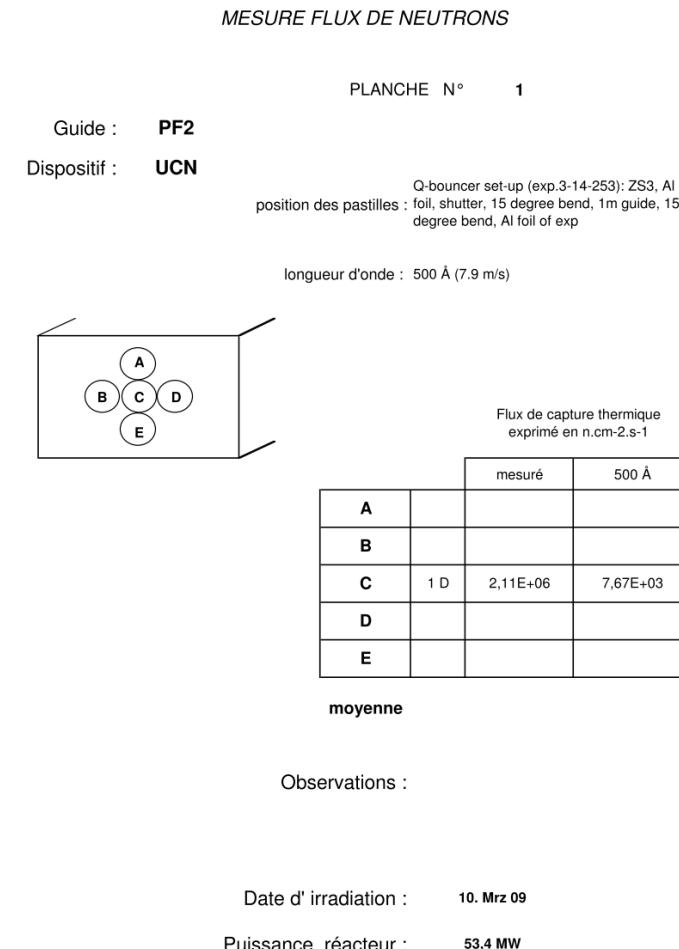
$$T = 22 \cdot 10^{-3} \frac{cts}{s}$$

- 2012:

- 50cm long setup (conventional Rabi)
- Beam pipe connected to chamber
- Velocity: 5,7..9,5 m/s

$$T = 10 \cdot 10^{-3} \frac{cts}{s}$$

24.03.2013



Systematic Influences – State of the Art 2012...

- **remarks:**

- We measure differences.
- Some experiments are Null Experiments. (q, Axions...)
- The numbers given are State-of-the-Art 2012 and may be corrected for or improved...

- general frontiers:

- The moon rotates around the Earth....
- ...and causes tides (ebb and flow), that change g!

- stabilization of the experiment:

- total, absolute inclination error (2012 @ PF2): $\delta_{inc} = \pm 30 \mu rad$
- stabilization of inclination of the neutron mirrors (2012 @ PF2):

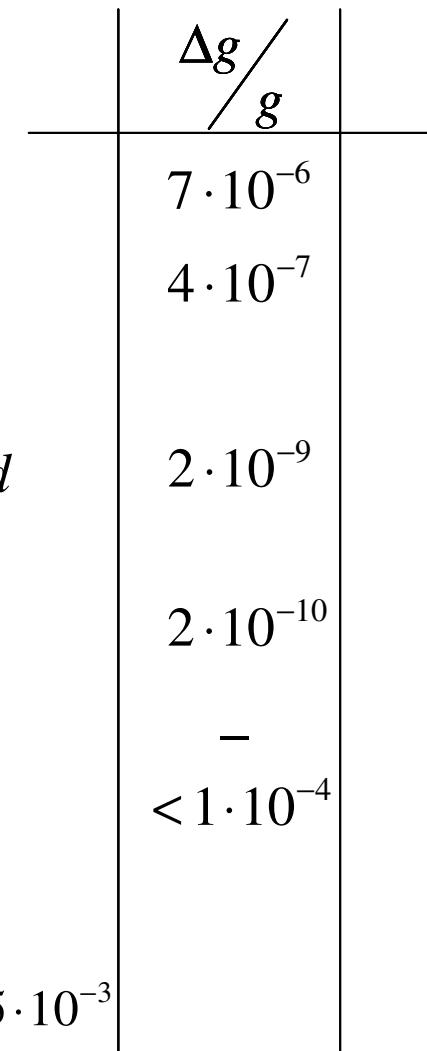
$$\delta_{rms} = 2 \mu rad$$

- experimental environment:

- constant magnetic fields cancel
- external magnetic gradient fields: (2012, 1 layer μ -metal)

$$\Delta B / \Delta z < \pm 1 \mu T / cm$$

- external vibrations (resonant, 2010@PF2) $u_{res}(v_{pq}) / u_{env}(v_{pq}) < 5 \cdot 10^{-3}$



Systematic Frontiers II...

- quality of the neutron mirrors:
 - roughness*: $R_a < 1.7 \text{ nm}$
 - waviness (cm)*: $W_{z,\max} < 10 \text{ nm}$

(* measured externally by SDH, Heidelberg and U. Schmidt, PI Heidelberg)

- magnetic fields due to ferromagnetic enclosures inside the neutron mirrors*

(* in measurement at ATI by P. Kregsamer)

- quality of the external vibrations in region II:

- frequency stability
(externally measured via laser beams, 2012)
- homogeneity of the vibration amplitude
-> linear inhomogeneities cancels

$$\Delta\nu/\nu < 1.1 \cdot 10^{-5}$$

$$< 1.6 \cdot 10^{-5}$$

- correction for incoming flux

(reactor power adjustments, control rods, temp. of cold source):
correction linear at least to 1% (measured)

$$\Delta T/T \approx (0.5 - 3) \cdot 10^{-2}$$

$$< 1 \cdot 10^{-4}$$

- roughness of the absorber + inability to measure the slit size

- influence of up to 10% in Damped-Rabi-Setup (1-region)
- No influence with ordinary Rabi/Ramsey method

—

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Conclusion: Requirements ESS

- **UCN source**

- in-flight experiments:
 - high flux
 - high stability
 - simple and reliable...
- storing experiments (future)
 - high density

- **space & environment:**

- length > 5m, width > 3m
- separation of experimental area from experiment control advantageous
- stable floor
- decoupling of the experimental area from other experiments
 - (electromagnetic, vibrations...)
- cleanroom conditions (partly)

The Team

Atominstut



Hartmut Abele (project leader)
T.J.
Gunther Cronenberg
Hanno Filter
Martin Thalhammer
Thomas Bittner

ILL, Grenoble

Peter Geltenbort



TU Wien

Andrei N. Ivanov
Stefan Rotter
Larisa Chizhova



Phys. Institut, University of Heidelberg

Ulrich Schmidt

FRMII, Garching

Thorsten Lauer